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## Discovering Deep Sky Objects

Over 600 Galaxies, Clusters, Nebulae, Variables and Carbon Stars

## Discovering Deep Sky Objects

Over 600 Galaxies, Clusters, Nebulae, Variables and Carbon Stars (Version 2022-1)

Agnes Clarke
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For Wurzel

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## Using This Book

This book collects some of the most well-known and interesting deep sky objects. The Messier and Caldwell catalogs are included, in addition to many other interesting objects, covering nebulae, star clusters, galaxies and variable and carbon stars.

The 110 objects of the Messier Catalog were logged by Charles Messier in the late 18th century, as an aid to comet hunters by highlighting fuzzy objects that might be mistaken for comets. At some point Messier must have been simply trying to pad out his list, as bright, well-known objects like Messier 45 (the Pleiades) could not reasonably be mistaken for a comet. A drawback of the Messier Catalog is that is biased to the northern hemisphere, but the objects it covers are generally bright and quite accessible to most amateur equipment.

The Caldwell Catalog is much more recent, published by Patrick Moore in 1995. It covers the full sky pole-to-pole, including 109 clusters, galaxies and nebulae. Some of the Caldwell objects, while intriguing, are very difficult to observe. For example, the North America Nebula (Caldwell 20) typically can only be observed under the darkest of skies.

These two catalogs form the the backbone of this book, but a further 169 objects have been included, covering some notable carbon and variable stars and brighter clusters and galaxies. Carbon stars are typically variable in brightness, and are a deep red because of the preponderance of carbon dust in their outer layers. Most are giants at the end of their lives.

While many different types of variable star exist, the majority of those in this book are long period variables that are moderately bright at their maxima, and fade out of the reach of medium-sized telescopes at their dimmest. Where possible I point out nearby stars that can be used to estimate the current magnitude.
I locate targets using a 50 mm finderscope and a red dot finder (RDF). The RDF helps me to point the telescope within a few degrees of the target by simply moving the projected red dot in the sight to the right place in the sky, and then the finderscope takes over as it shows about 5-7 degrees of sky with
about 8-10 times magnification (it varies because my finderscope accepts different eyepieces).
This book has two types of diagrams to help with these two phases. Firstly, large diagrams show brighter stars visible from an urban location with the positions of the deep sky objects indicated by a small color-coded circle - blue targets are easiest, followed by orange and then red targets. This diagram helps with the rough pointing of the telescope via the red dot finder. Stars down to magnitude 5 are included in these overview charts, which is somewhat fainter than I can normally see from a light polluted location, unless it is a good clear night and it's past midnight!

To get the "right place in the sky" for the RDF, I rely on geometrical patterns and projections, and I have include some of these in the finding instructions for many of the objects. However, for some objects, there are no good signposts and one must simply squint at increasingly fainter nearby stars.
The second type of chart shows the faint stars visible in a finderscope (a $5^{\circ}$ finder circle is shown on each chart), so that I can finally locate the object by matching the pattern to the book. I normally have the finderscope and telescope precisely aligned, so that I can keep a medium magnification eyepiece in the telescope and not have to keep switching between lower and higher magnification in the telescope to center the object in the telescope view. The finder chart covers $10^{\circ}$ or the sky, a width spanning 20 full Moons, so even quite rough pointing with the RDF should put you somewhere on the chart.

The finder charts also list up to five other nearby objects. The number is capped to avoid cluttering the view of the star patterns, which is of course the main purpose of the charts.

## Astronomical Abbreviations

Constellations are often referred to in this work by three-letter abbreviations. When the constellation name is used in a star name, the constellation's genitive form is used. For example, Aquarius is abbreviated as Aqr, and one of its stars would be 5 Aquarii.

The following list gives the abbreviation and the full and genitive names.

```
And - Andromeda (Andromedae)
Ant - Antlia (Antliae)
Aps - Apus (Apodis)
Aqr - Aquarius (Aquarii)
Aql - Aquila (Aquilae)
Ara - Ara (Arae)
Ari - Aries (Arietis)
Aur - Auriga (Aurigae)
Boo - Bootes (Bootis)
Cae - Caelum (Caeli)
Cam - Camelopardalis (Camelopardalis)
Cnc - Cancer (Cancri)
CVn - Canes Venatici (Canum Venaticorum)
CMa - Canis Major (Canis Majoris)
CMi - Canis Minor (Canis Minoris)
Cap - Capricornus (Capricorni)
Car - Carina (Carinae)
Cas - Cassiopeia (Cassiopeiae)
Cen - Centaurus (Centauri)
Cep - Cepheus (Cephei)
Cet - Cetus (Ceti)
Cha - Chamaeleon (Chamaeleontis)
Cir - Circinus (Circini)
Col - Columba (Columbae)
Com - Coma Berenices (Comae Berenices)
```

CrA - Corona Australis (Coronae Australis)
CrB - Corona Borealis (Coronae Borealis)
Crv - Corvus (Corvi)
Crt - Crater (Crateris)
Cru - Crux (Crucis)
Cyg - Cygnus (Cygni)
Del - Delphinus (Delphini)
Dor - Dorado (Doradus)
Dra - Draco (Draconis)
Equ - Equuleus (Equulei)
Eri - Eridanus (Eridani)
For - Fornax (Fornacis)
Gem - Gemini (Geminorum)
Gru - Grus (Gruis)
Her - Hercules (Herculis)
Hor - Horologium (Horologii)
Hya - Hydra (Hydrae)
Hyi - Hydrus (Hydri)
Ind - Indus (Indi)
Lac - Lacerta (Lacertae)
Leo - Leo (Leonis)
LMi - Leo Minor (Leonis Minoris)
Lep - Lepus (Leporis)
Lib - Libra (Librae)
Lup - Lupus (Lupi)
Lyn - Lynx (Lyncis)
Lyr - Lyra (Lyrae)
Men - Mensa Mensae)
Mic - Microscopium (Microscopii)
Mon - Monoceros (Monocerotis)
Mus - Musca (Muscae)

| Nor - Norma (Normae) | Greek letters are used in Bayer designations of stars, |
| :---: | :---: |
| Oct - Octans (Octantis) | such as Alpha Canis Majoris (Sirius). Greek letters |
| Oph - Ophiuchus (Ophiuchi) | follows: |
| Ori - Orion (Orionis) | $\alpha$ - alf - alpha |
| Pav - Pavo (Pavonis) | $\beta$ - bet - beta |
| Peg - Pegasus (Pegasi) | $\gamma$ - gam - gamma |
| Per - Perseus (Persei) | $\delta$ - del - delta |
| Phe - Phoenix (Phoenicis) | $\varepsilon$ - eps - epsilon |
| Pic - Pictor (Pictoris) | $\zeta$ - zet - zeta |
| Psc - Pisces (Piscium) | $\eta$ - eta - eta |
| PsA - Piscis Austrinus (Piscis Austrini) | $\theta$ - tet - theta |
| Pup - Puppis (Puppis) | l - iot - iota |
| Pyx - Pyxis (Pyxidis) | $\kappa$ - kap - kappa |
| Ret - Reticulum (Reticuli) | $\lambda$ - lam -lamda, lambda |
| Sge - Sagitta (Sagittae) | $\mu$ - mu. - mu |
| Sgr - Sagittarius (Sagittarii) | $v$ - nu. - nu |
| Sco - Scorpius (Scorpii) | $\xi$ - ksi-xi |
| Scl - Sculptor (Sculptoris) | o-omi-omicron |
| Sct - Scutum (Scuti) | $\pi$ - pi. - pi |
| Ser - Serpens (Serpentis) | $\rho-$ rho - rho |
| Sex - Sextans (Sextantis) | $\sigma$ - sig - sigma |
| Tau - Taurus (Tauri) | $\tau$ - tau-tau |
| Tel - Telescopium (Telescopii) | $v$ - ups - upsilon |
| Tri - Triangulum (Trianguli) | $\varphi$ - phi - phi |
| TrA - Triangulum Australe (Trianguli Australis) | $\chi$ - khi - chi |
| Tuc - Tucana (Tucanae) | $\psi$ - psi - psi |
| UMa - Ursa Major (Ursae Majoris) | $\omega$ - ome - omega |
| UMi - Ursa Minor (Ursae Minoris) |  |
| Vel - Vela (Velorum) |  |
| Vir - Virgo (Virginis) |  |
| Vol - Volans (Volantis) |  |
| Vul - Vulpecula (Vulpeculae) |  |

## Deep Sky Objects by Constellation

## Andromeda

M110 (NGC 205) is a magnitude 8.5 elliptical galaxy. Angular size is $17 \times 10^{\prime}$. (page 65)
Observed:
M32 (NGC 221) is a magnitude 8.1 elliptical galaxy. Angular size is $8 \times 66^{\prime}$. (page 66)
Observed:
C28 (NGC 752) is a magnitude 5.7 open cluster.
Angular size is 50 '. (page 87)
Observed:
NGC 7686 is a magnitude 5.6 open cluster. Angular size is $15^{\prime}$. (page 64)
Observed:
C22 (Blue Snowball, NGC 7662) is a magnitude 9.2 planetary nebula. Angular size is 0.6 '. (page 61) Observed:

C23 (NGC 891) is a magnitude 9.9 barred spiral galaxy. Angular size is $14 \times 3{ }^{\prime}$. (page 87)

Observed:
M31 (Andromeda Galaxy, NGC 224) is a magnitude 3.4 spiral galaxy. Angular size is $178 \times 63^{\prime}$. (page 62)

Observed:
R And (HD 1967) is a magnitude 5.8 variable star.
Magnitude ranges from 14.9 to 5.8 ( $\Delta$ mag. 9.1)
with a period of 409 d . (page 62)
Observed:

## Antlia

U Ant (HD 91793) is a magnitude 5.7 carbon star. Magnitude ranges from 8.6 to 5.7 ( $\Delta$ mag. 2.9) with a period of 170d. (page 199)
Observed:

## Apus

C107 (NGC 6101) is a magnitude 9.3 globular cluster.
Angular size is 11 '. (page 308)
Observed:

## Aquarius

M73 (NGC 6994) is a magnitude 9.0 asterism. Angular size is $3^{\prime}$. (page 296)
Observed:
M2 (NGC 7089) is a magnitude 6.5 globular cluster. Angular size is $12.9^{\prime}$. (page 293)
Observed:
M72 (NGC 6981) is a magnitude 9.3 globular cluster. Angular size is 6 '. (page 293)
Observed:
C55 (Saturn Nebula, NGC 7009) is a magnitude 8.3 planetary nebula. Angular size is $0.7^{\prime}$. (page 295) Observed:

C63 (Helix Nebula, NGC 7293) is a magnitude 6.5 planetary nebula. Angular size is $13^{\prime}$. (page 297) Observed:

R Aqr (HD 222800) is a magnitude 5.8 variable star. Magnitude ranges from 12.4 to 5.8 ( $\Delta$ mag. 6.6) with a period of 387 d . (page 73)

Observed:

## Aquila

NGC 6709 is a magnitude 6.7 open cluster. Angular size is $13^{\prime}$. (page 241)
Observed:
R Aql (HD 177940) is a magnitude 5.5 variable star. Magnitude ranges from 12.0 to 5.5 ( $\Delta$ mag. 6.5) with a period of 284d. (page 242)

Observed:
Ara
C81 (NGC 6352) is a magnitude 8.1 globular cluster. Angular size is 7'. (page 269) Observed:

C86 (NGC 6397) is a magnitude 5.6 globular cluster.
Angular size is $26^{\prime}$. (page 270)

## Observed:

C82 (NGC 6193) is a magnitude 5.2 open cluster.
Angular size is $15^{\prime}$. (page 265)
Observed:
NGC 6250 is a magnitude 5.9 open cluster. Angular size is $8^{\prime}$. (page 264)
Observed:

## Aries

V Ari (HD 13826) is a magnitude 8.0 carbon star. Magnitude ranges from 8.6 to 8.0 ( $\Delta$ mag. 0.6 ) with a period of 75d. (page 92)
Observed:
R Ari (HD 13913) is a magnitude 7.4 variable star. Magnitude ranges from 13.7 to 7.4 ( $\Delta$ mag. 6.3) with a period of 187 d . (page 90)
Observed:
U Ari (HD 19737) is a magnitude 7.2 variable star. Magnitude ranges from 15.2 to 7.2 ( $\Delta$ mag. 8.0) with a period of 371d. (page 92)
Observed:

## Auriga

C31 (Flaming Star Nebula, IC 405) is a magnitude 6.0 bright nebula. Angular size is $30 \times 19^{\prime}$. (page 109) Observed:
UU Aur (HD 46687) is a magnitude 5.1 carbon star. Magnitude ranges from 7.0 to 5.1 ( $\Delta$ mag. 1.9) with a period of 235d. (page 107)

Observed:
M36 (NGC 1960) is a magnitude 6.3 open cluster.
Angular size is 12 '. (page 109)
Observed:
M37 (NGC 2099) is a magnitude 6.2 open cluster.
Angular size is $24^{\prime}$. (page 110)
Observed:
M38 (NGC 1912) is a magnitude 7.4 open cluster.
Angular size is $21^{\prime}$. (page 108)
Observed:
NGC 2281 is a magnitude 5.4 open cluster. Angular size is $15^{\prime}$. (page 107)
Observed:

AE Aur (HD 4078) is a magnitude 5.78 variable star. Magnitude ranges from 6.08 to 5.78 ( $\Delta$ mag. 0.3). (page 108)
Observed:
R Aur (HD 34019) is a magnitude 6.7 variable star. Magnitude ranges from 13.9 to 6.7 ( $\Delta$ mag. 7.2 ) with a period of 458 d . (page 106)
Observed:

## Boötes

C45 (NGC 5248) is a magnitude 10.2 spiral galaxy. Angular size is $6 \times 44^{\prime}$. (page 217)
Observed:
R Boo (HD 128609) is a magnitude 6.2 variable star. Magnitude ranges from 13.1 to 6.2 ( $\Delta$ mag. 6.9) with a period of 223d. (page 215)

## Observed:

## Camelopardalis

NGC 1502 is a magnitude 5.7 open cluster. Angular size is $8^{\prime}$. (page 81 )
Observed:
C5 (Maffei 1 Group, IC342) is a magnitude 9.2 spiral galaxy. Angular size is $21 \times 21^{\prime}$. (page 54 )
Observed:
C7 (NGC 2403) is a magnitude 8.9 spiral galaxy.
Angular size is $22 \times 12^{\prime}$. (page 136)
Observed:

## Cancer

M44 (Beehive Cluster, NGC 2632) is a magnitude 3.7 open cluster. Angular size is $95^{\prime}$. (page 140) Observed:

M67 (NGC 2682) is a magnitude 6.1 open cluster.
Angular size is $30^{\prime}$. (page 141)
Observed:
C48 (NGC 2775) is a magnitude 10.3 spiral galaxy.
Angular size is $4.5 \times 3^{\prime}$. (page 143)

## Observed:

R Cnc (HD 69243) is a magnitude 6.07 variable star. Magnitude ranges from 11.8 to 6.07 ( $\Delta$ mag. 5.7) with a period of 362d. (page 141)
Observed:
X Cnc (HD 76221) is a magnitude 5.6 variable star. Magnitude ranges from 7.5 to 5.6 ( $\Delta \mathrm{mag}$. 1.9) with a period of 195d. (page 140)

Observed:

## Canes Venatici

Y Cvn (La Superba, HD 110914) is a magnitude 5.0 carbon star. Magnitude ranges from 6.4 to 5.0 ( $\Delta$ mag. 1.4) with a period of 158 d . (page 169)
Observed:
M3 (NGC 5272) is a magnitude 6.2 globular cluster.
Angular size is $16.2^{\prime}$. (page 214)
Observed:
C21 (NCG 4449, NGC 4449) is a magnitude 9.4 irregular galaxy. Angular size is $6 \times 4^{\prime}$. (page 173)
Observed:
C29 (NGC 5005) is a magnitude 9.8 barred spiral galaxy. Angular size is $5 \times 2^{\prime}$. (page 170)
Observed:
C26 (NGC 4244) is a magnitude 10.6 spiral galaxy. Angular size is $17 \times 2^{\prime}$. (page 174)
Observed:
C32 (Whale Galaxy, NGC 4631) is a magnitude 9.3 spiral galaxy. Angular size is $15 \times 3$ '. (page 170)
Observed:
M106 (NGC 4258) is a magnitude 8.4 spiral galaxy. Angular size is $19 \mathrm{x} 8^{\prime}$. (page 172)
Observed:
M51 (Whirlpool Galaxy, NGC 5194) is a magnitude 8.4 spiral galaxy. Angular size is $11 \times 7$ '. (page 168)

Observed:
M63 (Sunflower Galaxy, NGC 5055) is a magnitude 8.6 spiral galaxy. Angular size is $10 \times 6$ '. (page 169)

Observed:

M94 (NGC 4736) is a magnitude 8.2 spiral galaxy. Angular size is 7x3'. (page 173)
Observed:

## Canis Major

C58 (Caroline's Cluster, NGC 2360) is a magnitude
7.2 open cluster. Angular size is $13^{\prime}$. (page 129)

Observed:
C64 (Pirate's Jewels Cluster, NGC 2362) is a magnitude 4.1 open cluster. Angular size is $8^{\prime}$. (page 130)
Observed:
M41 (NGC 2287) is a magnitude 4.6 open cluster. Angular size is $38^{\prime}$. (page 129)
Observed:
NGC 2354 is a magnitude 6.5 open cluster. Angular size is 20 '. (page 126)
Observed:
R CMa (HD 57167) is a magnitude 5.7 variable star. Magnitude ranges from 6.34 to 5.7 ( $\Delta$ mag. 0.6) with a period of 1.13594 d . (page 125)
Observed:
VY CMa (HD 58061) is a magnitude 6.5 variable star. Magnitude ranges from 9.6 to 6.5 ( $\Delta$ mag. 3.1). (page 130)
Observed:

## Canis Minor

S CMi (HD 59950) is a magnitude 6.6 variable star. Magnitude ranges from 13.2 to 6.6 ( $\Delta$ mag. 6.6) with a period of 333d. (page 142)

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Observed:
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## Capricornus

RT Cap (HD 192737) is a magnitude 6.5 carbon star. Magnitude ranges from 8.1 to 6.5 ( $\Delta$ mag. 1.6) with a period of 395 d . (page 294)
Observed:
M30 (NGC 7099) is a magnitude 7.2 globular cluster. Angular size is 11 '. (page 297)

Observed:

## Carina

C92 (Eta Carinae Nebula, NGC 3372) is a magnitude 6.2 bright nebula. Angular size is $120 \times 120$ '. (page 200)

Observed:
NGC 2808 (NGC) is a magnitude 6.2 globular cluster. Angular size is $14^{\prime}$. (page 159)
Observed:
C102 (Theta Car Cluster, IC 2602) is a magnitude 1.9 open cluster. Angular size is 50 '. (page 202)
Observed:
C91 (Wishing Well Cluster, NGC 3532) is a magnitude
3.0 open cluster. Angular size is $55^{\prime}$. (page 200)

Observed:
C96 (Southern Beehive, NGC 2516) is a magnitude 3.8 open cluster. Angular size is $30^{\prime}$. (page 158)

Observed:
IC 2581 is a magnitude 4.3 open cluster. Angular size is $8^{\prime}$. (page 157)
Observed:
NGC 3114 is a magnitude 4.2 open cluster. Angular size is $35^{\prime}$. (page 162)
Observed:
C90 (NGC 2867) is a magnitude 9.7 planetary nebula. Angular size is $0.2^{\prime}$. (page 161)

Observed:
R Car (HD 82901) is a magnitude 3.9 variable star. Magnitude ranges from 10.5 to 3.9 ( $\Delta$ mag. 6.6) with a period of 309d. (page 163)
Observed:
S Car (HD 88366) is a magnitude 4.5 variable star. Magnitude ranges from 9.9 to 4.5 ( $\Delta$ mag. 5.4) with a period of 149d. (page 162)
Observed:
1 Car (HD 84810) is a magnitude 3.28 variable star. Magnitude ranges from 4.18 to 3.28 ( $\Delta$ mag. 0.9) with a period of 35.53584 d . (page 158)
Observed:

## Cassiopeia

C11 (Bubble Nebula, NGC 7635) is a magnitude 7.0 bright nebula. Angular size is $15 \times 8^{1}$. (page 63)

## Observed:

C17 (NGC 147) is a magnitude 9.3 elliptical galaxy. Angular size is $13 \times 8^{\prime}$. (page 61) Observed:

C18 (NGC 185) is a magnitude 9.2 elliptical galaxy. Angular size is $12 \mathrm{x} 9^{\prime}$. (page 65 )

## Observed:

C10 (NGC 663) is a magnitude 7.1 open cluster. Angular size is 16 '. (page 81)
Observed:
C13 (Owl Cluster, NGC 457) is a magnitude 6.4 open cluster. Angular size is $13{ }^{\prime}$. (page 64)
Observed:
C8 (NGC 559) is a magnitude 9.5 open cluster. Angular size is $7^{\prime}$. (page 59)

Observed:
M103 (NGC 581) is a magnitude 7.4 open cluster. Angular size is 6 '. (page 85 )

Observed:
M52 (NGC 7654) is a magnitude 7.3 open cluster. Angular size is $13^{\prime}$. (page 59) Observed:

NGC 129 is a magnitude 6.5 open cluster. Angular size is $21^{\prime}$. (page 60 )

Observed:
NGC 654 is a magnitude 6.5 open cluster. Angular size is $5^{\prime}$. (page 85 )

Observed:
R Cas (HD 224490) is a magnitude 4.7 variable star. Magnitude ranges from 13.5 to 4.7 ( $\Delta$ mag. 8.8) with a period of 430d. (page 60)

Observed:

## Centaurus

C80 (Omega Centauri, NGC 5139) is a magnitude 3.6 globular cluster. Angular size is 36 '. (page 199)
Observed:
C84 (NGC 5286) is a magnitude 7.6 globular cluster. Angular size is $9^{\prime}$. (page 230)
Observed:
C100 (Lambda Centauri Nebula, IC 2944) is a magnitude 4.5 open cluster. Angular size is $15^{\prime}$. (page 204)
Observed:
C97 (Pearl Cluster, NGC 3766) is a magnitude 5.3 open cluster. Angular size is 12 . (page 204)
Observed:
NGC 5281 is a magnitude 5.9 open cluster. Angular size is $5^{\prime}$. (page 232)
Observed:
NGC 5316 is a magnitude 6.0 open cluster. Angular size is $14^{\prime}$. (page 228)
Observed:
NGC 5460 is a magnitude 5.6 open cluster. Angular size is $25^{\prime}$. (page 229)
Observed:
NGC 5617 is a magnitude 6.3 open cluster. Angular size is $10^{\prime}$. (page 228)
Observed:
NGC 5662 is a magnitude 5.5 open cluster. Angular size is $12^{\prime}$. (page 231)

Observed:
C77 (Centaurus A, NGC 5128) is a magnitude 7.0 peculiar galaxy. Angular size is $18 \times 14^{\prime}$. (page 202)

Observed:
C83 (NGC 4945) is a magnitude 9.5 spiral galaxy. Angular size is 20x4'. (page 203)
Observed:

R Cen (HD 124601) is a magnitude 5.3 variable star. Magnitude ranges from 11.8 to 5.3 ( $\Delta$ mag. 6.5) with a period of 546d. (page 231)
Observed:
T Cen (HD 119090) is a magnitude 5.5 variable star. Magnitude ranges from 9.0 to 5.5 ( $\Delta$ mag. 3.5) with a period of 90.44 d . (page 229)
Observed:

## Cepheus

C4 (Iris Nebula, NGC 7023) is a magnitude 6.8 bright nebula. Angular size is $18 \times 18^{\prime}$. (page 51)
Observed:
C9 (Cave Nebula, Sh2-155) is a magnitude 7.7 bright nebula. Angular size is $50 \times 30^{\prime}$. (page 63)
Observed:
mu Cep (Garnet Star, NGC 206936) is a magnitude 3.7 carbon star. Magnitude ranges from 5.0 to 3.7 ( $\Delta$ mag. 1.3). (page 273)
Observed:
C1 (Polarissima Cluster, NGC 188) is a magnitude 8.1 open cluster. Angular size is $15^{\prime}$. (page 53)
Observed:
NGC 7160 is a magnitude 6.1 open cluster. Angular size is $7^{\prime}$. (page 273)
Observed:
C2 (Bow Tie Nebula, NGC 40) is a magnitude 11.6 planetary nebula. Angular size is 0.6 '. (page 52)
Observed:
C12 (NGC 6946) is a magnitude 9.7 spiral galaxy. Angular size is $12 \times 10^{\prime}$. (page 277)

## Observed:

AR Cep (HD 217158) is a magnitude 7.0 variable star. Magnitude ranges from 7.9 to 7.0 ( $\Delta$ mag. 0.9 ). (page 54)
Observed:
T Cep (HD 202012) is a magnitude 5.2 variable star. Magnitude ranges from 11.3 to 5.2 ( $\Delta$ mag. 6.1) with a period of 388 d . (page 55)

## Observed:

U Cep (HD 5679) is a magnitude 6.75 variable star. Magnitude ranges from 9.24 to 6.75 ( $\Delta$ mag. 2.5) with a period of 2.49305 d . (page 51)
Observed:

## Cetus

C51 (IC 1613) is a magnitude 9.0 irregular galaxy. Angular size is $12 \times 11^{\prime}$. (page 70)

Observed:
C56 (Skull Nebula, NGC 246) is a magnitude 8.0 planetary nebula. Angular size is $3.8^{\prime}$. (page 72) Observed:

C62 (Needle's Eye Galaxy, NGC 247) is a magnitude 8.9 spiral galaxy. Angular size is $20 \times 7{ }^{\prime}$. (page 74)

Observed:
M77 (NGC 1068) is a magnitude 8.9 spiral galaxy. Angular size is 7x6'. (page 94)
Observed:
T Cet (HD 1760) is a magnitude 5.0 variable star. Magnitude ranges from 6.9 to $5.0(\Delta \mathrm{mag} .1 .9)$ with a period of 159 d . (page 73)

Observed:
U Cet (HD 15971) is a magnitude 6.8 variable star. Magnitude ranges from 13.4 to 6.8 ( $\Delta$ mag. 6.6) with a period of 235d. (page 96)
Observed:
W Cet (HD 224960) is a magnitude 7.1 variable star. Magnitude ranges from 14.8 to 7.1 ( $\Delta$ mag. 7.7) with a period of 351d. (page 72)
Observed:
o Cet (Mira, HD 14386) is a magnitude 2.0 variable star. Magnitude ranges from 10.1 to 2.0 ( $\Delta$ mag. 8.1 ) with a period of 332 d . (page 96)

Observed:

## Chamaeleon

C109 (NGC 3195) is a magnitude 11.6 planetary nebula. Angular size is $1.3^{\prime}$. (page 311)

Observed:

## Circinus

C88 (NGC 5823) is a magnitude 7.9 open cluster. Angular size is $10^{\prime}$. (page 227)
Observed:

## Columba

C73 (NGC 1851) is a magnitude 7.3 globular cluster. Angular size is $11^{\prime}$. (page 133)
Observed:
T Col (HD 34897) is a magnitude 6.6 variable star. Magnitude ranges from 12.7 to 6.6 ( $\Delta$ mag. 6.1) with a period of 226d. (page 132)

## Observed:

## Coma Berenices

M91 (NGC 4548) is a magnitude 10.2 barred spiral galaxy. Angular size is $5^{\prime}$. (page 186)
Observed:
C35 (NGC 4889) is a magnitude 11.4 elliptical galaxy.
Angular size is $3 \times 2^{\prime}$. (page 177)
Observed:
M53 (NGC 5024) is a magnitude 7.6 globular cluster.
Angular size is $13^{\prime}$. (page 178)
Observed:
M85 (NGC 4382) is a magnitude 9.1 lenticular galaxy. Angular size is $7 \times 5^{\prime}$. (page 185)
Observed:
C38 (Needle Galaxy, NGC 4565) is a magnitude 9.6 barred spiral galaxy. Angular size is $16 \times 3{ }^{\prime}$. (page 177)

Observed:
C36 (Koi Fish Galaxy, NGC 4559) is a magnitude 9.8 spiral galaxy. Angular size is $10 \times 4{ }^{\prime}$. (page 184)
Observed:
M100 (NGC 4321) is a magnitude 9.3 spiral galaxy.
Angular size is 7x6'. (page 185)
Observed:

M64 (Black Eye Galaxy, NGC 4826) is a magnitude 8.5 spiral galaxy. Angular size is $9 \times 5$ '. (page 184)

Observed:
M88 (NGC 4501) is a magnitude 9.6 spiral galaxy. Angular size is $7 \mathrm{x} 4^{\prime}$. (page 179)
Observed:
M98 (NGC 4192) is a magnitude 10.1 spiral galaxy.
Angular size is $10 \times 3{ }^{\prime}$. (page 179)
Observed:
M99 (NGC 4254) is a magnitude 9.9 spiral galaxy. Angular size is 5'. (page 186)
Observed:

## Corona Australis

C68 (NGC 6729) is a magnitude 9.7 bright nebula.
Angular size is $1.0^{\prime}$. (page 263)
Observed:
C78 (Cacciatore Cluster, NGC 6541) is a magnitude 6.6 globular cluster. Angular size is 13 '. (page 269)

Observed:

## Corona Borealis

R CrB (HD 141527) is a magnitude 5.71 variable star. Magnitude ranges from 14.8 to 5.71 ( $\Delta$ mag. 9.1). (page 214)
Observed:
S CrB (HD 136753) is a magnitude 5.8 variable star. Magnitude ranges from 14.1 to 5.8 ( $\Delta$ mag. 8.3) with a period of 360d. (page 211)
Observed:
T CrB (HD 143454) is a magnitude 2.0 variable star. Magnitude ranges from 10.8 to 2.0 ( $\Delta$ mag. 8.8) with a period of 80 y . (page 215)
Observed:
U CrB (HD 136175) is a magnitude 7.66 variable star. Magnitude ranges from 8.79 to 7.66 ( $\Delta$ mag. 1.1) with a period of 3.45220 d . (page 211)

Observed:

V CrB (HD 141826) is a magnitude 6.9 variable star. Magnitude ranges from 12.6 to 6.9 ( $\Delta$ mag. 5.7) with a period of 358d. (page 210)
Observed:
W CrB (HD 146560) is a magnitude 7.8 variable star. Magnitude ranges from 14.3 to 7.8 ( $\Delta$ mag. 6.5) with a period of 238d. (page 210)
Observed:

## Corvus

C60 (Antennae Galaxies, NGC 4038) is a magnitude 11.3 spiral galaxy. Angular size is $2.6 \times 1.8^{\prime}$. (page 193)

Observed:
C61 (Antennae Galaxies, NGC 4039) is a magnitude 13.0 spiral galaxy. Angular size is $3.2 \times 2.2^{\prime}$. (page 194)

Observed:
R Crv (HD 107199) is a magnitude 6.7 variable star. Magnitude ranges from 14.4 to 6.7 ( $\Delta$ mag. 7.7) with a period of 317 d . (page 194)

Observed:

## Crux

C99 (Coalsack Nebula, TGU H1867) is a magnitude 20.0 dark nebula. Angular size is $400 \times 300$ '. (page 205)

Observed:
C94 (Jewel Box, NGC 4755) is a magnitude 4.2 open cluster. Angular size is 10 '. (page 203)
Observed:
C98 (Coalsack Cluster, NGC 4609) is a magnitude 6.9 open cluster. Angular size is $5^{\prime}$. (page 201)

Observed:
R Cru (HD 107805) is a magnitude 6.4 variable star. Magnitude ranges from 7.23 to 6.4 ( $\Delta$ mag. 0.8 ) with a period of 5.82575 d . (page 201)

Observed:

## Cygnus

C19 (Cocoon Nebula, IC 5146) is a magnitude 10.0 bright nebula. Angular size is $12^{\prime}$. (page 279)
Observed:
C20 (North America Nebula, NGC 7000) is a magnitude 6.0 bright nebula. Angular size is $120 \times 100$ '. (page 275)
Observed:
C27 (Crescent Nebula, NGC 6888) is a magnitude 7.5 bright nebula. Angular size is $20 \times 10^{\prime}$. (page 280) Observed:
M29 (NGC 6913) is a magnitude 7.1 open cluster. Angular size is 7'. (page 276)
Observed:
M39 (NGC 7092) is a magnitude 4.6 open cluster. Angular size is $32^{\prime}$. (page 279)
Observed:
NGC 6871 is a magnitude 5.2 open cluster. Angular size is $20^{\prime}$. (page 280)
Observed:
C15 (Blinking Planetary, NGC 6826) is a magnitude 9.8 planetary nebula. Angular size is $0.5^{\prime}$. (page 274)

Observed:
C33 (East Veil Nebula, NGC 6992/5) is a magnitude 8.0 supernova remnant. Angular size is $60 \mathrm{x} 8^{\prime}$. (page 277)

Observed:
C34 (West Veil Nebula, NGC 6960) is a magnitude 8.0 supernova remnant. Angular size is $70 \times 6$ '. (page 281)

Observed:
CH Cyg (HD 182917) is a magnitude 5.6 variable star. Magnitude ranges from 8.49 to 5.6 ( $\Delta$ mag. 2.9). (page 234)
Observed:

R Cyg (HD 185456) is a magnitude 6.1 variable star. Magnitude ranges from 14.4 to 6.1 ( $\Delta$ mag. 8.3) with a period of 426d. (page 278)
Observed:
RT Cyg (HD 186686) is a magnitude 6.0 variable star. Magnitude ranges from 13.1 to 6.0 ( $\Delta$ mag. 7.1) with a period of 190d. (page 274)
Observed:
U Cyg (HD 193680) is a magnitude 5.9 variable star. Magnitude ranges from 12.1 to 5.9 ( $\Delta$ mag. 6.2) with a period of 463d. (page 275)

## Observed:

X Cyg (HD 197572) is a magnitude 5.85 variable star. Magnitude ranges from 6.91 to 5.85 ( $\Delta$ mag. 1.1) with a period of 16.38633 d . (page 276)

## Observed:

$\chi \operatorname{Cyg}$ (HD 187796) is a magnitude 3.3 variable star. Magnitude ranges from 14.2 to 3.3 ( $\Delta$ mag. 10.9) with a period of 408d. (page 281)

Observed:

## Delphinus

C42 (NGC 7006) is a magnitude 10.6 globular cluster. Angular size is $2.8^{\prime}$. (page 289)
Observed:
C47 (NGC 6934) is a magnitude 8.9 globular cluster. Angular size is 5.9'. (page 289)

## Observed:

EU Del (HD 196610) is a magnitude 5.79 variable star. Magnitude ranges from 6.9 to 5.79 ( $\Delta$ mag. 1.1) with a period of 59.7 d . (page 288)

## Observed:

## Dorado

C103 (Tarantula Nebula, NGC 2070) is a magnitude 4.0 bright nebula. Angular size is $40 \times 25{ }^{\prime}$. (page 308)

Observed:
LMC is a magnitude 0.91 Magellanic barred spiral galaxy. Angular size is $645^{\prime}$. (page 310)
Observed:
$\beta$ Dor (HD 037350 ) is a magnitude 3.46 variable star. Magnitude ranges from 4.08 to 3.46 ( $\Delta$ mag. 0.6) with a period of 9.8426 d . (page 134)
Observed:

## Draco

RY $\operatorname{Dra}(H D 112559)$ is a magnitude 6.0 carbon star. Magnitude ranges from 8.2 to 6.0 ( $\Delta$ mag. 2.2) with a period of 170d. (page 167)
Observed:
UX $\operatorname{Dra}$ (HD 183556) is a magnitude 6.2 carbon star. Magnitude ranges from 7.0 to 6.2 ( $\Delta \mathrm{mag} .0 .8$ ) with a period of 170d. (page 50)
Observed:
M102 (NGC 5866) is a magnitude 9.9 lenticular galaxy. Angular size is $5 \times 2^{\prime}$. (page 208)
Observed:
C6 (Cat's Eye Nebula, NGC 6543) is a magnitude 8.8 planetary nebula. Angular size is $0.3^{\prime}$. (page 234)
Observed:
C3 (NGC 4236) is a magnitude 9.7 barred spiral galaxy. Angular size is $22 \times 7^{\prime}$. (page 52 )
Observed:

## Eridanus

NGC 1291 is a magnitude 9.39 barred spiral galaxy. Angular size is $10^{\prime}$. (page 101) Observed:

NGC 1535 (Cleopatra's Eye Nebula) is a magnitude 9.5 planetary nebula. Angular size is $18{ }^{\prime \prime}$. The central star is magnitude 12.2. (page 97)
Observed:
T Eri (HD 024754) is a magnitude 7.2 variable star. Magnitude ranges from 13.2 to 7.2 ( $\Delta$ mag. 6.0) with a period of 252d. (page 97)

Observed:

## Fornax

Fornax Dwarf is a magnitude 9.3 irregular galaxy. Angular size is $17^{\prime}$. (page 100) Observed:

NGC 1316 is a magnitude 9.42 lenticular galaxy. Angular size is $12^{\prime}$. (page 101)
Observed:
NGC 1360 (Robin's Egg Nebula) is a magnitude 9.5 planetary nebula. Angular size is $380{ }^{\prime \prime}$. The central star is magnitude 11.4. (page 98)

Observed:
C67 (NGC 1097) is a magnitude 9.2 barred spiral galaxy. Angular size is $9 \times 6{ }^{\prime}$. (page 100)
Observed:
R For (HIP 011582) is a magnitude 7.5 variable star. Magnitude ranges from 13.0 to 7.5 ( $\Delta$ mag. 5.5) with a period of 389 d . (page 98)
Observed:

## Gemini

M35 (NGC 2168) is a magnitude 5.3 open cluster. Angular size is $28^{\prime}$. (page 113)

## Observed:

C39 (Eskimo Nebula, NGC 2392) is a magnitude 9.9 planetary nebula. Angular size is $0.8^{\prime}$. (page 117) Observed:
R Gem (HD 053791) is a magnitude 6.0 variable star. Magnitude ranges from 14.0 to 6.0 ( $\Delta$ mag. 8.0) with a period of 370 d . (page 113)
Observed:

## Grus

S Gru (HD 212539) is a magnitude 6.0 variable star. Magnitude ranges from 15.0 to 6.0 ( $\Delta$ mag. 9.0) with a period of 402d. (page 302)
Observed:

## Hercules

M13 (Great Hercules Globular, NGC 6205) is a magnitude 5.8 globular cluster. Angular size is $17^{\prime}$. (page 236)
Observed:
M92 (NGC 6341) is a magnitude 6.4 globular cluster. Angular size is $11^{\prime}$. (page 235)

Observed:

S Her (HD 152276) is a magnitude 6.4 variable star. Magnitude ranges from 13.8 to 6.4 ( $\Delta$ mag. 7.4) with a period of 307 d . (page 240)
Observed:
U Her (HD 148206) is a magnitude 6.4 variable star. Magnitude ranges from 13.4 to 6.4 ( $\Delta$ mag. 7.0) with a period of 406d. (page 216)
Observed:
X Her (HD 144205) is a magnitude 6.0 variable star. Magnitude ranges from 7.0 to $6.0(\Delta \mathrm{mag} .1 .0)$ with a period of 95.0 d . (page 209)
Observed:
$\mathrm{g} \operatorname{Her}$ (HD 148783) is a magnitude 4.3 variable star. Magnitude ranges from 6.3 to 4.3 ( $\Delta$ mag. 2.0) with a period of 89.2 d. (page 209)

Observed:
u $\operatorname{Her}(H D 156633)$ is a magnitude 4.69 variable star. Magnitude ranges from 5.37 to 4.69 ( $\Delta$ mag. 0.7) with a period of 2.05103 d . (page 237)

Observed:

## Horologium

C87 (NGC 1261) is a magnitude 8.4 globular cluster. Angular size is 7 '. (page 103)

Observed:
R Hor (HD 018242) is a magnitude 4.7 variable star. Magnitude ranges from 14.3 to 4.7 ( $\Delta$ mag. 9.6) with a period of 408 d . (page 102)

Observed:
U Hor (HD 024607) is a magnitude 7.8 variable star. Magnitude ranges from 15.1 to 7.8 ( $\Delta$ mag. 7.3) with a period of 348d. (page 102)

Observed:

## Hydra

U Hya (HD 92055) is a magnitude 4.7 carbon star. Magnitude ranges from 6.2 to 4.7 ( $\Delta$ mag. 1.5). (page 193)

Observed:

V Hya (SAO 179278) is a magnitude 6.5 carbon star. Magnitude ranges from 12.0 to 6.5 ( $\Delta$ mag. 5.5) with a period of 533d. (page 195)
Observed:
C66 (NGC 5694) is a magnitude 10.2 globular cluster. Angular size is 3.6 '. (page 221)
Observed:
M68 (NGC 4590) is a magnitude 7.8 globular cluster. Angular size is $12^{\prime}$. (page 196)
Observed:
M48 (NGC 2548) is a magnitude 5.5 open cluster. Angular size is $54^{\prime}$. (page 147)
Observed:
C59 (Ghost of Jupiter, NGC 3242) is a magnitude 8.6 planetary nebula. Angular size is $0.4^{\prime}$. (page 149)
Observed:
M83 (Southern Pinwheel, NGC 5236) is a magnitude 7.6 spiral galaxy. Angular size is $11 \times 10^{\prime}$. (page 225)

Observed:
R Hya (HD 117287) is a magnitude 3.5 variable star. Magnitude ranges from 10.9 to 3.5 ( $\Delta$ mag. 7.4) with a period of 389 d . (page 195)

Observed:
S Hya (HD 076011) is a magnitude 7.2 variable star. Magnitude ranges from 13.3 to 7.2 ( $\Delta$ mag. 6.1) with a period of 257 d . (page 143)
Observed:

## Lacerta

C16 (NGC 7243) is a magnitude 6.4 open cluster.
Angular size is $21^{\prime}$. (page 278)
Observed:

## Leo

M95 (NGC 3351) is a magnitude 9.7 barred spiral galaxy. Angular size is $4 \times 3^{\prime}$. (page 189)

Observed:
M105 (NGC 3379) is a magnitude 9.3 elliptical galaxy. Angular size is $2^{\prime}$. (page 180)
Observed:

C40 (NGC 3626) is a magnitude 10.9 barred spiral galaxy. Angular size is $3 \times 2^{\prime}$. (page 178)
Observed:
M65 (NGC 3623 ) is a magnitude 9.3 spiral galaxy. Angular size is $8 \times 1.5^{\prime}$. (page 187) Observed:
M66 (NGC 3627) is a magnitude 8.9 spiral galaxy. Angular size is $8 \times 2.5^{\prime}$. (page 187) Observed:

M96 (NGC 3368) is a magnitude 9.2 spiral galaxy. Angular size is $6 \times 4^{\prime}$. (page 182)

Observed:
R Leo (HD 084748) is a magnitude 4.4 variable star. Magnitude ranges from 11.3 to 4.4 ( $\Delta$ mag. 6.9) with a period of 310 d . (page 142)

## Observed:

## Leo Minor

R LMi (HD 084346) is a magnitude 6.3 variable star. Magnitude ranges from 13.2 to 6.3 ( $\Delta$ mag. 6.9) with a period of 372 d . (page 138)

## Observed:

## Lepus

M79 (NGC 1904) is a magnitude 7.7 globular cluster. Angular size is 9'. (page 126)
Observed:
R Lep (HD 031996) is a magnitude 5.5 variable star. Magnitude ranges from 11.7 to 5.5 ( $\Delta$ mag. 6.2) with a period of 427d. (page 125)
Observed:
RX Lep (HD 033664) is a magnitude 5.0 variable star. Magnitude ranges from 7.4 to $5.0(\Delta \mathrm{mag} .2 .4)$ with a period of 60 d . (page 128)
Observed:

## Lynx

Y Lyn (HD 58521) is a magnitude 6.9 carbon star. Magnitude ranges from 7.5 to $6.9(\Delta$ mag. 0.6$)$ with a period of 110 d . (page 106)

Observed:

C25 (Intergalactic Wanderer, NGC 2419) is a magnitude 10.4 globular cluster. Angular size is $6^{\prime}$. (page 137)
Observed:

## Lyra

M56 (NGC 6779) is a magnitude 8.3 globular cluster. Angular size is 7'. (page 237)
Observed:
M57 (Ring Nebula, NGC 6720) is a magnitude 8.8 planetary nebula. Angular size is $1.4 \times 1.0$ '. (page 236)

Observed:
$R \operatorname{Lyr}$ (HD 175865) is a magnitude 3.88 variable star. Magnitude ranges from 5.0 to 3.88 ( $\Delta \mathrm{mag}$. 1.1) with a period of 46d. (page 235)
Observed:

## Microscopium

U Mic (HD 194814) is a magnitude 7.0 variable star. Magnitude ranges from 14.4 to 7.0 ( $\Delta$ mag. 7.4) with a period of 334d. (page 301)
Observed:

## Monoceros

C46 (Hubble's Variable Nebula, NGC 2261) is a magnitude 10.0 bright nebula. Angular size is $2 \times 1$ '. (page 119)

Observed:
C49 (NGC 2237-9) is a magnitude 9.0 bright nebula. Angular size is $80 \times 60^{\prime}$. (page 119)

Observed:
C50 (Rosette Nebula, NGC 2244) is a magnitude 4.8 open cluster. Angular size is 24 . (page 116) Observed:

C54 (NGC 2506) is a magnitude 7.6 open cluster. Angular size is 7 '. (page 150)

Observed:
M50 (NGC 2323) is a magnitude 6.3 open cluster. Angular size is $16^{\prime}$. (page 124)

Observed:

NGC 2232 is a magnitude 3.9 open cluster. Angular size is $30^{\prime}$. (page 127)
Observed:
NGC 2301 is a magnitude 6.0 open cluster. Angular size is $12^{\prime}$. (page 116)
Observed:
U Mon (HD 059693) is a magnitude 6.1 variable star. Magnitude ranges from 8.8 to 6.1 ( $\Delta$ mag. 2.7) with a period of 91.3d. (page 147)
Observed:
V Mon (HD 044639) is a magnitude 6.0 variable star. Magnitude ranges from 13.9 to 6.0 ( $\Delta$ mag. 7.9) with a period of 341d. (page 127)
Observed:

## Musca

C105 (The Southern Butterfly, NGC 4833) is a magnitude 7.3 globular cluster. Angular size is $14^{\prime}$. (page 311)
Observed:
C108 (NGC 4372) is a magnitude 7.8 globular cluster. Angular size is 19 '. (page 309)
Observed:

## Norma

C89 (S Norma Cluster, NGC 6087) is a magnitude 5.4 open cluster. Angular size is 12 . (page 227)
Observed:
NGC 6067 is a magnitude 5.6 open cluster. Angular size is $13^{\prime}$. (page 230)
Observed:
R Nor (HD 138743) is a magnitude 6.5 variable star.
Magnitude ranges from 13.9 to 6.5 ( $\Delta$ mag. 7.4) with a period of 508d. (page 226)
Observed:
T Nor (HD 140041) is a magnitude 6.2 variable star. Magnitude ranges from 13.6 to 6.2 ( $\Delta$ mag. 7.4) with a period of 241d. (page 226)

[^0]
## Octans

R Oct (HD 040857) is a magnitude 6.4 variable star. Magnitude ranges from 13.2 to 6.4 ( $\Delta$ mag. 6.8) with a period of 405 d . (page 310)
Observed:

## Ophiuchus

M10 (NGC 6254) is a magnitude 6.6 globular cluster. Angular size is $15^{\prime}$. (page 247)
Observed:
M107 (NGC 6171) is a magnitude 7.9 globular cluster.
Angular size is $10^{\prime}$. (page 249)
Observed:
M12 (NGC 6218) is a magnitude 6.7 globular cluster. Angular size is $15^{\prime}$. (page 247)
Observed:
M14 (NGC 6402) is a magnitude 7.6 globular cluster. Angular size is $12^{\prime}$. (page 252)
Observed:
M19 (NGC 6273) is a magnitude 6.8 globular cluster. Angular size is $14^{\prime}$. (page 257)
Observed:
M62 (NGC 6266) is a magnitude 6.5 globular cluster. Angular size is $14^{\prime}$. (page 261)
Observed:
M9 (NGC 6333) is a magnitude 7.7 globular cluster. Angular size is $9.3^{\prime}$. (page 255)
Observed:
IC 4665 is a magnitude 4.2 open cluster. Angular size is $41^{\prime}$. (page 243)
Observed:
NGC 6633 is a magnitude 4.6 open cluster. Angular size is $27^{\prime}$. (page 242)
Observed:
RS Oph (HD 162214) is a magnitude 4.3 variable star. Magnitude ranges from 12.5 to 4.3 ( $\Delta$ mag. 8.2). (page 253)

## Observed:

V Oph (HD 148182) is a magnitude 7.3 variable star. Magnitude ranges from 11.6 to 7.3 ( $\Delta$ mag. 4.3) with a period of 297 d . (page 220)
Observed:
X Oph (HD 172171) is a magnitude 5.9 variable star. Magnitude ranges from 9.2 to 5.9 ( $\Delta$ mag. 3.3) with a period of 329 d . (page 241 )

Observed:

## Orion

Betelgeuse (HD 39801) is a magnitude 0.4 carbon star. Magnitude ranges from 1.3 to 0.4 ( $\Delta$ mag. 0.9 ). (page 115)

Observed:
W Ori (HD 32736) is a magnitude 6.5 carbon star. Magnitude ranges from 10.0 to 6.5 ( $\Delta$ mag. 3.5) with a period of 210d. (page 120)

Observed:
M42 (Great Nebula in Orion, NGC 1976) is a magnitude 4.0 diffuse nebula. Angular size is $85 \times 60$ '. (page 128)

Observed:
M43 (De Mairan's Nebula, NGC 1982) is a magnitude 9.0 diffuse nebula. Angular size is 20x15'. (page 124)

Observed:
M78 (NGC 2068) is a magnitude 8.3 diffuse nebula. Angular size is $8 \times 6$ '. (page 120)

Observed:
NGC 1662 is a magnitude 6.4 open cluster. Angular size is 20 '. (page 115)

Observed:
NGC 1981 is a magnitude 4.6 open cluster. Angular size is 25 '. (page 123)

Observed:
NGC 2169 is a magnitude 5.9 open cluster. Angular size is $7^{\prime}$. (page 118)

Observed:

U Ori (HD 039816) is a magnitude 4.8 variable star. Magnitude ranges from 13.0 to 4.8 ( $\Delta$ mag. 8.2) with a period of 368d. (page 118)
Observed:
VV Ori (HD 036695) is a magnitude 5.31 variable star. Magnitude ranges from 5.66 to 5.31 ( $\Delta$ mag. 0.4) with a period of 1.48538 d . (page 123)
Observed:

## Pavo

C93 (Pavo Globular Cluster, NGC 6752) is a magnitude 5.4 globular cluster. Angular size is 20 '. (page 265)

Observed:
C101 (NGC 6744) is a magnitude 9.0 barred spiral galaxy. Angular size is $16 \times 10^{\prime}$. (page 270)
Observed:
S Pav (HD 187835) is a magnitude 6.6 variable star. Magnitude ranges from 10.4 to 6.6 ( $\Delta$ mag. 3.8) with a period of 381d. (page 303)

Observed:

## Pegasus

TW Peg (HD 209598) is a magnitude 7.0 carbon star. Magnitude ranges from 9.2 to 7.0 ( $\Delta$ mag. 2.2) with a period of 956d. (page 285)

Observed:
M15 (Great Pegasus Globular, NGC 7078) is a magnitude 6.2 globular cluster. Angular size is $12^{\prime}$. (page 287)

Observed:
C30 (Deer Lick Group, NGC 7331) is a magnitude 9.5 barred spiral galaxy. Angular size is $11 \times 4{ }^{\prime}$. (page 66)

Observed:
C43 (Little Sombrero Galaxy, NGC 7814) is a magnitude 10.5 barred spiral galaxy. Angular size is $6 \times 2^{\prime}$. (page 68)

## Observed:

C44 (NGC 7479) is a magnitude 11.0 barred spiral galaxy. Angular size is $4 \times 3$ '. (page 69)
Observed:
R Peg (HD 218292) is a magnitude 6.9 variable star. Magnitude ranges from 13.8 to 6.9 ( $\Delta$ mag. 6.9) with a period of 378 d . (page 69)
Observed:

## Perseus

C14 (Double Cluster, NGC 869/884) is a magnitude 4.3 open cluster. Angular size is $60 \times 30^{\prime}$. (page 82)

Observed:
M34 (NGC 1039) is a magnitude 5.5 open cluster.
Angular size is $35^{\prime}$. (page 83)
Observed:
NGC 1528 is a magnitude 6.4 open cluster. Angular size is $24^{\prime}$. (page 82 )
Observed:
NGC 1545 is a magnitude 6.2 open cluster. Angular size is $18^{\prime}$. (page 86 )
Observed:
M76 (Little Dumbbell Nebula, NGC 650) is a magnitude 10.1 planetary nebula. Angular size is $3 \times 2$ '. (page 86)
Observed:
C24 (Perseus A, NGC 1275) is a magnitude 11.6 seyfert galaxy. Angular size is $2^{\prime}$. (page 83)
Observed:
$\mathrm{X} \operatorname{Per}(\mathrm{HD} 024534)$ is a magnitude 6.03 variable star. Magnitude ranges from 7.0 to 6.03 ( $\Delta$ mag. 1.0). (page 84)
Observed:

## Pictor

R Pic (HD 030551) is a magnitude 6.35 variable star. Magnitude ranges from 10.1 to 6.35 ( $\Delta$ mag. 3.8) with a period of 171d. (page 133)

[^1]
## Pisces

Z Psc (HD 7561) is a magnitude 7.0 carbon star. Magnitude ranges from 7.9 to 7.0 ( $\Delta$ mag. 0.9 ) with a period of 144 d . (page 68)
Observed:
M74 (NGC 628) is a magnitude 9.4 spiral galaxy. Angular size is 10 '. (page 91)
Observed:
R Psc (HD 009203) is a magnitude 7.0 variable star. Magnitude ranges from 14.8 to 7.0 ( $\Delta$ mag. 7.8) with a period of 345 d . (page 93)

## Observed:

TX Psc (HD 223075) is a magnitude 4.79 variable star. Magnitude ranges from 5.2 to 4.79 ( $\Delta$ mag. 0.4). (page 70)
Observed:

## Puppis

NGC 2298 (NGC) is a magnitude 9.29 globular cluster. Angular size is $5^{\prime}$. (page 132)
Observed:
C71 (NGC 2477) is a magnitude 5.8 open cluster. Angular size is $27^{\prime}$. (page 155)
Observed:
M46 (NGC 2437) is a magnitude 6.0 open cluster.
Angular size is $27^{\prime}$. (page 151)
Observed:
M47 (NGC 2422) is a magnitude 5.2 open cluster. Angular size is $30^{\prime}$. (page 148)
Observed:
M93 (NGC 2447) is a magnitude 6.0 open cluster. Angular size is $22^{\prime}$. (page 152)
Observed:
NGC 2423 is a magnitude 6.7 open cluster. Angular size is $19^{\prime}$. (page 151)
Observed:
NGC 2451 is a magnitude 2.8 open cluster. Angular size is $45^{\prime}$. (page 159)
Observed:

NGC 2527 is a magnitude 6.5 open cluster. Angular size is $22^{\prime}$. (page 149)
Observed:
NGC 2539 is a magnitude 6.5 open cluster. Angular size is $22^{\prime}$. (page 148)

Observed:
NGC 2546 is a magnitude 6.3 open cluster. Angular size is $41^{\prime}$. (page 155)

Observed:

## Reticulum

NGC 1313 is a magnitude 9.2 barred spiral galaxy. Angular size is $9^{\prime}$. (page 103)
Observed:

## Sagitta

M71 (NGC 6838) is a magnitude 8.2 globular cluster. Angular size is $7^{\prime}$. (page 286)
Observed:
U Sge (HD 181182) is a magnitude 6.45 variable star. Magnitude ranges from 9.28 to 6.45 ( $\Delta$ mag. 2.8) with a period of 3.38062 d . (page 240)
Observed:
WZ Sge (HV 03518) is a magnitude 7.0 variable star. Magnitude ranges from 15.53 to 7.0 ( $\Delta$ mag. 8.5) with a period of 33 y . (page 286)
Observed:

## Sagittarius

AQ Sgr (HD 184283) is a magnitude 6.6 carbon star. Magnitude ranges from 7.7 to 6.6 ( $\Delta$ mag. 1.1) with a period of 200d. (page 296)
Observed:
M17 (Omega Nebula, NGC 6618) is a magnitude 7.0 diffuse nebula. Angular size is $11^{\prime}$. (page 249) Observed:

M20 (Trifid Nebula, NGC 6514) is a magnitude 9.0 diffuse nebula. Angular size is $28^{\prime}$. (page 256) Observed:

M8 (Lagoon Nebula, NGC 6523) is a magnitude 6.0 diffuse nebula. Angular size is $90 \times 40^{\prime}$. (page 257)

Observed:
M22 (Sagittarius Cluster, NGC 6656) is a magnitude 5.1 globular cluster. Angular size is $24^{\prime}$. (page 251)

Observed:
M28 (NGC 6626) is a magnitude 6.8 globular cluster. Angular size is $11^{\prime}$. (page 252)
Observed:
M54 (NGC 6715) is a magnitude 7.6 globular cluster. Angular size is $9^{\prime}$. (page 266)
Observed:
M55 (NGC 6809) is a magnitude 6.3 globular cluster. Angular size is $19^{\prime}$. (page 300)
Observed:
M69 (NGC 6637) is a magnitude 7.6 globular cluster. Angular size is $7^{\prime}$. (page 267)
Observed:
M70 (NGC 6681) is a magnitude 7.9 globular cluster. Angular size is $8^{\prime}$. (page 262)
Observed:
M75 (NGC 6864) is a magnitude 8.5 globular cluster. Angular size is 6 '. (page 295)
Observed:
C57 (Barnard's Galaxy, NGC 6822) is a magnitude
9.3 irregular galaxy. Angular size is $10 \times 9^{\prime}$. (page 294)

Observed:
M24 (Sagittarius Star Cloud, IC 4715) is a magnitude 4.6 star cloud. Angular size is $90^{\prime}$. (page 250)

Observed:
M18 (NGC 6613) is a magnitude 7.5 open cluster. Angular size is $9^{\prime}$. (page 255)
Observed:
M21 (NGC 6531) is a magnitude 6.5 open cluster.
Angular size is $13^{\prime}$. (page 251)
Observed:

M23 (NGC 6494) is a magnitude 6.9 open cluster. Angular size is $27^{\prime}$. (page 250)
Observed:
M25 (IC 4725) is a magnitude 6.5 open cluster.
Angular size is $40^{\prime}$. (page 256)
Observed:
RR $\operatorname{Sgr}$ (HD 188378) is a magnitude 5.4 variable star. Magnitude ranges from 14.0 to 5.4 ( $\Delta$ mag. 8.6) with a period of 336 d . (page 300)
Observed:
RT $\operatorname{Sgr}(H D$ 192702) is a magnitude 6.0 variable star. Magnitude ranges from 14.1 to $6.0(\Delta$ mag. 8.1) with a period of 306d. (page 301)

Observed:
RU $\operatorname{Sgr}$ (HD 188813) is a magnitude 6.0 variable star. Magnitude ranges from 13.8 to 6.0 ( $\Delta$ mag. 7.8) with a period of 240d. (page 302)
Observed:
RY $\operatorname{Sgr}(H D 180093)$ is a magnitude 5.8 variable star. Magnitude ranges from 14.0 to 5.8 ( $\Delta$ mag. 8.2) with a period of Period. (page 263)
Observed:

## Sc

M11 (Wild Duck Cluster, NGC 6705) is a magnitude 6.3 open cluster. Angular size is 14'. (page 248)

Observed:

## Scorpius

Antares (HD 148478) is a magnitude 0.9 carbon star. Magnitude ranges from 1.0 to 0.9 ( $\Delta$ mag. 0.1 ). (page 222)
Observed:
M4 (NGC 6121) is a magnitude 5.6 globular cluster.
Angular size is 26.3 '. (page 222)
Observed:
M80 (NGC 6093) is a magnitude 7.3 globular cluster.
Angular size is 10 '. (page 221)
Observed:

C75 (NGC 6124) is a magnitude 5.8 open cluster. Angular size is $29^{\prime}$. (page 225)
Observed:
C76 (Northern Jewel Box, NGC 6231) is a magnitude 2.6 open cluster. Angular size is $15^{\prime}$. (page 268) Observed:

M6 (Butterfly Cluster, NGC 6405) is a magnitude 4.2 open cluster. Angular size is $25^{\prime}$. (page 266)
Observed:
M7 (Ptolemy's Cluster, NGC 6475) is a magnitude 3.3 open cluster. Angular size is $80^{\prime}$. (page 267)

Observed:
NGC 6322 is a magnitude 6.0 open cluster. Angular size is $10^{\prime}$. (page 264)
Observed:
NGC 6416 is a magnitude 5.7 open cluster. Angular size is 18 '. (page 262)

Observed:
C69 (Bug Nebula, NGC 6302) is a magnitude 12.8 planetary nebula. Angular size is $0.8^{\prime}$. (page 268) Observed:

RR Sco (HD 152783) is a magnitude 5.0 variable star. Magnitude ranges from 12.4 to 5.0 ( $\Delta$ mag. 7.4) with a period of 281d. (page 261)

Observed:

## Sculptor

R Scl (HD 8879) is a magnitude 6.1 carbon star. Magnitude ranges from 8.8 to 6.1 ( $\Delta$ mag. 2.7) with a period of 363 d . (page 76)
Observed:
C72 (String of Pearls, NGC 55) is a magnitude 8.2 barred spiral galaxy. Angular size is $32 \times 6$ '. (page 77)

Observed:
C65 (Sculptor Galaxy, NGC 253) is a magnitude 7.1 spiral galaxy. Angular size is $25 \times 7^{\prime}$. (page 74)
Observed:

C70 (Southern Pinwheel Galaxy, NGC 300) is a magnitude 8.1 spiral galaxy. Angular size is 20x13'. (page 77)
Observed:
S Scl (HD 001115) is a magnitude 5.5 variable star. Magnitude ranges from 13.6 to 5.5 ( $\Delta$ mag. 8.1) with a period of 363d. (page 76)
Observed:

## Scutum

M26 (NGC 6694) is a magnitude 8.0 open cluster.
Angular size is $15^{\prime}$. (page 248)
Observed:
R Sct (HD 173819) is a magnitude 4.2 variable star. Magnitude ranges from 8.6 to 4.2 ( $\Delta$ mag. 4.4) with a period of 146.5 d . (page 253)
Observed:

## Serpens

M5 (NGC 5904) is a magnitude 5.6 globular cluster. Angular size is $17.4^{\prime}$. (page 218)
Observed:
IC 4756 is a magnitude 5.0 open cluster. Angular size is $52^{\prime}$. (page 243)
Observed:
M16 (Eagle Nebula, NGC 6611) is a magnitude 6.4 open cluster. Angular size is $7^{\prime}$. (page 254)
Observed:
NGC 6605 is a magnitude 6.0 open cluster. Angular size is $29^{\prime}$. (page 254)
Observed:
$R \operatorname{Ser}(H D 141850)$ is a magnitude 5.16 variable star. Magnitude ranges from 14.4 to 5.16 ( $\Delta$ mag. 9.2) with a period of 356d. (page 216)
Observed:
S Ser (HD 136695) is a magnitude 7.0 variable star. Magnitude ranges from 14.1 to 7.0 ( $\Delta$ mag. 7.1) with a period of 372d. (page 217)

[^2]
## Sextans

C53 (Spindle Galaxy, NGC 3115) is a magnitude 9.1 elliptical galaxy. Angular size is $8 \times 3^{\prime}$. (page 150) Observed:

## Taurus

C41 (Hyades, Mel 25) is a magnitude 1.0 open cluster. Angular size is 330 '. (page 91) Observed:
M45 (Pleiades, Mel 22) is a magnitude 1.6 open cluster. Angular size is $110^{\prime}$. (page 90) Observed:

NGC 1647 is a magnitude 6.4 open cluster. Angular size is $45^{\prime}$. (page 114)
Observed:
NGC 1746 is a magnitude 6.0 open cluster. Angular size is $42^{\prime}$. (page 117)
Observed:
M1 (Crab Nebula, NGC 1952) is a magnitude 8.4 supernova remnant. Angular size is $6 \times 4$ '. (page 114)

Observed:
R Tau (HD 028309) is a magnitude 7.6 variable star. Magnitude ranges from 15.8 to 7.6 ( $\Delta$ mag. 8.2) with a period of 321 d . (page 93)
Observed:

## Triangulum

M33 (Triangulum Galaxy, NGC 598) is a magnitude 5.7 spiral galaxy. Angular size is $73 \times 45^{\prime}$. (page 88)

Observed:
R Tri (HD 016210 ) is a magnitude 5.4 variable star. Magnitude ranges from 12.6 to 5.4 ( $\Delta$ mag. 7.2) with a period of 267d. (page 84)

## Observed:

## Triangulum Australe

C95 (NGC 6025) is a magnitude 5.1 open cluster. Angular size is $12^{\prime}$. (page 232)

## Observed:

## Tucana

SMC is a magnitude 2.7 Magellanic barred spiral galaxy. Angular size is $315^{\prime}$. (page 307)
Observed:
C104 (NGC 362) is a magnitude 6.6 globular cluster. Angular size is $13^{\prime}$. (page 307)
Observed:
C106 (47 Tucanae, NGC 104) is a magnitude 4.0 globular cluster. Angular size is $31^{\prime}$. (page 309) Observed:

## Ursa Major

M109 (NGC 3992) is a magnitude 9.8 barred spiral galaxy. Angular size is $7 \times 4^{\prime}$. (page 172)
Observed:
RT UMa (TYC 3431-229-1) is a magnitude 8.6 carbon star. Magnitude ranges from 9.6 to 8.6 ( $\Delta$ mag. 1.0). (page 137)

Observed:
VY UMa (HD 92839) is a magnitude 6.0 carbon star. Magnitude ranges from 6.6 to 6.0 ( $\Delta$ mag. 0.6 ). (page 49)

Observed:
M40 (Winnecke 4, Win 4) is a magnitude 8.4 double star. Angular size is $0.8^{\prime}$. (page 171)
Observed:
M82 (Cigar Galaxy, NGC 3034) is a magnitude 8.4 irregular galaxy. Angular size is $9 x 4$ '. (page 50) Observed:
M97 (Owl Nebula, NGC 3587) is a magnitude 9.9 planetary nebula. Angular size is $3 \times 3.3^{\prime}$. (page 168)

Observed:
M101 (Pinwheel Galaxy, NGC 5457) is a magnitude
7.9 spiral galaxy. Angular size is $22^{\prime}$. (page 208)

Observed:
M108 (NGC 3556) is a magnitude 10.0 spiral galaxy. Angular size is $8 \times 1^{\prime}$. (page 171)
Observed:

M81 (Bode's Galaxy, NGC 3031) is a magnitude 6.9 spiral galaxy. Angular size is $21 \times 10^{\prime}$. (page 53) Observed:

R UMa (HD 092763) is a magnitude 6.5 variable star. Magnitude ranges from 13.7 to 6.5 ( $\Delta$ mag. 7.2) with a period of 302d. (page 49)
Observed:
W UMa (HD 083950) is a magnitude 7.75 variable star. Magnitude ranges from 8.48 to 7.75 ( $\Delta$ mag. 0.7 ) with a period of 0.3336 d . (page 136)

Observed:
Z UMa (HD 103681) is a magnitude 6.2 variable star. Magnitude ranges from 9.4 to $6.2(\Delta$ mag. 3.2) with a period of 196d. (page 167)
Observed:

## Vela

C79 (NGC 3201) is a magnitude 6.7 globular cluster. Angular size is $18^{\prime}$. (page 156)
Observed:
IC 2395 is a magnitude 4.6 open cluster. Angular size is $8^{\prime}$. (page 160 )
Observed:
NGC 2547 is a magnitude 4.7 open cluster. Angular size is $20^{\prime}$. (page 156)
Observed:
NGC 2669 is a magnitude 6.1 open cluster. Angular size is 12 '. (page 161)
Observed:
NGC 3228 is a magnitude 6.0 open cluster. Angular size is $18^{\prime}$. (page 157)
Observed:
C74 (Eight Burst Nebula, NGC 3132) is a magnitude 8.2 planetary nebula. Angular size is $0.8^{\prime}$. (page 160)

Observed:

## Virgo

M58 (NGC 4579) is a magnitude 9.7 barred spiral galaxy. Angular size is $6 \times 5^{\prime}$. (page 182)

Observed:

SS Vir (HD 108105) is a magnitude 6.0 carbon star. Magnitude ranges from 9.6 to $6.0(\Delta$ mag. 3.6) with a period of 355 d . (page 190)
Observed:
C52 (NGC 4697) is a magnitude 9.3 elliptical galaxy. Angular size is $6 \times 3^{\prime}$. (page 192)

Observed:
M49 (NGC 4472) is a magnitude 8.4 elliptical galaxy. Angular size is $9 \times 7.5^{\prime}$. (page 190)
Observed:
M59 (NGC 4621) is a magnitude 9.6 elliptical galaxy. Angular size is $5 \times 4^{\prime}$. (page 189)

Observed:
M60 (NGC 4649) is a magnitude 8.8 elliptical galaxy. Angular size is 7x6'. (page 183)

Observed:
M87 (NGC 4486) is a magnitude 8.6 elliptical galaxy. Angular size is $7^{\prime}$. (page 188)

Observed:
M89 (NGC 4552) is a magnitude 9.8 elliptical galaxy. Angular size is $4^{\prime}$. (page 181)

Observed:
M84 (NGC 4374) is a magnitude 9.1 lenticular galaxy. Angular size is $5^{\prime}$. (page 188)

Observed:
M86 (NGC 4406) is a magnitude 8.9 lenticular galaxy. Angular size is $8 \times 6{ }^{\prime}$. (page 181) Observed:

M104 (Sombrero Galaxy, NGC 4594) is a magnitude 8.0 spiral galaxy. Angular size is $9 \times 44^{\prime}$. (page 192)

Observed:
M61 (NGC 4303) is a magnitude 9.7 spiral galaxy. Angular size is 6 '. (page 183)

Observed:
M90 (NGC 4569) is a magnitude 9.5 spiral galaxy.
Angular size is $10 \times 55^{\prime}$. (page 180)
Observed:
$\mathrm{S} \operatorname{Vir}(\mathrm{HD} 117833)$ is a magnitude 6.3 variable star. Magnitude ranges from 13.2 to 6.3 ( $\Delta$ mag. 6.9) with a period of 375 d . (page 220)

Observed:

## Vulpecula

C37 (20 Vulpeculae Cluster, NGC 6885) is a magnitude 5.7 open cluster. Angular size is $7^{\prime}$. (page 285)
Observed:
NGC 6940 is a magnitude 6.3 open cluster. Angular size is 31 '. (page 287)

Observed:
M27 (Dumbbell Nebula, NGC 6853) is a magnitude 7.4 planetary nebula. Angular size is $8 \times 6$ '. (page 288)

Observed:

# Bortle Light Pollution Scale 

The Bortle Light Pollution Scale was developed by John E Bortle in 2001 to categorize local light pollution levels into nine categories, with "one" being truly dark, pristine skies and "nine" being highly polluted inner city skies.

To help you judge your local light pollution, the following pages show selected regions at Bortle levels nine, seven five and three. The northern and southern polar regions are included, as are the more equatorial constellations of Leo, Scorpius, Orion and Pegasus. I also include two bright and easily recognized circumpolar constellations, Crux and Cassiopeia. Although only some Bortle levels are illustrated, it should be fairly simple to interpolate the remaining levels - for example if the sky is slightly better than the Bortle 7 chart but not quite as good as Bortle 5, then it is reasonable to estimate a level of Bortle 6 even though it is not separately illustrated.

The targets in this book are color-coded into three categories: blue targets have a magnitude of at least 7 and are the brightest and red targets are the faintest, no brighter than magnitude 9 , while orange targets fall between. This is only a rough guide, as an object might be relatively bright but very spread out (that is, having a low surface brightness) so it might be classified as a blue object but actually be somewhat more difficult.

In Bortle 9 skies, all targets are extremely difficult, with only a few of the brightest blue targets such as the Pleiades being easily visible. Finding the targets presents its own challenge as even the relatively bright signpost stars used in this book are at the edge of visibility. Very bright skies make galaxies in particular more difficult, while open clusters and globular clusters fare somewhat better. This is because higher magnification can be used to stretch out and dim the background, while stars, being point sources, are not significantly dimmed by higher magnifications. Likewise, colored stars like carbon stars can still be enjoyed regardless of light pollution.

Bortle 8 skies show a slight improvement, and most blue targets should be visible though a small
telescope. Other targets may only show on the best nights. Red targets will mostly prove elusive.

In Bortle 7 skies, the blue targets should present themselves beautifully in the eyepiece, and almost all other targets should at least be visible.

In Bortle 6 skies, orange targets should show well in your telescope, and the milky way should be visible. While these skies are regarded as relatively badly light polluted, a Bortle 6 sky nevertheless retains some traces of its true majesty.
Light pollution is worse nearer to the horizon, and stars are partially extinguished by the atmosphere at lower elevations, so it is not unusual to for different parts of the sky to be a different Bortle.

Aside from being a nuisance to stargazers, light pollution has serious consequences for ecosystems and human health. Light pollution can affect the reproductive cycles of animals, the migration of birds, and contributes to the catastrophic decline in the numbers of insects. Light intrusion at night can affect humans too, not only disrupting sleep but also increasing the risk of conditions such as obesity, depression, diabetes and breast cancer. In addition, unnecessary and inefficient lighting wastes energy and contributes to climate change.

The International Dark-Sky Association campaigns against light pollution:
https://www.darksky.org/

# Bortle Scale (Ursa Minor) 



## Bortle Scale (Octans)



Bortle 9 (limiting magnitude 4.0)


Bortle 5 (limiting magnitude 6.0)


## Bortle Scale (Cassiopeia)



Bortle 9 (limiting magnitude 4.0)
Bortle 5 (limiting magnitude 6.0)


## Bortle Scale (Orion)



Bortle 9 (limiting magnitude 4.0)


Bortle 7 (limiting magnitude 5.0)


Bortle 5 (limiting magnitude 6.0)


Bortle 3 (limiting magnitude 7.0)

## Bortle Scale (Leo)



Bortle 9 (limiting magnitude 4.0)
Bortle 5 (limiting magnitude 6.0)


## Bortle Scale (Scorpius)



Bortle 9 (limiting magnitude 4.0)


Bortle 5 (limiting magnitude 6.0)


Bortle 7 (limiting magnitude 5.0)


Bortle 3 (limiting magnitude 7.0)

## Bortle Scale (Pegasus)



## Bortle Scale (Crux)



Bortle 9 (limiting magnitude 4.0)


Bortle 7 (limiting magnitude 5.0)

Bortle 5 (limiting magnitude 6.0)



Bortle 3 (limiting magnitude 7.0)

## Northern Circumpolar Sky (1)



## Northern Circumpolar Sky (2)




RA: $161.25^{\circ} \mid 10 \mathrm{~h} 45.0^{\prime}-$ DEC: $67.4^{\circ} \mid 67^{\circ} 24^{\prime}$

VY UMa (HD 92839) is a magnitude 6.0 carbon star.
Magnitude ranges from 6.6 to 6.0 ( $\Delta$ mag. 0.6).
Travel from Merak to Dubhe in Ursa Major, then head toward Polaris for an equal distance.

VY UMa has less range in brightness than many carbon stars. Its B-V color index is a very red 2.4. The $\mathrm{B}-\mathrm{V}$ color index is the difference in magnitude of the star measured in blue (B) light and yellow-green light (V).

Also visible:
(1) M81 ( $6.9_{m}$ spiral galaxy)
(2) M82 (8.4m irregular galaxy)
(3) NGC 3077 (10.61m irregular galaxy)
(4) IC 2574 (10.8m Magellanic barred spiral galaxy)
(5) R UMa ( $6.5_{m}$ variable star)


RA: $161.16^{\circ} \mid 10 \mathrm{~h} 44.64^{\prime}-$ DEC: $68.78^{\circ} \mid 68^{\circ} 47^{\prime}$
R UMa (HD 092763) is a magnitude 6.5 variable star. Magnitude ranges from 13.7 to 6.5 ( $\Delta$ mag. 7.2 ) with a period of 302 d .

Travel from Merak to Dubhe in Ursa Major, then head toward Polaris by an one-and-a-half times the distance.

R UMa is just over a degree north of the carbon star VY UMa.

Also visible:
(1) M81 (6.9m spiral galaxy)
(2) M82 (8.4m irregular galaxy)
(3) NGC 3077 (10.61m irregular galaxy)
(4) IC 2574 (10.8m Magellanic barred spiral galaxy)
(5) VY UMa ( $6.0_{m}$ carbon star)


RA: $148.95^{\circ} \mid 9$ h $55.79^{\prime}-$ DEC: $69.68^{\circ} \mid 69^{\circ} 41^{\prime}$
M82 (Cigar Galaxy, NGC 3034) is a magnitude 8.4
irregular galaxy. Angular size is $9 \mathrm{x} 4^{\prime}$.
Draw a line from Phecda through Dubhe, and extend it an equal length further. This brings you very close to M81 and M82.

This is the nearest large starburst galaxy; its core contains nearly 200 titanic clusters with masses averaging 200,000 solar masses. The many supernovae that occur in this region have ejected huge quantities of hydrogen gas that can be seen as reddish streams in photographs. A 100 mm telescope can reveal the lumpy texture of this galaxy caused by the prominent dark dust clouds.

Also visible:
(1) M81 ( $6.9_{m}$ spiral galaxy)
(2) NGC 3077 ( $10.61_{m}$ irregular galaxy)
(3) IC 2574 ( $10.8_{m}$ Magellanic barred spiral galaxy)
(4) R UMa ( $6.5_{m}$ variable star)
(5) VY UMa ( $6.0_{m}$ carbon star)


RA: $290.4^{\circ} \mid 19 \mathrm{~h} 21.59^{\prime}-$ DEC: $76.6^{\circ} \mid 76^{\circ} 36^{\prime}$

UX Dra (HD 183556) is a magnitude 6.2 carbon star. Magnitude ranges from 7.0 to $6.2(\Delta$ mag. 0.8$)$ with a period of 170 d .
$5.5^{\circ} \mathrm{NE}$ from mag.3.69 chi Dra.

UX Dra is a deep red with a B-V color index of 2.68. Its absolute magnitude is around -2.25 .


RA: $15.58^{\circ} \mid$ 1h $2.3^{\prime}-$ DEC: $81.88^{\circ} \mid 81^{\circ} 53^{\prime}$
U Cep (HD 5679) is a magnitude 6.75 variable star.
Magnitude ranges from 9.24 to 6.75 ( $\Delta$ mag. 2.5 ) with a period of 2.49305 d .

A quarter of the distance from Polaris to Ruchbah in Cassiopeia.

This eclipsing binary system is also a visual double, with a faint companion separated by 14 " at position angle $62^{\circ}$.


Also visible:
(1) C 1 ( $8.1_{m}$ open cluster)
(2) AR Cep ( $7.0_{m}$ variable star)


RA: $315.45^{\circ} \mid 21 \mathrm{~h} 1.79^{\prime}-$ DEC: $68.2^{\circ} \mid 68^{\circ} 12^{\prime}$
C 4 (Iris Nebula, NGC 7023) is a magnitude 6.8 bright nebula. Angular size is $18 \times 18^{\prime}$.
3.3${ }^{\circ}$ SW from mag.3.32 Alphirk.

The beautiful Iris Nebula, one of the brighter reflection nebula. Surrounded by dark dust clouds. This nebula shares the same finder circle as the variable T Cep, which lies one degree to the north east.

Also visible:
(1) T Cep (5.2m variable star)


RA: $3.25^{\circ} \mid$ 0h $13.0^{\prime}-$ DEC: $72.53^{\circ} \mid 72^{\circ} 32^{\prime}$
C2 (Bow Tie Nebula, NGC 40) is a magnitude 11.6
planetary nebula. Angular size is $0.6^{\prime}$.
Slightly over halfway from Polaris to Caph.

O:
A relatively faint planetary nebula. The central white dwarf has a surface temperature of approximately $50,000^{\circ} \mathrm{C}$.


RA: $184.18^{\circ} \mid 12 \mathrm{~h} 16.7^{\prime}-$ DEC: $69.47^{\circ} \mid 69^{\circ} 28^{\prime}$
C3 (NGC 4236) is a magnitude 9.7 barred spiral galaxy. Angular size is $22 \times 7$ '.

Just less than halfway from Dubhe to Kochab.

A faint, obliquely oriented barred spiral galaxy. This galaxy is a neglected member of the M81 Group.

Also visible:
(1) NGC 4125 ( $10.65_{m}$ elliptical galaxy)
(2) RY $\operatorname{Dra}$ ( $6.0_{m}$ carbon star)


RA: $148.9^{\circ} \mid 9 \mathrm{~h} 55.6^{\prime}-$ DEC: $69.07^{\circ} \mid 69^{\circ} 4^{\prime}$
M81 (Bode's Galaxy, NGC 3031) is a magnitude 6.9
spiral galaxy. Angular size is $21 \times 10^{\prime}$.
Draw a line from Phecda through Dubhe, and extend it an equal length further. This brings you very close to M81 and M82.


This iconic spiral galaxy forms a delightful pair at lower magnifications with the nearby M82. The core contains a supermassive black hole with 70 million solar masses. M81 is the largest galaxy of the eponymous M81 Group, and at 11.7 million light years is one of the nearest large galaxies.

Also visible:
(1) M82 (8.4m irregular galaxy)
(2) NGC 3077 (10.61m irregular galaxy)
(3) IC 2574 (10.8 $8_{m}$ Magellanic barred spiral galaxy)
(4) R UMa ( $6.5_{m}$ variable star)
(5) VY UMa ( $6.0_{m}$ carbon star)


RA: $11.1^{\circ}\left|0 \mathrm{~h} 44.4^{\prime}-\mathrm{DEC}: 85.33^{\circ}\right| 85^{\circ} 20^{\prime}$

C1 (Polarissima Cluster, NGC 188) is a magnitude 8.1 open cluster. Angular size is 15 '.

Five degrees from Polaris, in the direction of the W of Cassiopeia.

A sparse, loose grouping of middle-aged stars. With an age of nearly 7 billion years, this is an unusually old open cluster. Open clusters normally disperse after only a few hundred million years.

Also visible:
(1) U Cep ( $6.75_{m}$ variable star)
(2) AR Cep (7.0 $0_{m}$ variable star)


RA: $342.89^{\circ} \mid 22 \mathrm{~h} 51.56^{\prime}-$ DEC: $85.05^{\circ} \mid 85^{\circ} 3^{\prime}$

AR Cep (HD 217158) is a magnitude 7.0 variable star.
Magnitude ranges from 7.9 to 7.0 ( $\Delta$ mag. 0.9 ).
Travel from Kochab to Polaris, then turn left and travel a short distance roughly equal to the distance from Kochab to Pherkad.


An orange-yellow variable in a busy field with many points of reference for its narrow 1-magnitude range. The small Polarissima Cluster (Caldwell 1) is on the northeastern edge of the finder circle.


Also visible:
(1) C 1 ( $8.1_{m}$ open cluster)
(2) U Cep ( $6.75_{m}$ variable star)


RA: $56.7^{\circ}\left|3 \mathrm{~h} 46.8^{\prime}-\mathrm{DEC}: 68.1^{\circ}\right| 68^{\circ} 6^{\prime}$

C5 (Maffei 1 Group, IC342) is a magnitude 9.2 spiral galaxy. Angular size is $21 \times 21^{\prime}$.

Draw a line from Shedar to Segin in Cassiopeia, then double its length.

A bright face-on spiral galaxy easily lost in a busy star field. This large and bright spiral was initially thought to be part of our Local Group of galaxies, but is now assigned to the neighboring Maffei Group.


## T Cep

RA: $317.38^{\circ} \mid 21 \mathrm{~h} 9.52^{\prime}-$ DEC: $68.49^{\circ} \mid 68^{\circ} 29^{\prime}$

T Cep (HD 202012) is a magnitude 5.2 variable star. Magnitude ranges from 11.3 to 5.2 ( $\Delta$ mag. 6.1 ) with a period of 388 d .
2.6 ${ }^{\circ}$ SW from mag.3.32 Alphirk.


T Cep can appear somewhat yellow-orange at its peak brightness.

Also visible:
(1) C 4 ( $6.8_{m}$ bright nebula)

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## October: $45^{\circ}$ North (1)



## October: $45^{\circ}$ North (2)




RA: $22.38^{\circ} \mid$ 1h $29.5^{\prime}-$ DEC: $63.3^{\circ} \mid 63^{\circ} 18^{\prime}$
C8 (NGC 559) is a magnitude 9.5 open cluster.
Angular size is $7^{\prime}$.
2.8 ${ }^{\circ} \mathrm{W}$ from mag.3.44 Segin.


A small and faint open cluster.

Also visible:
(1) C 10 (7.1m open cluster)
(2) M103 ( $7.4_{m}$ open cluster)
(3) NGC 654 ( $6.5_{m}$ open cluster)
(4) W Cas ( $7.8_{m}$ variable star)


RA: $351.05^{\circ} \mid 23 \mathrm{~h} 24.2^{\prime}-$ DEC: $61.58^{\circ} \mid 61^{\circ} 35^{\prime}$


M52 (NGC 7654) is a magnitude 7.3 open cluster.
Angular size is 13 '.
Draw a line from Shedar is Cassiopeia to Caph, and extend it an equal length.

A large open cluster (1,200 solar masses), M52 is 4,600 light-years from Earth. The southern half of this cluster is partially obscured by dust.

Also visible:
(1) C9 (7.7 ${ }_{m}$ bright nebula)
(2) C 11 ( $7.0_{m}$ bright nebula)
(3) NGC 7789 ( $6.7_{m}$ open cluster)
(4) WZ Cas (6.3 $3_{m}$ variable star)


RA: $7.48^{\circ} \mid$ Oh $29.9^{\prime}-$ DEC: $60.23^{\circ} \mid 60^{\circ} 14^{\prime}$
NGC 129 is a magnitude 6.5 open cluster. Angular size is $21^{\prime}$.

Midway between Caph and Navi.


A large, scattered open cluster, its age is estimated at 76 million years.

Also visible:
(1) NGC 7789 ( $6.7_{m}$ open cluster)
(2) W Cas ( $7.8_{m}$ variable star)
(3) WZ Cas ( $6.3_{m}$ variable star)


## R Cas

RA: $359.6^{\circ} \mid 23 \mathrm{~h} 58.41^{\prime}-$ DEC: $51.39^{\circ} \mid 51^{\circ} 23^{\prime}$

R Cas (HD 224490) is a magnitude 4.7 variable star.
Magnitude ranges from 13.5 to 4.7 ( $\Delta$ mag. 8.8) with a period of 430 d .

Forms the pointy end of long isosceles triangle with Caph and Shedar, twice as long as it is wide.

This light red star is about 610 light-years from Earth. R Cas has an exceedingly faint companion (magnitude 13.2) separated by 22 ", position angle $266^{\circ}$.

Also visible:
(1) NGC 7686 ( $5.6_{m}$ open cluster)
(2) ST Cas $\left(9.0_{m}\right.$ carbon star)


RA: $8.3^{\circ} \mid$ Oh $33.2^{\prime}-$ DEC: $48.5^{\circ} \mid 48^{\circ} 30^{\prime}$

C17 (NGC 147) is a magnitude 9.3 elliptical galaxy.
Angular size is $13 \times 8$.
$5.4^{\circ} \mathrm{S}$ from mag. 3.72 zet Cas.


A diffuse and therefore difficult target, C 17 is a satellite galaxy of M31 (the Andromeda Galaxy). As such C17 is a member of the Local Group of which the Milky way is a member. C17 seems to have ceased forming new stars three billion years ago.

Also visible:
(1) C 18 (9.2m elliptical galaxy)
(2) VX And ( $8.0_{m}$ carbon star)
(3) ST Cas $\left(9.0_{m}\right.$ carbon star)


RA: $351.48^{\circ} \mid 23 \mathrm{~h} 25.9^{\prime}-$ DEC: $42.55^{\circ} \mid 42^{\circ} 33^{\prime}$

C22 (Blue Snowball, NGC 7662) is a magnitude 9.2
planetary nebula. Angular size is $0.6^{\prime}$.
Forms a rough equilateral triangle with Alpheratz and Scheat (the northern two stars of the Square of Pegasus).


This planetary nebula is 2,500 light-years from Earth.


RA: $10.45^{\circ} \mid 0 \mathrm{~h} 41.8^{\prime}-$ DEC: $41.27^{\circ} \mid 41^{\circ} 16^{\prime}$
M31 (Andromeda Galaxy, NGC 224) is a magnitude 3.4 spiral galaxy. Angular size is $178 \times 63$ '.

The deepest V in the W of Cassiopeia (formed by Navi, Shedar and Caph) points to M31. Also, Almach, Mirach and M31 form a right-angle triangle.

The most fabulous galaxy in the northern hemisphere, M31 is our nearest grand spiral galaxy, and the largest member of our Local Group of galaxies. The Milky Way is on course to collide with it in billions of years time. Given the huge extent of the galaxy (six Moons wide), the galaxy is best appreciated in dark skies with a small wide-field telescope or binoculars.

Also visible:
(1) M32 (8.1 $m_{m}$ elliptical galaxy)
(2) M110 ( $8.5_{m}$ elliptical galaxy)
(3) R And ( $5.8_{m}$ variable star)
(4) VX And ( $8.0_{m}$ carbon star)


RA: $6.01^{\circ} \mid$ 0h $24.03^{\prime}-$ DEC: $38.58^{\circ} \mid 38^{\circ} 35^{\prime}$
R And (HD 1967) is a magnitude 5.8 variable star.
Magnitude ranges from 14.9 to 5.8 ( $\Delta$ mag. 9.1) with a period of 409 d .
6.3 ${ }^{\circ} \mathrm{W}$ from mag.3.94 mu. And.

This red star is 790 light-years from Earth. R And is also a double star with a faint magnitude 12.37 companion separated by $83.5^{\prime \prime}$, position angle $142^{\circ}$.

Also visible:
(1) M31 (3.4m spiral galaxy)
(2) M32 (8.1 $1_{m}$ elliptical galaxy)
(3) M110 ( $8.5_{m}$ elliptical galaxy)
(4) AQ And ( $6.9_{m}$ carbon star)


RA: $344.2^{\circ} \mid 22 \mathrm{~h} 56.79^{\prime}-$ DEC: $62.62^{\circ} \mid 62^{\circ} 37^{\prime}$
C9 (Cave Nebula, Sh2-155) is a magnitude 7.7 bright nebula. Angular size is $50 \times 30^{\prime}$.

Draw a line from Shedar (Alpha Cas) to to Caph (Beta Cas) and extend it twice the distance.

A very faint emission nebula, the Cave Nebula is one of the most difficult Caldwell objects to spot visually. The Cave Nebula appears to be in the very earliest stages of collapsing to form a star cluster.

Also visible:
(1) C 11 ( $7.0_{m}$ bright nebula)
(2) M52 ( $7.3_{m}$ open cluster)
(3) NGC 7354 ( $12.5_{m}$ planetary nebula)


RA: $350.18^{\circ}\left|23 \mathrm{~h} 20.7^{\prime}-\mathrm{DEC}: 61.2^{\circ}\right| 61^{\circ} 12^{\prime}$
C11 (Bubble Nebula, NGC 7635) is a magnitude 7.0 bright nebula. Angular size is $15 \times 8^{\prime}$.
6.0 ${ }^{\circ}$ SE from mag.3.68 iot Cep.

A diffuse but fascinating emission nebula, best known as the Bubble Nebula. This delicate nebula is formed by the radiation of a hot, massive star ( 15 solar masses) clearing its immediate neighborhood of hydrogen gas by the force of its radiation pressure. The bubble is about 6 light-years in diameter.

Also visible:
(1) C9 ( $7.7_{m}$ bright nebula)
(2) M52 (7.3m open cluster)


RA: $19.77^{\circ} \mid$ 1h $19.09^{\prime}-$ DEC: $58.33^{\circ} \mid 58^{\circ} 20^{\prime}$
C13 (Owl Cluster, NGC 457) is a magnitude 6.4 open cluster. Angular size is 13'.
2.0 ${ }^{\circ}$ SSW from mag. 2.8 Rukba.

A bright grouping known as the Owl Cluster. The two bright stars forming the eyes of the owl are not members of the cluster.

Also visible:
(1) C 10 (7.1m open cluster)
(2) M103 ( $7.4_{m}$ open cluster)
(3) NGC 654 ( $6.5_{m}$ open cluster)
(4) W Cas ( $7.8_{m}$ variable star)
(5) WW Cas ( $9.1_{m}$ carbon star)


## NGC 7686

RA: $352.55^{\circ} \mid 23 \mathrm{~h} 30.2^{\prime}-$ DEC: $49.13^{\circ} \mid 49^{\circ} 8^{\prime}$

NGC 7686 is a magnitude 5.6 open cluster. Angular size is $15^{\prime}$.

Forms the pointy end of a long triangle with Caph and Shedar, roughly three times as long it is wide.

This bright cluster of stars is an easy binocular target. H L Johnson argued in 1961 that this is not an open cluster as the stars did not appear to share a common age.

Also visible:
(1) R Cas (4.7m variable star)


## C18

RA: $9.75^{\circ} \mid 0 \mathrm{~h} 39.0^{\prime}-$ DEC: $48.33^{\circ} \mid 48^{\circ} 20^{\prime}$
C18 (NGC 185) is a magnitude 9.2 elliptical galaxy.
Angular size is $12 \mathrm{x} 9^{\prime}$.
$5.5^{\circ} \mathrm{S}$ from mag. 3.72 zet Cas.


Like nearby C17, this is is a satellite galaxy of M31 (the Andromeda Galaxy). Being both less diffuse and somewhat brighter than C17, C18 is somewhat less difficuly target in light-polluted skies. C18 contains much younger stars than C17 and has an active nucleus; it may be classified as a Seyfert galaxy and would be the closest galaxy of such type to Earth.

Also visible:
(1) C 17 ( $9.3_{m}$ elliptical galaxy)
(2) VX And ( $8.0_{m}$ carbon star)
(3) ST Cas $\left(9.0_{m}\right.$ carbon star)


## M110

RA: $10.1^{\circ} \mid 0 \mathrm{~h} 40.4^{\prime}-$ DEC: $41.68^{\circ} \mid 41^{\circ} 41^{\prime}$
M110 (NGC 205) is a magnitude 8.5 elliptical galaxy.
Angular size is $17 \times 10^{\prime}$.
Once you have found M31, M110 is visible at the same time, slightly to the north west.

M110 is a dwarf elliptical galaxy orbiting M31. The full classification is "pec dE5": peculiar because it has traces of dust and star formation near its center, and " 5 " because it is $50 \%$ flattened. There is no evidence for a supermassive black hole in M110. Fourteen of the satellite galaxies orbiting M31 follow a similar orbit to M110, suggesting they were originally satellites of M110 that were separated by interactions with the far more massive M31.

Also visible:
(1) M31 (3.4m spiral galaxy)
(2) M32 (8.1 $1_{m}$ elliptical galaxy)
(3) R And ( $5.8_{m}$ variable star)
(4) VX And ( $8.0_{m}$ carbon star)


RA: $10.7^{\circ} \mid 0 \mathrm{~h} 42.8^{\prime}-$ DEC: $40.87^{\circ} \mid 40^{\circ} 52^{\prime}$
M32 (NGC 221) is a magnitude 8.1 elliptical galaxy.
Angular size is $8 \times 66^{\prime}$.
Once you have found M31, M32 is visible at the same time, slightly to the south.

M32 is a compact elliptical (cE) galaxy orbiting M31. This kind of very dense and featureless dwarf galaxy is widely believed to form as a remnant of a larger galaxy that has had it outer parts stripped away. Possibly gas from the outer regions was forced into the core, resulting in a burst of star formation that helps to explain the high density of stars in the inner regions of a cE galaxy.

Also visible:
(1) M31 (3.4m spiral galaxy)
(2) M110 ( $8.5_{m}$ elliptical galaxy)
(3) R And ( $5.8_{m}$ variable star)
(4) VX And ( $8.0_{m}$ carbon star)


RA: $339.28^{\circ} \mid 22 \mathrm{~h} 37.09^{\prime}-$ DEC: $34.42^{\circ} \mid 34^{\circ} 25^{\prime}$
C30 (Deer Lick Group, NGC 7331) is a magnitude 9.5 barred spiral galaxy. Angular size is $11 \times 4{ }^{\prime}$.
4.3${ }^{\circ}$ NNW from mag.3.1 Matar.

This photogenic spiral galaxy has a rich spiral structure with many dark dust clouds, and a prominent and bright yellow core. Dark skies and larger telescopes are needed to see the spiral structure, but the core can be seen under less perfect conditions.

## October: $\mathbf{1 5}^{\circ}$ North




RA: $19.02^{\circ} \mid$ 1h $16.09^{\prime}-$ DEC: $25.8^{\circ} \mid 25^{\circ} 48^{\prime}$
Z Psc (HD 7561) is a magnitude 7.0 carbon star.
Magnitude ranges from 7.9 to $7.0(\Delta$ mag. 0.9$)$ with a period of 144 d .
Halfway between Hamal of Aries and Alpheratz of Pegasus.

This carbon star has a B-V color index of 2.62.
Messier 33, the Triangulum Galaxy is just over one finder circle to the north-east.


RA: $0.83^{\circ} \mid$ 0h $3.3^{\prime}-$ DEC: $16.15^{\circ} \mid 16^{\circ} 9^{\prime}$

C43 (Little Sombrero Galaxy, NGC 7814) is a magnitude 10.5 barred spiral galaxy. Angular size is $6 \times 2$.
2.5 ${ }^{\circ}$ NWW from mag.2.87 Algenib.

An ideal Caldwell object: a small galaxy graced with a bright core and dark dust lane, close to a prominent signpost in the form of bright Algenib, the southeastern star of the Square of Pegasus.


RA: $346.23^{\circ} \mid 23 \mathrm{~h} 4.9^{\prime}-$ DEC: $12.32^{\circ} \mid 12^{\circ} 19^{\prime}$
C44 (NGC 7479) is a magnitude 11.0 barred spiral galaxy. Angular size is $4 \times 3$ '.
2.8 ${ }^{\circ} \mathrm{S}$ from mag.2.57 Marchab.

A face-on galaxy with relatively bright spiral arms, this object benefits greatly from any increase in aperture. This galaxy is a starburst galaxy, with many bright star-forming regions, and two recent supernovae (1990 and 2009).

Also visible:
(1) Palomar 13 ( $13.47_{m}$ globular cluster)
(2) R Peg ( $6.9_{m}$ variable star)


RA: $346.66^{\circ} \mid 23 \mathrm{~h} 6.65^{\prime}-$ DEC: $10.54^{\circ} \mid 10^{\circ} 33^{\prime}$

R Peg (HD 218292) is a magnitude 6.9 variable star.
Magnitude ranges from 13.8 to 6.9 ( $\Delta$ mag. 6.9 ) with a period of 378 d .
4.6 ${ }^{\circ} \mathrm{S}$ from mag.2.57 Marchab.

At its brightest this star can appear yellowish.
Caldwell 44 is $2^{\circ}$ to the NNW.
Also visible:
(1) C44 (11.0 $0_{m}$ barred spiral galaxy)
(2) Palomar 13 ( $13.47_{m}$ globular cluster)


RA: $356.6^{\circ} \mid 23 \mathrm{~h} 46.39^{\prime}-$ DEC: $3.49^{\circ} \mid 3^{\circ} 29^{\prime}$

TX Psc (HD 223075) is a magnitude 4.79 variable star. Magnitude ranges from 5.2 to 4.79 ( $\Delta$ mag. 0.4).

Forms a triangle with Markab and Algenib or Pegasus.

This carbon star has a color index of 2.52 and spectral class of C-N6. It is about 650 light-years from Earth.


RA: $16.2^{\circ} \mid$ 1h $4.79^{\prime}-$ DEC: $2.12^{\circ} \mid 2^{\circ} 7^{\prime}$
C51 (IC 1613) is a magnitude 9.0 irregular galaxy.
Angular size is $12 \times 11^{\prime}$.
Two and a half finder circles east and slightly south of Betelgeuse.

As might be guessed from its discovery date of 1906, this is a diffuse and difficult object, even under dark skies with a larger telescope. Caldwell 51 is a member of our Local Group of Galaxies and is roughly the same distance from us as the famous and vastly larger Andromeda Galaxy (M31).

Also visible:
(1) NGC 488 (11.15m barred spiral galaxy)

## October: - $\mathbf{1 5}^{\circ}$ South



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RA: $11.75^{\circ}\left|0 \mathrm{~h} 47.0^{\prime}-\mathrm{DEC}:-11.88^{\circ}\right|-11^{\circ} 52^{\prime}$

C56 (Skull Nebula, NGC 246) is a magnitude 8.0
planetary nebula. Angular size is $3.8^{\prime}$.
5.5 ${ }^{\circ}$ SWW from mag.3.6 Deneb Algenubi.

This diffuse planetary nebula is more challenging with light pollution, but under good skies it can be viewed with a small telescope.

Also visible:
(1) NGC 157 (11.0 $0_{m}$ barred spiral galaxy)


RA: $0.53^{\circ}\left|0 \mathrm{~h} 2.12^{\prime}-\mathrm{DEC}:-14.68^{\circ}\right|-14^{\circ} 40^{\prime}$

W Cet (HD 224960) is a magnitude 7.1 variable star.
Magnitude ranges from 14.8 to 7.1 ( $\Delta$ mag. 7.7) with a period of 351 d .
7.2 ${ }^{\circ}$ SW from mag.3.75 Deneb Kaitos Shemali.
©
Also visible:
(1) WLM (11.03 ${ }_{m}$ irregular galaxy)
(2) R Aqr ( $5.8_{m}$ variable star)


RA: $355.96^{\circ} \mid 23 \mathrm{~h} 43.82^{\prime}-$ DEC: $-15.28^{\circ} \mid-15^{\circ} 16^{\prime}$
R Aqr (HD 222800) is a magnitude 5.8 variable star. Magnitude ranges from 12.4 to 5.8 ( $\Delta$ mag. 6.6) with a period of 387 d .
$10.0^{\circ} \mathrm{NE}$ from mag. 3.8 c 02 Aqr .


This moderately red variable is around 1045 lightyears from Earth.

Also visible:
(1) WLM (11.03 ${ }_{m}$ irregular galaxy)
(2) W Cet ( $7.1_{m}$ variable star)


## T Cet

RA: $5.44^{\circ} \mid 0 \mathrm{~h} 21.77^{\prime}-$ DEC: $-20.06^{\circ} \mid-20^{\circ} 2^{\prime}$

T Cet (HD 1760) is a magnitude 5.0 variable star. Magnitude ranges from 6.9 to $5.0(\Delta$ mag. 1.9$)$ with a period of 159 d .
5.5 SWW from mag.2.24 Diphda.

This moderately red variable has a less extreme range in brightness compared to many others in this book, but there are good stars in the finder view to make a comparison. Mag. 6.95 SAO 166174 is just over a degree to the southwest, and marks the minimum brightness for T Cet. 2 degrees to the northwest is 7 Cet, which at approximately mag. 4.4 is just a bit brighter than T Cet at its brightest. SAO 147275 is 2 degrees northeast, and marks the midpoint at mag. 6.2.


RA: $11.78^{\circ} \mid 0 \mathrm{~h} 47.1^{\prime}-$ DEC: $-20.77^{\circ} \mid-20^{\circ} 45^{\prime}$
C62 (Needle's Eye Galaxy, NGC 247) is a magnitude 8.9 spiral galaxy. Angular size is $20 \times 7$ '.
2.9 ${ }^{\circ}$ SSE from mag.2.24 Diphda.

The Needle's Eye galaxy is so named because it is almost edge-on, with a dark void in its disk. While this void is filled with older stars, it lacks gas clouds and bight young stars; it is believed the void is the result of an interaction with another galaxy. This dwarf spiral galaxy is about 11 million light-years from us in the nearby Sculptor Group of galaxies.

Also visible:
(1) C65 (7.1 $1_{m}$ spiral galaxy)


C65 (Sculptor Galaxy, NGC 253) is a magnitude 7.1 spiral galaxy. Angular size is $25 \times 7$ '.
$7.3^{\circ} \mathrm{S}$ from mag.2.24 Diphda.

Discovered in 1783 by Caroline Herschel, this large and luminous galaxy is a gorgeous sight in any telescope.

Also visible:
(1) C62 (8.9m spiral galaxy)
(2) NGC 288 ( $8.09_{m}$ globular cluster)

October: - $\mathbf{4 5}^{\circ}$ South



RA: $3.84^{\circ} \mid 0$ h $15.37^{\prime}-$ DEC: $-32.05^{\circ} \mid-32^{\circ} 2^{\prime}$
$\mathrm{S} \mathrm{Scl}(\mathrm{HD} 001115)$ is a magnitude 5.5 variable star.
Magnitude ranges from 13.6 to 5.5 ( $\Delta$ mag. 8.1) with a period of 363 d .
10.4 ${ }^{\circ}$ NNW from mag.2.44 Ankaa.


This yellowish variable star fades out of sight (for most telescopes) on an almost exactly annual basis. It is 3,507 light-years from Earth.

Also visible:
(1) NGC 7793 ( $9.63_{m}$ spiral galaxy)
(2) NGC 134 (11.23 ${ }_{m}$ barred spiral galaxy)


RA: $21.75^{\circ} \mid$ 1h $27.0^{\prime}-$ DEC: $-32.5^{\circ} \mid-32^{\circ} 29^{\prime}$
R Scl (HD 8879) is a magnitude 6.1 carbon star.
Magnitude ranges from 8.8 to 6.1 ( $\Delta$ mag. 2.7) with a period of 363 d .
$10.8^{\circ} \mathrm{N}$ from mag.3.4 gam Phe.

This carbon star has a B-V color index of 2.46 and spectral type C-N5+. It is approximately 1436 lightyears from Earth.

Also visible:
(1) NGC 613 (10.73 $3_{m}$ barred spiral galaxy)


RA: $13.73^{\circ} \mid 0 \mathrm{~h} 54.9^{\prime}-$ DEC: $-37.68^{\circ} \mid-37^{\circ} 40^{\prime}$
C70 (Southern Pinwheel Galaxy, NGC 300) is a
magnitude 8.1 spiral galaxy. Angular size is 20x13'.
7.1 ${ }^{\circ}$ NE from mag.2.44 Ankaa.


This large and bright galaxy is only six million lightyears from us and is a member of the Sculptor Group.

Also visible:
(1) Sculptor (10.5m irregular galaxy)


RA: $3.73^{\circ} \mid 0$ h $14.9^{\prime}-$ DEC: $-39.18^{\circ} \mid-39^{\circ} 10^{\prime}$
C72 (String of Pearls, NGC 55) is a magnitude 8.2 barred spiral galaxy. Angular size is $32 \times 6{ }^{\prime}$.
3.7 ${ }^{\circ}$ NW from mag.2.44 Ankaa.

This is another bright member of the nearby Sculptor Group of galaxies and is Magellanic-type galaxy, meaning it is intermediate between a small irregular galaxy and dwarf spiral galaxy.

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## November: $\mathbf{4 5}^{\circ}$ North (1)



## November: $\mathbf{4 5}^{\circ}$ North (2)




RA: $61.93^{\circ}\left|4 \mathrm{~h} 7.69^{\prime}-\mathrm{DEC}: 62.33^{\circ}\right| 62^{\circ} 20^{\prime}$
NGC 1502 is a magnitude 5.7 open cluster. Angular size is $8^{\prime}$.
$11.7^{\circ} \mathrm{NE}$ from mag.3.93 Miram.

This young open cluster is estimated at 5-15 million years age, and contains about 60 stars of widely varying brightness. It is at one end of the unrelated Kemble's Cascade asterism, a three degree long straight line of over 205 th to 10th magnitude stars.

Also visible:
(1) NGC 1501 ( $12.5_{m}$ planetary nebula)


RA: $26.5^{\circ} \mid$ 1h $46.0^{\prime}-$ DEC: $61.25^{\circ} \mid 61^{\circ} 15^{\prime}$

C 10 (NGC 663) is a magnitude 7.1 open cluster.
Angular size is 16 .
2.6 ${ }^{\circ}$ SSW from mag.3.44 Segin.

A rich open cluster, possibly bright enough to spotted with the naked eye from a dark site. The cluster is roughly 22 million years old, with many of its larger stars now in their later (and brighter) stages of life. The cluster is enhanced by an unrelated, more distant dust cloud behind it which clears the view of competing background stars.

Also visible:
(1) C13 ( $6.4_{m}$ open cluster)
(2) C14 (4.3m open cluster)
(3) M103 (7.4m open cluster)
(4) NGC 654 ( $6.5_{m}$ open cluster)
(5) WW Cas (9.1m carbon star)


RA: $35.0^{\circ} \mid 2 \mathrm{~h} 20.0^{\prime}-$ DEC: $57.13^{\circ} \mid 57^{\circ} 8^{\prime}$
C14 (Double Cluster, NGC 869/884) is a magnitude 4.3 open cluster. Angular size is $60 \times 30^{\prime}$.
4.4 ${ }^{\circ}$ NWW from mag.3.93 Miram.


A gorgeous pair of brilliant open clusters, both bright enough to be spotted by the naked eye under dark skies. In addition to many hot blue members, the Double Cluster also hosts a number of older clearly red giants.

Also visible:
(1) C10 (7.1 $1_{m}$ open cluster)
(2) NGC 1027 ( $6.7_{m}$ open cluster)


RA: $63.85^{\circ} \mid 4 \mathrm{~h} 15.4^{\prime}-$ DEC: $51.23^{\circ} \mid 51^{\circ} 14^{\prime}$

NGC 1528 is a magnitude 6.4 open cluster. Angular size is $24^{\prime}$.

Draw a line from Shedar through Ruchbah in
Cassiopeia, and extend it halfway to Capella.
At 110 million years old, this cluster of 165 stars can be seen with binoculars. The brightest star in the cluster is magnitude 8.7.

Also visible:
(1) NGC 1444 ( $6.6_{m}$ open cluster)
(2) NGC 1545 ( $6.2_{m}$ open cluster)


RA: $40.5^{\circ} \mid 2 \mathrm{~h} 42.0^{\prime}-$ DEC: $42.78^{\circ} \mid 42^{\circ} 47^{\prime}$
M34 (NGC 1039) is a magnitude 5.5 open cluster.
Angular size is $35^{\prime}$.
5.2 ${ }^{\circ}$ NWW from mag.2.12 Algol.


Covering a patch of sky the size of the Moon, the small stars of this cluster can be resolved with binoculars in suburban skies. It is estimated to be at least 200 million years old, so it also contains 19 white dwarfs which are the remnants of the larger, shorter-lived members of this cluster.

Also visible:
(1) C23 (9.9 barred spiral galaxy)
(2) NGC 1023 (10.35m lenticular galaxy)


RA: $49.95^{\circ} \mid 3 \mathrm{~h} 19.8^{\prime}-$ DEC: $41.52^{\circ} \mid 41^{\circ} 31^{\prime}$
C24 (Perseus A, NGC 1275) is a magnitude 11.6 seyfert galaxy. Angular size is $2^{\prime}$.
2.2 ${ }^{\circ}$ NEE from mag.2.12 Algol.

A faint and very difficult Caldwell object, this object consists of two interacting galaxies - a massive elliptical galaxy, and a highly disrupted spiral galaxy. In deep photographs, C24 is surrounded by a mysterious web of hydrogen filaments that are much cooler than the intergalactic medium. It is uncertain why they are cooler and also why they have not collapsed to form new stars.

Also visible:
(1) NGC 1342 ( $6.7_{m}$ open cluster)
(2) Y Per (8.1m carbon star)


RA: $39.26^{\circ}$ | 2h $37.03^{\prime}-$ DEC: $34.26^{\circ} \mid 34^{\circ} 16^{\prime}$
R Tri (HD 016210) is a magnitude 5.4 variable star.
Magnitude ranges from 12.6 to 5.4 ( $\Delta$ mag. 7.2 ) with a period of 267 d .
Halfway between Algol and Hamal.

This orange-red variable is 3,000 light-years from Earth. There are four faint galaxies in the finder circle. NGC 959 is 2 degrees to the northwest (mag 12.4); NGC 925 is 2.5 degrees to the west (mag. 10.1); NGC 969 is less than 2 degrees south (mag. 12.4); NGC 1060 is 2.5 degrees to the southeast (mag. 11.8).

Also visible:
(1) NGC 925 (10.69 $9_{m}$ barred spiral galaxy)


RA: $58.85^{\circ}\left|3 \mathrm{~h} 55.38^{\prime}-\mathrm{DEC}: 31.05^{\circ}\right| 31^{\circ} 3{ }^{\prime}$

X Per (HD 024534) is a magnitude 6.03 variable star.
Magnitude ranges from 7.0 to 6.03 ( $\Delta$ mag. 1.0).
$0.8^{\circ}$ SSE from mag. 2.91 zet Per.

This white variable star is roughly 2,651 light-years from Earth. It is also a double star, with a magnitude 12.3 companion separated by 23.4 ", position angle $13^{\circ}$.

Also visible:
(1) NGC 1514 ( $9.5_{m}$ planetary nebula)
(2) IC 351 ( $12.0_{m}$ planetary nebula)
(3) IC 2003 (12.0m planetary nebula)


## NGC 654

RA: $26.03^{\circ} \mid$ 1h $44.09^{\prime}-$ DEC: $61.88^{\circ} \mid 61^{\circ} 53^{\prime}$
NGC 654 is a magnitude 6.5 open cluster. Angular size is $5^{\prime}$.
2.1 ${ }^{\circ}$ SW from mag.3.44 Segin.

This cluster is possibly as young as 15 million years, but there is some uncertainty as star formation may have spanned 20 million years. The cluster has about 80 members and is dominated by HD 10494, a yellow supergiant with class F5Ia.

Also visible:
(1) C8 (9.5m open cluster)
(2) C 10 (7.1 $1_{m}$ open cluster)
(3) C13 ( $6.4_{m}$ open cluster)
(4) M103 ( $7.4_{m}$ open cluster)
(5) WW Cas (9.1 $1_{m}$ carbon star)


RA: $23.3^{\circ} \mid$ 1h $33.2^{\prime}-$ DEC: $60.7^{\circ} \mid 60^{\circ} 42^{\prime}$

M103 (NGC 581) is a magnitude 7.4 open cluster.
Angular size is $6^{\prime}$.
$1.0^{\circ}$ NEE from mag.2.8 Rukba.

This open cluster is quite distant compared to other open clusters in the Messier catalog, being roughly 9,000 light-years from Earth, and is composed of about 40 faint stars. The view of the cluster is dominated by Struve 131, which is a foreground star and not a member of the cluster. Struve 131 is a double star (mag. 7.3 and 9.9, separation 13.9", position angle $143^{\circ}$ ).

Also visible:
(1) C8 (9.5m open cluster)
(2) C 10 (7.1 $\mathrm{l}_{m}$ open cluster)
(3) C13 ( $6.4_{m}$ open cluster)
(4) NGC 654 ( $6.5_{m}$ open cluster)
(5) WW Cas (9.1m carbon star)


RA: $25.6^{\circ} \mid$ 1h $42.4^{\prime}-$ DEC: $51.57^{\circ} \mid 51^{\circ} 34^{\prime}$

M76 (Little Dumbbell Nebula, NGC 650) is a magnitude 10.1 planetary nebula. Angular size is $3 \times 2^{\prime}$.

Midway between Almach and Ruchbah.

This planetary nebula is the same angular size as four Jupiters. It is graced with two NGC numbers, NGC 650 and NGC 651 , as was originally thought to be two distinct nebulae (the double-lobed nature of this nebula is only apparent in telescopes over around 150 mm aperture). This object is quite difficult in lightpolluted skies.


RA: $65.22^{\circ} \mid 4 \mathrm{~h} 20.89^{\prime}-$ DEC: $50.25^{\circ} \mid 50^{\circ} 15^{\prime}$

NGC 1545 is a magnitude 6.2 open cluster. Angular size is $18^{\prime}$.

Slightly north of a point halfway between Capella and Mirfak.

Only 1.5 degrees from southeast from the similar NGC 1528, this open cluster is around 2,320 lightyears from Earth and around 280 million years old.

Also visible:
(1) NGC 1528 ( $6.4_{m}$ open cluster)


RA: $35.65^{\circ} \mid 2 \mathrm{~h} 22.59^{\prime}-$ DEC: $42.35^{\circ} \mid 42^{\circ} 21^{\prime}$

C23 (NGC 891) is a magnitude 9.9 barred spiral galaxy. Angular size is $14 \times 3$ '.
3.4 ${ }^{\circ}$ E from mag.2.28 Almach.


This edge-on spiral galaxy in challenging under lightpolluted skies, but rewards those who find it with a prominent central dust lane. Studies of the extended halo of this galaxy (the halo is the diffuse region containing the globular clusters of a galaxy) showed that the halo material was most likely ejected from the galaxy, rather than being pulled in from intergalactic space.

Also visible:
(1) M34 ( $5.5_{m}$ open cluster)
(2) NGC 1023 (10.35m lenticular galaxy)


RA: $29.45^{\circ} \mid$ 1h $57.8^{\prime}-$ DEC: $37.68^{\circ} \mid 37^{\circ} 41^{\prime}$

C28 (NGC 752) is a magnitude 5.7 open cluster.
Angular size is 50 '.
$3.5^{\circ} \mathrm{NW}$ from mag. 3.08 bet Tri.

A relatively bright cluster discovered by Caroline Herschel in 1783 . Under very good skies this large, loose cluster is visible by the naked eye, but is an easy binocular target as it has a number of stars approaching 9th magnitude in brightness.


RA: $23.48^{\circ} \mid$ 1h $33.9^{\prime}-$ DEC: $30.65^{\circ} \mid 30^{\circ} 39^{\prime}$
M33 (Triangulum Galaxy, NGC 598) is a magnitude 5.7 spiral galaxy. Angular size is $73 \times 45$ '.

Travel from the lovely double star Almach to Mirach and turn right; continue half the distance again.
":
A large object spreading over a degree of sky, M33 (the Triangulum Galaxy) is the third largest galaxy in the Local Group, after M31 and the Milky Way. M33 is a challenging naked-eye target under dark skies, a challenging telescopic target in suburban skies, and virtually impossible under city skies. It has giant emission nebulae, including NGC 604 (spanning 1500 light-years) and three smaller regions (NGC 588, NGC 592, and NGC 595).

## November: $\mathbf{1 5}^{\circ}$ North




RA: $34.03^{\circ} \mid 2 \mathrm{~h} 16.11^{\prime}-$ DEC: $25.06^{\circ} \mid 25^{\circ} 3^{\prime}$
R Ari (HD 13913) is a magnitude 7.4 variable star.
Magnitude ranges from 13.7 to 7.4 ( $\Delta$ mag. 6.3) with a period of 187 d .
2.5 ${ }^{\circ}$ NE from mag.2.23 Hamal.

At its brightest, this pulsating variable star forms a binocular double with 21 Ari (magnitude 6.4, separation 6').


M45 (Pleiades, Mel 22) is a magnitude 1.6 open cluster. Angular size is $110^{\prime}$.
$0.1^{\circ} \mathrm{W}$ from mag.2.96 Alcyone.

One of the nearest clusters to Earth, the Pleiades cuts through even the worst light pollution. Galileo sketched the Pleiades in 1610 , noting 36 stars. In long exposures the Pleiades are veiled in dust; the cluster is drifting through a dusty region, and the dust is not in fact part of the cluster. It is possible to glimpse this dusty glow with a small telescope.


RA: $66.75^{\circ} \mid 4 \mathrm{~h} 27.0^{\prime}-$ DEC: $16.0^{\circ} \mid 16^{\circ} 0^{\prime}$
C41 (Hyades, Mel 25) is a magnitude 1.0 open cluster.
Angular size is 330 '.
0.4 ${ }^{\circ}$ NWW from mag. 3.62 the 02 Tau.

The Hyades are an obvious feature of the northern sky, clustering around Aldebaran and easily seen with the naked eye even in light polluted conditions. The Hyades are the nearest open cluster to Earth. Despite appearances, Aldebaran is not a member of this cluster. This target is best viewed with binoculars or a small telescope.

Also visible:
(1) NGC 1647 ( $6.4_{m}$ open cluster)
(2) T Tau ( $9.3_{m}$ variable star)


RA: $24.18^{\circ} \mid$ 1h $36.7^{\prime}-$ DEC: $15.78^{\circ} \mid 15^{\circ} 47^{\prime}$

M74 (NGC 628) is a magnitude 9.4 spiral galaxy.
Angular size is $10^{\prime}$.
1.3 ${ }^{\circ}$ NEE from mag.3.72 Al'farg.

A face-on grand design spiral galaxy, this object has a very low surface brightness and even larger telescopes struggle to reveal its two spiral arms, which are however clearly defined in photographs.


RA: $47.76^{\circ} \mid 3 \mathrm{~h} 11.05^{\prime}-$ DEC: $14.8^{\circ} \mid 14^{\circ} 48^{\prime}$
U Ari (HD 19737) is a magnitude 7.2 variable star.
Magnitude ranges from 15.2 to 7.2 ( $\Delta$ mag. 8.0) with a period of 371 d .

Forms a parallel zig-zag with Aldebaran, the Pleiades and Hamal.

This Mira-type variable is over 3,300 light-years from Earth.


RA: $33.75^{\circ} \mid 2 \mathrm{~h} 15.0^{\prime}-$ DEC: $12.2^{\circ} \mid 12^{\circ} 12^{\prime}$
V Ari (HD 13826) is a magnitude 8.0 carbon star.
Magnitude ranges from 8.6 to $8.0(\Delta$ mag. 0.6$)$ with a period of 75d.
9.8 ${ }^{\circ}$ SSE from mag.2.72 Sheratan.

With a moderate B-V color index of 2.11 , this isn't one of the reddest carbon stars. It is an innately bright object with an absolute magnitude of -2.07 , dimmed by its great distance of roughly 4235 light-years.


RA: $67.08^{\circ} \mid 4$ h $28.3^{\prime}-$ DEC: $10.16^{\circ} \mid 10^{\circ} 10^{\prime}$
R Tau (HD 028309) is a magnitude 7.6 variable star.
Magnitude ranges from 15.8 to 7.6 ( $\Delta$ mag. 8.2) with a period of 321 d .
$5.7^{\circ} \mathrm{S}$ from mag. 3.62 the02 Tau.


RA: $22.66^{\circ} \mid$ 1h $30.64^{\prime}-$ DEC: $2.88^{\circ} \mid 2^{\circ} 53^{\prime}$

R Psc (HD 009203) is a magnitude 7.0 variable star.
Magnitude ranges from 14.8 to 7.0 ( $\Delta$ mag. 7.8) with a period of 345 d .
$11.1^{\circ} \mathrm{N}$ from mag. 3.83 the Cet.

This Mira-type variable is 1,853 light-years from Earth.

Also visible:
(1) NGC 488 (11.15m barred spiral galaxy)


RA: $40.67^{\circ} \mid 2 \mathrm{~h} 42.69^{\prime}-$ DEC: $0.03^{\circ} \mid 0^{\circ} 2^{\prime}$
M77 (NGC 1068) is a magnitude 8.9 spiral galaxy.
Angular size is $7 \times 6$.
$3.2^{\circ} \mathrm{S}$ from mag.3.58 Kaffaljidhma.
"
Messier 77 is one of the closest type 2 Seyfert galaxies, and is accordingly well-studied. Type 2 Seyfert galaxies have an active core that is thought to be veiled from sight by intervening clouds of dust.

Also visible:
(1) NGC 936 (11.12 $2_{m}$ lenticular galaxy)

## November: - $\mathbf{1 5}^{\circ}$ South



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RA: $34.84^{\circ} \mid 2$ h $19.34^{\prime}-$ DEC: $-2.98^{\circ} \mid-2^{\circ} 58^{\prime}$
o Cet (Mira, HD 14386) is a magnitude 2.0 variable star. Magnitude ranges from 10.1 to 2.0 ( $\Delta$ mag. 8.1) with a period of 332 d .

At it brightest, a comfortable naked-eye target.

The first periodic variable star to be discovered,
Mira's changeful nature was first observed by David Fabricius in 1596 . Although it dims by a factor of around 600 in the visible spectrum, a lot of this energy is nevertheless radiated in the infrared spectrum instead. In infrared, Mira only varies by 2 magnitudes. It spends one third of its cycle brightening, and twothirds fading. Mira is very close, only 200-400 lightyears from Earth.

Also visible:
(1) NGC 936 (11.12 $2_{m}$ lenticular galaxy)


U Cet (HD 15971) is a magnitude 6.8 variable star. Magnitude ranges from 13.4 to 6.8 ( $\Delta$ mag. 6.6) with a period of 235 d .
10.7 ${ }^{\circ}$ SEE from mag.3.92 Baten Kaitos.

This red giant is 4,000 light years from Earth.


RA: $63.55^{\circ} \mid 4 \mathrm{~h} 14.19^{\prime}-$ DEC: $-12.73^{\circ} \mid-12^{\circ} 43^{\prime}$
NGC 1535 (Cleopatra's Eye Nebula) is a magnitude 9.5 planetary nebula. Angular size is $18{ }^{\prime \prime}$. The central star is magnitude 12.2.
4.0 ${ }^{\circ}$ E from mag.3.19 Zaurak.


This faint planetary nebula is relatively remote at 5,500 to 7,500 light-years distance, five times more distant than the first planetary to be discovered, Messier 27.


RA: $58.81^{\circ} \mid 3 \mathrm{~h} 55.23^{\prime}-$ DEC: $-24.03^{\circ} \mid-24^{\circ} 1^{\prime}$

T Eri (HD 024754) is a magnitude 7.2 variable star.
Magnitude ranges from 13.2 to 7.2 ( $\Delta$ mag. 6.0) with a period of 252 d .
$8.5^{\circ}$ SEE from mag. 3.95 tau04 Eri.

5,000 to 7,000 light years from Earth, this red giant has spectral class M5/6IIIe.

Also visible:
(1) NGC 1395 (10.55m elliptical galaxy)
(2) NGC 1398 (10.57m barred spiral galaxy)


## NGC 1360

RA: $53.33^{\circ}\left|3 \mathrm{~h} 33.3^{\prime}-\mathrm{DEC}:-25.85^{\circ}\right|-25^{\circ} 50^{\prime}$

NGC 1360 (Robin's Egg Nebula) is a magnitude 9.5 planetary nebula. Angular size is 380 ". The central star is magnitude 11.4.
$5.1^{\circ}$ SE from mag.3.95 tau04 Eri.


This planetary nebula has an interesting pair of stars in its center: a white dwarf (as is usual for a planetary nebula) and a subdwarf O-type star, about half the mass of the Sun but somewhat brighter. The subdwarf O-type star is composed of a dense carbon-oxygen core surrounded by a helium burning shell. The atmosphere of the star is highly enriched in helium (or depleted in hydrogen).

Also visible:
(1) NGC 1395 (10.55m elliptical galaxy)
(2) NGC 1398 (10.57m barred spiral galaxy)


RA: $37.31^{\circ} \mid 2 \mathrm{~h} 29.25^{\prime}-$ DEC: $-26.1^{\circ} \mid-26^{\circ} 5^{\prime}$
R For (HIP 011582) is a magnitude 7.5 variable star. Magnitude ranges from 13.0 to 7.5 ( $\Delta$ mag. 5.5) with a period of 389 d .
$9.9^{\circ}$ NWW from mag. 3.95 alf For.

Two degrees to the northwest is the faint galaxy NGC 922 (magnitude 12.21). The star has the spectral class of $\mathrm{C} 4,3 \mathrm{e}$ indicating it is a carbon star, but it has a low $\mathrm{B}-\mathrm{V}$ color index of 1.61 , so it is only moderately red. The star is roughly 2,000 light years from Earth.
Also visible:
(1) C67 (9.2 $2_{m}$ barred spiral galaxy)

## November: - $\mathbf{4 5}^{\circ}$ South




RA: $41.58^{\circ} \mid 2 h 46.3^{\prime}-$ DEC: $-30.28^{\circ} \mid-30^{\circ} 16^{\prime}$
C67 (NGC 1097) is a magnitude 9.2 barred spiral galaxy. Angular size is $9 \times 6{ }^{\prime}$.
5.7 ${ }^{\circ}$ SWW from mag. 3.95 alf For.

คั
While the galaxy's core is bright and easy to observe, the spiral arms are much fainter.

Also visible:
(1) Fornax Dwarf (9.3 $3_{m}$ irregular galaxy)
(2) R For ( $7.5_{m}$ variable star)


RA: $39.99^{\circ} \mid 2 \mathrm{~h} 39.97^{\prime}-$ DEC: $-34.45^{\circ} \mid-34^{\circ} 26^{\prime}$

Fornax Dwarf is a magnitude 9.3 irregular galaxy. Angular size is 17 '.
6.8 ${ }^{\circ}$ NNW from mag.3.42 Acamar.

Only discovered in 1938 in South Africa from photographic plates, this inconsequential satellite of the Milky Way is virtually invisible. It possesses six globular clusters, and one of these, NGC 1049, was discovered by John Herschel over a century before the parent galaxy was detected.

Also visible:
(1) C67 (9.2 $2_{m}$ barred spiral galaxy)


RA: $50.68^{\circ} \mid 3 \mathrm{~h} 22.69^{\prime}-$ DEC: $-37.2^{\circ} \mid-37^{\circ} 11^{\prime}$
NGC 1316 is a magnitude 9.42 lenticular galaxy.
Angular size is 12 .
5.6 ${ }^{\circ}$ NEE from mag.3.42 Acamar.

This lenticular galaxy is a respectable 60 million light-years from Earth, and despite the distance still managed to be the fourth-brightest radio sources in the sky. Since 1980, NGC 1316 has hosted four supernovae.

Also visible:
(1) NGC 1291 (9.39m barred spiral galaxy)
(2) NGC 1365 (10.32 $2_{m}$ barred spiral galaxy)
(3) NGC 1399 ( $10.55_{m}$ elliptical galaxy)
(4) NGC 1380 (10.87m lenticular galaxy)
(5) NGC 1350 (11.16 $6_{m}$ barred spiral galaxy)


RA: $49.33^{\circ}\left|3 \mathrm{~h} 17.3^{\prime}-\mathrm{DEC}:-41.1^{\circ}\right|-41^{\circ} 5^{\prime}$

NGC 1291 is a magnitude 9.39 barred spiral galaxy.
Angular size is 10 '.
3.6º SEE from mag.3.42 Acamar.

This galaxy has a relatively bright, elongated core encircled by a dim ring with twice the diameter. If you spot this one, you can cross two NGCs off your list as it is also cataloged as NGC 1269.

Also visible:
(1) NGC 1316 (9.42m lenticular galaxy)


RA: $58.2^{\circ} \mid 3 \mathrm{~h} 52.78^{\prime}-$ DEC: $-45.83^{\circ} \mid-45^{\circ} 49^{\prime}$
U Hor (HD 024607) is a magnitude 7.8 variable star. Magnitude ranges from 15.1 to 7.8 ( $\Delta$ mag. 7.3 ) with a period of 348 d .
$5.2^{\circ} \mathrm{SW}$ from mag. 3.83 alf Hor.

On the south-western edge of the finder circle, you can find the barred spiral galaxy NGC 1433 (magnitude 10).

Also visible:
(1) NGC 1433 (10.7m barred spiral galaxy)
(2) NGC 1512 (11.13 $3_{m}$ barred spiral galaxy)


R Hor (HD 018242) is a magnitude 4.7 variable star. Magnitude ranges from 14.3 to 4.7 ( $\Delta$ mag. 9.6) with a period of 408 d .
6.1 ${ }^{\circ}$ NEE from mag. 3.78 phi Eri.

With a range from magnitude 4.7 to 14.3, R Horologii has one of the largest brightness ranges of stars visible to the unaided eye.


RA: $48.08^{\circ}\left|3 \mathrm{~h} 12.3^{\prime}-\mathrm{DEC}:-55.22^{\circ}\right|-55^{\circ} 12^{\prime}$
C87 (NGC 1261) is a magnitude 8.4 globular cluster. Angular size is $7^{\prime}$.
$9.0^{\circ}$ SEE from mag. 3.78 phi Eri.

O:
The globular cluster is quite distant at 50,000 lightyears. Consequently it is dimmer and its member stars are exceedingly faint and unresolvable in anything but the most massive amateur telescopes.


RA: $49.58^{\circ} \mid 3 \mathrm{~h} 18.3^{\prime}-$ DEC: $-66.5^{\circ} \mid-66^{\circ} 29^{\prime}$

NGC 1313 is a magnitude 9.2 barred spiral galaxy.
Angular size is $9^{\prime}$.
3.1 ${ }^{\circ}$ SW from mag. 3.8 bet Ret.

Also known as the Topsy Turvy Galaxy, NGC 1313 is a field galaxy, which means it is not part of a galaxy group. It has a diameter of 50,000 light years and is 12.9 million light-years from Earth.

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## December: $\mathbf{4 5}{ }^{\circ}$ North




RA: $79.32^{\circ} \mid 5 \mathrm{~h} 17.29^{\prime}-$ DEC: $53.59^{\circ} \mid 53^{\circ} 35^{\prime}$

R Aur (HD 34019) is a magnitude 6.7 variable star.
Magnitude ranges from 13.9 to 6.7 ( $\Delta$ mag. 7.2) with a period of 458 d .
$6.2^{\circ} \mathrm{W}$ from mag. 3.88 del Aur.

The moderately red variable is also a double star, with a faint magnitude 10.1 companion separated by $45.1^{\prime \prime}$ at position angle $53^{\circ}$.


RA: $112.05^{\circ} \mid 7 \mathrm{~h} 28.19^{\prime}-$ DEC: $46.0^{\circ} \mid 46^{\circ} 0^{\prime}$

Y Lyn (HD 58521) is a magnitude 6.9 carbon star.
Magnitude ranges from 7.5 to 6.9 ( $\Delta$ mag. 0.6 ) with a period of 110 d .
$13.7^{\circ}$ NNE from mag. 3.64 the Gem.


RA: $102.33^{\circ} \mid 6 \mathrm{~h} 49.3^{\prime}-$ DEC: $41.07^{\circ} \mid 41^{\circ} 4^{\prime}$

NGC 2281 is a magnitude 5.4 open cluster. Angular size is $15^{\prime}$.
$7.1^{\circ} \mathrm{N}$ from mag.3.64 the Gem.

A bright, concentrated open cluster, with member stars ranging widely in brightness. The double star Psi7 Aurigae is one degree to the north east (magnitudes 5.1 and 10.9 ; separation 43 "; position angle $90^{\circ}$ ).

Also visible:
(1) UU Aur ( $5.1_{m}$ carbon star)


RA: $99.13^{\circ} \mid 6 \mathrm{~h} 36.5^{\prime}-$ DEC: $38.5^{\circ} \mid 38^{\circ} 30^{\prime}$
UU Aur (HD 46687) is a magnitude 5.1 carbon star. Magnitude ranges from 7.0 to 5.1 ( $\Delta$ mag. 1.9) with a period of 235 d .

On a line between Capella and Pollux, slightly closer to Capella.

UU Aurigae is an interesting star, being a variable star, a carbon star and a double star. With a color index of 2.78 and spectral class of C5II, the star is intensely red. It has a moderate brightness range of just over two magnitudes but even at its faintest it remains bright enough to show in small telescopes. It has a faint magnitude 11.8 companion separated by $117^{\prime \prime}$ at position angle $223^{\circ}$.

Also visible:
(1) NGC 2281 (5.4m open cluster)


RA: $82.18^{\circ} \mid 5 \mathrm{~h} 28.69^{\prime}-$ DEC: $35.83^{\circ} \mid 35^{\circ} 50^{\prime}$

M38 (NGC 1912) is a magnitude 7.4 open cluster.
Angular size is $21^{\prime}$.
$6.3^{\circ} \mathrm{W}$ from mag. 2.71 the Aur.

Discovered by Giovanni Batista Hodierna before 1654 , this 250 million year old cluster is 3,480 lightyears from Earth. It is my personal favorite of all the wonderful Auriga clusters.

Also visible:
(1) C31 ( $6.0_{m}$ bright nebula)
(2) M36 ( $6.3_{m}$ open cluster)
(3) AE Aur ( $5.78_{m}$ variable star)


RA: $79.08^{\circ} \mid 5 h 16.3^{\prime}-$ DEC: $34.31^{\circ} \mid 34^{\circ} 19^{\prime}$

AE Aur (HD 4078) is a magnitude 5.78 variable star.
Magnitude ranges from 6.08 to 5.78 ( $\Delta$ mag. 0.3 ).
4.1 ${ }^{\circ}$ NEE from mag.2.9 Hassaleh.

This eruptive blue variable star lights the appropriately named Flaming Star Nebula (Caldwell 31). The magnitude range of 0.3 is hardly to be seen, but magnitude 5.9 HIP 25471 is two degrees to the east, marking the midpoint of this star's brightness range.

Also visible:
(1) C 31 ( $6.0_{m}$ bright nebula)
(2) M36 (6.3m open cluster)
(3) M38 ( $7.4_{m}$ open cluster)


RA: $79.05^{\circ} \mid 5 \mathrm{~h} 16.19^{\prime}-$ DEC: $34.27^{\circ} \mid 34^{\circ} 16^{\prime}$
C31 (Flaming Star Nebula, IC 405) is a magnitude 6.0
bright nebula. Angular size is $30 \times 19^{\prime}$.
4.1 ${ }^{\circ}$ NEE from mag.2.9 Hassaleh.

O
The Flaming Star nebula is dominated by its eponymous stellar component which is surrounded a by a whorling halo of glowing gas and dust. The central star, AE Aurigae, originates from Orion's Belt.

Also visible:
(1) M36 (6.3 ${ }_{m}$ open cluster)
(2) M38 ( $7.4_{m}$ open cluster)
(3) AE Aur ( $5.78_{m}$ variable star)


M36 (NGC 1960) is a magnitude 6.3 open cluster.
Angular size is $12^{\prime}$.
5.6 ${ }^{\circ}$ SWW from mag.2.71 the Aur.

This bright open cluster is a more distant analog of the more famous Pleiades. It can be aged at around 22 million years based on the stars which have not yet consumed their lithium supplies. In a small telescope, the cluster is quite sparse as only ten stars are brighter than tenth magnitude.

Also visible:
(1) C 31 ( $6.0_{m}$ bright nebula)
(2) M37 ( $6.2_{m}$ open cluster)
(3) M38 ( $7.4_{m}$ open cluster)
(4) AE Aur (5.78m variable star)


RA: $88.1^{\circ} \mid 5 h 52.39^{\prime}-$ DEC: $32.55^{\circ} \mid 32^{\circ} 33^{\prime}$
M37 (NGC 2099) is a magnitude 6.2 open cluster.
Angular size is $24^{\prime}$.
$4.8^{\circ} \mathrm{SSW}$ from mag. 2.71 the Aur.
$\%$
The brightest of the renowned Auriga open clusters, this cluster has 150 stars that can be visible to a small telescope from a dark site. It is at least 347 million years old, far older than M36. In total it has at least 500 member stars with the collective mass of 1,500 Suns.

Also visible:
(1) M36 (6.3 $3_{m}$ open cluster)

## December: $\mathbf{1 5}^{\circ}$ North (1)



## December: $\mathbf{1 5}^{\circ}$ North (2)




RA: $92.23^{\circ} \mid 6$ h $8.89^{\prime}-$ DEC: $24.33^{\circ} \mid 24^{\circ} 20^{\prime}$

M35 (NGC 2168) is a magnitude 5.3 open cluster.
Angular size is $28^{\prime}$.
3.6 ${ }^{\circ}$ NWW from mag.3.19 Tejat Posterior.

This large cluster spanning half a degree is a fine sight at lower magnification in a telescope; given the delicacy of its stars it is best viewed on darker nights. It has a mass of 1,600 Suns. Less than a degree to the southwest is a smaller but denser open cluster, NGC 2158 (magnitude 8.6). Because of its compact form, NGC 2158 was initially misclassified as a globular cluster.

Also visible:
(1) NGC 2129 ( $6.7_{m}$ open cluster)
(2) NGC 2175 ( $6.8_{m}$ open cluster)
(3) U Ori (4.8m variable star)
(4) TU Gem ( $7.5_{m}$ carbon star)


RA: $106.84^{\circ} \mid 7 h 7.35^{\prime}-$ DEC: $22.7^{\circ} \mid 22^{\circ} 42^{\prime}$

R Gem (HD 053791) is a magnitude 6.0 variable star. Magnitude ranges from 14.0 to 6.0 ( $\Delta$ mag. 8.0) with a period of 370 d .
3.0 ${ }^{\circ}$ NWW from mag.3.51 Wasat.

This fairly red star is a technetium star, indicating its spectrum is enriched with this radioactive element. Technetium has a half-life of 4.2 million years, which means the element was formed by nucleosynthesis in the star. Specifically heavy elements like technetium are formed by slow neutron capture (the s-process) in the boundary between shells of fusing hydrogen and helium.


RA: $83.63^{\circ} \mid 5 \mathrm{~h} 34.5^{\prime}-$ DEC: $22.02^{\circ} \mid 22^{\circ} 1^{\prime}$
M1 (Crab Nebula, NGC 1952) is a magnitude 8.4 supernova remnant. Angular size is $6 \times 4{ }^{\prime}$.
$1.1^{\circ}$ NW from mag.3.0 zet Tau.


A bright and compact supernova remnant. The star exploded in 1054. When Charles Messier rediscovered the object he initially mistook it for Halley's Comet. To avoid future errors, he decided to create his famous catalog of fuzzy patches that were not comets, and Messier 1 was naturally the first entry.

Also visible:
(1) SU Tau (9.1m variable star)
(2) Y Tau (7.1m carbon star)


RA: $71.5^{\circ} \mid 4 \mathrm{~h} 46.0^{\prime}-$ DEC: $19.07^{\circ} \mid 19^{\circ} 4^{\prime}$
NGC 1647 is a magnitude 6.4 open cluster. Angular size is $45^{\prime}$.
3.4 ${ }^{\circ}$ NE from mag.1.06 Aldebaran.

150 million years old and 1,800 light-years from Earth, this cluster contains about 90 stars.

Also visible:
(1) C 41 (1.0 $0_{m}$ open cluster)


RA: $72.13^{\circ} \mid 4 \mathrm{~h} 48.5^{\prime}-$ DEC: $10.93^{\circ} \mid 10^{\circ} 56^{\prime}$
NGC 1662 is a magnitude 6.4 open cluster. Angular size is 20 '.
3.9 ${ }^{\circ} \mathrm{N}$ from mag.3.31 Tabit.

Located midway between Bellatrix and Aldebaran, this loose cluster is about 1,425 light-years from Earth.


Also visible:
(1) R Ori ( $9.05_{m}$ variable star)


RA: $88.8^{\circ} \mid 5 \mathrm{~h} 55.19^{\prime}-$ DEC: $7.4^{\circ} \mid 7^{\circ} 24^{\prime}$

Betelgeuse (HD 39801) is a magnitude 0.4 carbon star. Magnitude ranges from 1.3 to 0.4 ( $\Delta$ mag. 0.9 ).
$0.0^{\circ}$ SWW from mag. 0.92 Betelgeuse.

While some sources list Betelgeuse as a carbon star, it does not appear to have more carbon than oxygen in its outer layers, a pre-requisite for true carbon star. It is moderately red with a B-V color index of only 1.52 , but its high apparent magnitude makes the color more vivid. Betelgeuse is variable and at its dimmest is equivalent to Bellatrix, and at its brightest it slightly outshines Rigel.

Also visible:
(1) NGC 2022 (11.5m planetary nebula)


RA: $98.1^{\circ} \mid 6 h 32.39^{\prime}-$ DEC: $4.87^{\circ} \mid 4^{\circ} 52^{\prime}$
C50 (Rosette Nebula, NGC 2244) is a magnitude 4.8 open cluster. Angular size is 24 '.

Two and a half finder circles east and slightly south of Betelgeuse.

This extremely young star cluster has not drifted away from it's surrounding nebulosity (Caldwell 49), and it still possesses several brilliant but short-lived O type stars. The brightest member of the cluster is HD 46223, a vast star with 50 times the mass of the sun and 400,000 times the luminosity.

Also visible:
(1) C 46 (10.0 $\mathrm{o}_{\mathrm{m}}$ bright nebula)
(2) C49 (9.0 bright nebula)
(3) GK Ori ( $9.5_{m}$ carbon star)


RA: $102.95^{\circ} \mid 6 \mathrm{~h} 51.8^{\prime}-$ DEC: $0.47^{\circ} \mid 0^{\circ} 28^{\prime}$

NGC 2301 is a magnitude 6.0 open cluster. Angular size is $12^{\prime}$.
$11.7^{\circ} \mathrm{SW}$ from mag.3.09 Gomeisa.

This bright and rich open cluster is 165 million years old and is 2,840 light-years from Earth.

Also visible:
(1) NGC 2346 (10.5m planetary nebula)


RA: $75.9^{\circ}\left|5 \mathrm{~h} 3.6^{\prime}-\mathrm{DEC}: 23.82^{\circ}\right| 23^{\circ} 49^{\prime}$
NGC 1746 is a magnitude 6.0 open cluster. Angular size is $42^{\prime}$.

Travel from the Pleiades to Aldebaran, turn right, and travel an equal distance further.

Originally thought to be a poor and relatively scattered open cluster, more recent studies suggest this object is an asterism, merely a chance alignment of unrelated stars.

Also visible:
(1) NGC 1647 (6.4m open cluster)


RA: $112.3^{\circ} \mid 7 \mathrm{~h} 29.19^{\prime}-$ DEC: $20.92^{\circ} \mid 20^{\circ} 55^{\prime}$

C39 (Eskimo Nebula, NGC 2392) is a magnitude 9.9 planetary nebula. Angular size is $0.8^{\prime}$.
$2.3^{\circ}$ SEE from mag.3.51 Wasat.

The Clown Face Nebula lies in an interesting field in the finderscope. Double star Delta Geminorum (mag. 3.5 , separation $5.6^{\prime \prime}$ ) is on the north-western edge of the finder view, while the small but concentrated open cluster NGC 2420 (mag. 8.3) is on the eastern edge.

Also visible:
(1) S Gem ( $8.0_{m}$ variable star)
(2) T Gem (8.0 $0_{m}$ variable star)


RA: $88.95^{\circ} \mid 5$ h $55.81^{\prime}-$ DEC: $20.18^{\circ} \mid 20^{\circ} 11^{\prime}$
U Ori (HD 039816) is a magnitude 4.8 variable star.
Magnitude ranges from 13.0 to 4.8 ( $\Delta$ mag. 8.2) with a period of 368 d .

Just less than halfway from Betelgeuse to Theta Aurigae.


A good example of a long-period variable, being an easy unaided eye object at its brightest (assuming the skies are unpolluted), while disappearing entirely from view in small to medium telescopes at its dimmest.


Also visible:
(1) M35 (5.3 $3_{m}$ open cluster)
(2) NGC 2129 ( $6.7_{m}$ open cluster)
(3) NGC 2175 ( $6.8_{m}$ open cluster)
(4) SU Tau (9.1m variable star)
(5) Y Tau (7.1 $1_{m}$ carbon star)


RA: $92.1^{\circ} \mid$ 6h $8.39^{\prime}-$ DEC: $13.95^{\circ} \mid 13^{\circ} 57^{\prime}$
NGC 2169 is a magnitude 5.9 open cluster. Angular size is $7^{\prime}$.

Halfway between Betelgeuse and Tejat in Gemini.

Known as the " 37 Cluster" as the brightest stars of the cluster form these numerals, the stars also form the letters "XY" in smaller telescopes that reveal fewer stars.


RA: $99.8^{\circ} \mid 6 \mathrm{~h} 39.19^{\prime}-$ DEC: $8.73^{\circ} \mid 8^{\circ} 44^{\prime}$
C46 (Hubble's Variable Nebula, NGC 2261) is a magnitude 10.0 bright nebula. Angular size is $2 \times 11^{\prime}$.
4.4 ${ }^{\circ}$ SSW from mag.3.4 Alzirr.

Small dust clouds drifting close to the variable star R Monocerotis cast vastly larger and highly mobile shadows on the more distant reflection nebula. The appearance can accordingly change significantly in a period of weeks, leading to the name of Hubble's Variable Nebula. As a reflection nebula, this object is sensitive to light pollution and particularly the blue cast of moonlight.

Also visible:
(1) C49 (9.0 $0_{m}$ bright nebula)
(2) C50 (4.8m open cluster)
(3) BG Mon ( $9.2_{m}$ carbon star)


C49 (NGC 2237-9) is a magnitude 9.0 bright nebula. Angular size is $80 \times 60^{\prime}$.

Two and a half finder circles east and slightly south of Betelgeuse.

This large nebula can be viewed with binoculars from a dark site. While diffuse nebulae can be a challenge in light-polluted surroundings, the location of Caldwell 49 is marked by the closely associated star cluster Caldwell 50.

Also visible:
(1) C46 ( $10.0_{m}$ bright nebula)
(2) C50 (4.8m open cluster)
(3) GK Ori ( $9.5_{m}$ carbon star)


RA: $76.35^{\circ} \mid 5 \mathrm{~h} 5.39^{\prime}-$ DEC: $1.2^{\circ} \mid 1^{\circ} 12^{\prime}$

W Ori (HD 32736) is a magnitude 6.5 carbon star.
Magnitude ranges from 10.0 to 6.5 ( $\Delta$ mag. 3.5) with a period of 210 d .

W Orionis forms a kite shape with Bellatrix, Mintaka and Rigel.

The unusual spectrum of this carbon star was first noted in the 1860 s and is now classified as a C5,4 carbon star. W Orionis is exceptionally red with a B-V color index of 3.34.


RA: $86.68^{\circ} \mid 5 \mathrm{~h} 46.69^{\prime}-$ DEC: $0.05^{\circ} \mid 0^{\circ} 3^{\prime}$

M78 (NGC 2068) is a magnitude 8.3 diffuse nebula.
Angular size is $8 \times 6$.
Follow Orion's Belt from Mintaka to Alnitak, then turn right and continue an equal distance.

This complex and confusing reflection nebula hides a young star cluster in its dust clouds. The cluster is however detectable in infrared wavelengths.

Also visible:
(1) NGC 1981 (4.6m open cluster)
(2) VV Ori (5.31m variable star)

## December: - $15^{\circ}$ South (1)



## December: -15 ${ }^{\circ}$ South (2)




RA: $83.38^{\circ} \mid 5 h 33.52^{\prime}-$ DEC: $-1.16^{\circ} \mid-1^{\circ} 8^{\prime}$
VV Ori (HD 036695) is a magnitude 5.31 variable star. Magnitude ranges from 5.66 to 5.31 ( $\Delta$ mag. 0.4 ) with a period of 1.48538 d .
$0.6^{\circ} \mathrm{W}$ from mag.1.75 Alnilam.


This eclipsing binary has a very slight magnitude variation, but half a degree to the SSE is the magnitude 5.30 star SAO 132234 (HD 36591), marking the brightest point of VV Ori's light curve. Conveniently SAO 132234 and VV Ori are the same color making comparison more easy. SAO 132234 is also a tight double with a magnitude 9.8 companion (separation 1.96 ", position angle $355^{\circ}$ ).

Also visible:
(1) M42 (4.0 $0_{m}$ diffuse nebula)
(2) M43 (9.0 $0_{m}$ diffuse nebula)
(3) M78 (8.3 ${ }_{m}$ diffuse nebula)
(4) NGC 1981 (4.6m open cluster)


RA: $83.8^{\circ}\left|5 \mathrm{~h} 35.19^{\prime}-\mathrm{DEC}:-4.43^{\circ}\right|-4^{\circ} 25^{\prime}$

NGC 1981 is a magnitude 4.6 open cluster. Angular size is $25^{\prime}$.
1.4 ${ }^{\circ} \mathrm{N}$ from mag.2.87 Hatysa.

This lively scattering of bright stars is often overlooked as it shares the eyepiece view with the Great Orion Nebula. Take a moment next time you admire the great nebula to look a degree north and acknowledge this open cluster too. In between these two objects is the much more elusive Running Man emission-reflection nebula.

Also visible:
(1) M42 (4.0 $0_{m}$ diffuse nebula)
(2) M43 (9.0 $0_{m}$ diffuse nebula)
(3) M78 (8.3 $3_{m}$ diffuse nebula)
(4) VV Ori (5.31 $1_{m}$ variable star)


## M43

RA: $83.9^{\circ} \mid 5 \mathrm{~h} 35.6^{\prime}-$ DEC: $-5.27^{\circ} \mid-5^{\circ} 15^{\prime}$

M43 (De Mairan's Nebula, NGC 1982) is a magnitude 9.0 diffuse nebula. Angular size is $20 \times 15$ '.
0.6 ${ }^{\circ} \mathrm{N}$ from mag.2.87 Hatysa.

Unless you knew it was cataloged separately as M43 by Charles Messier, you would simply consider M43 to be part of M42, merely separated by a dust lane and you would be correct. Beyond M43 is the much more challenging Running Man Nebula (cataloged as three nebulae by William Herschel: NGC 1973, NGC 1975 and NGC 1977).

Also visible:
(1) M42 (4.0 $0_{m}$ diffuse nebula)
(2) NGC 1981 (4.6m open cluster)
(3) VV Ori (5.31m variable star)


RA: $105.8^{\circ} \mid 7 h 3.19^{\prime}-$ DEC: $-8.33^{\circ} \mid-8^{\circ} 19^{\prime}$

M50 (NGC 2323) is a magnitude 6.3 open cluster.
Angular size is 16 .
One third of the distance from Sirius to Procyon.

Discovered on G D Cassini on or before 1711, the small cluster ( 285 solar masses) is just detectable by the naked eye in very dark conditions. It is quite compact with a core radius of just under 6 light-years. It is quite young at 140 million years.

Also visible:
(1) NGC 2343 ( $6.7_{m}$ open cluster)
(2) W CMa ( $7.0_{m}$ carbon star)


RA: $74.9^{\circ}\left|4 \mathrm{~h} 59.6^{\prime}-\mathrm{DEC}:-14.81^{\circ}\right|-14^{\circ} 47^{\prime}$
R Lep (HD 031996) is a magnitude 5.5 variable star. Magnitude ranges from 11.7 to 5.5 ( $\Delta$ mag. 6.2 ) with a period of 427 d .

Draw a line from Mintaka through Rigel and continue it for almost an equal length.

Better known as Hind's Crimson Star, this deeply red carbon star (B-V index of 2.71) is over 6,000 times brighter than the Sun. Like many carbon stars, R Lep is redder when it is fainter, meaning its deeper reds are better accessible by larger telescopes. When he discovered R Lep in 1845, Hind reported that it resembled "a drop of blood on a black field."

Also visible:
(1) RX Lep (5.0 $0_{m}$ variable star)


RA: $109.87^{\circ} \mid 7 \mathrm{~h} 19.46^{\prime}-$ DEC: $-16.4^{\circ} \mid-16^{\circ} 23^{\prime}$

R CMa (HD 57167) is a magnitude 5.7 variable star. Magnitude ranges from 6.34 to 5.7 ( $\Delta$ mag. 0.6 ) with a period of 1.13594 d .

Forms an equilateral triangle with Adhara and Mirzam.

This eclipsing binary star is a degree SSE of Caldwell 58.

Also visible:
(1) C58 (7.2m open cluster)
(2) M47 (5.2 $2_{m}$ open cluster)
(3) NGC 2423 ( $6.7_{m}$ open cluster)
(4) W CMa ( $7.0_{m}$ carbon star)


RA: $81.13^{\circ} \mid 5 \mathrm{~h} 24.5^{\prime}-$ DEC: $-24.55^{\circ} \mid-24^{\circ} 32^{\prime}$
M79 (NGC 1904) is a magnitude 7.7 globular cluster.
Angular size is $9^{\prime}$.
Draw a line from Arneb through Mihal in Lepus, and extend the line an equal distance.

This globular cluster is believed to be extragalactic, belonging not to the Milky way but to the Canis Major Dwarf Galaxy which, if it exists, is an elliptical concentration of a billion stars in the direction of Canis Major. Aside from its globular clusters (M79, NGC 1851, NGC 2298 and NGC 2808) this putative, highly disrupted galaxy is undetectable by amateur instruments.


RA: $108.58^{\circ} \mid 7 \mathrm{~h} 14.3^{\prime}-$ DEC: $-25.73^{\circ} \mid-25^{\circ} 43^{\prime}$

NGC 2354 is a magnitude 6.5 open cluster. Angular size is 20 .
$1.0^{\circ} \mathrm{N}$ from mag. 3.83 ome CMa .

Also visible:
(1) C64 (4.1m open cluster)
(2) VY CMa ( $6.5_{m}$ variable star)
(3) VY CMa ( $8.8_{m}$ carbon star)


RA: $95.68^{\circ} \mid 6$ h $22.72^{\prime}-$ DEC: $-2.2^{\circ} \mid-2^{\circ} 11^{\prime}$
V Mon (HD 044639) is a magnitude 6.0 variable star.
Magnitude ranges from 13.9 to 6.0 ( $\Delta$ mag. 7.9 ) with a period of 341 d .
$10.4^{\circ}$ E from mag. 2.05 Alnitak.


This orange long-period variable is roughly 1,388 light-years from Earth.

Also visible:
(1) NGC 2232 (3.9m open cluster)
(2) KS Mon ( $8.5_{m}$ carbon star)


RA: $96.65^{\circ} \mid 6$ h $26.6^{\prime}-$ DEC: $-4.75^{\circ} \mid-4^{\circ} 44^{\prime}$

NGC 2232 is a magnitude 3.9 open cluster. Angular size is $30^{\prime}$.
$10.8^{\circ}$ NEE from mag.2.2 Saiph.

At a distance of 1,060 light-years, this bright grouping is one of the closest open clusters to Earth. It is estimated to be roughly 31 million years old.

Also visible:
(1) V Mon ( $6.0_{m}$ variable star)
(2) KS Mon ( $8.5_{m}$ carbon star)


RA: $83.85^{\circ} \mid 5 \mathrm{~h} 35.39^{\prime}-$ DEC: $-5.45^{\circ} \mid-5^{\circ} 26^{\prime}$

M42 (Great Nebula in Orion, NGC 1976) is a magnitude 4.0 diffuse nebula. Angular size is $85 \times 60$ '.
0.4 ${ }^{\circ} \mathrm{N}$ from mag.2.87 Hatysa.

An infallible nebula, visible with the naked eye even from very light-polluted locations. This is the closest star forming region to Earth at a distance of around 1,344 light-years. At its heart is the Trapezium star cluster. The nebula spans 24 light-years and has a mass of 2,000 Suns.

Also visible:
(1) M43 (9.0 $0_{m}$ diffuse nebula)
(2) NGC 1981 (4.6 $\sigma_{m}$ open cluster)
(3) VV Ori ( $5.31_{m}$ variable star)


RA: $77.85^{\circ}\left|5 \mathrm{~h} 11.38^{\prime}-\mathrm{DEC}:-11.85^{\circ}\right|-11^{\circ} 50^{\prime}$
RX Lep (HD 033664) is a magnitude 5.0 variable star. Magnitude ranges from 7.4 to $5.0(\Delta$ mag. 2.4$)$ with a period of 60 d .

Travel from Saiph to Rigel and turn right, travel a further distance equal to Orion's Belt.

This pulsating variable star offers a moderate magnitude range over a relatively compressed period. $2.5^{\circ}$ to the southeast, magnitude 5.45 SAO 150304 (HD 34538) marks the brighter end of RX Lep's range. The fainter end is marked by magnitude 7.55 SAO 150215 (HD 33755), one degree to the north of RX Lep.

Also visible:
(1) R Lep ( $5.5_{m}$ variable star)
(2) R Lep ( $5.9_{m}$ carbon star)


RA: $109.45^{\circ} \mid 7 \mathrm{~h} 17.8^{\prime}-$ DEC: $-15.62^{\circ} \mid-15^{\circ} 36^{\prime}$
C58 (Caroline's Cluster, NGC 2360) is a magnitude 7.2 open cluster. Angular size is $13^{\prime}$.
$7.9^{\circ}$ E from mag.-1.58 Sirius.

Discovered by Caroline Herschel in 1783, she described this object as a "beautiful cluster of pretty compressed stars near $1 / 2$ degree in diameter". The cluster has a diameter of 15 light-years and is aged at 2.2 million years. Given the age, the cluster has a few short-lived bright blue stars. The "blue stragglers" are actually older, smaller stars that absorbed mass from a companion star, becoming bright and blue in consequence.

Also visible:
(1) M47 (5.2m open cluster)
(2) NGC 2423 ( $6.7_{m}$ open cluster)
(3) $\mathrm{R} \mathrm{CMa} \mathrm{(5.7m} \mathrm{variable} \mathrm{star)}$
(4) W CMa ( $7.0_{m}$ carbon star)


RA: $101.75^{\circ}\left|6 \mathrm{~h} 47.0^{\prime}-\mathrm{DEC}:-20.73^{\circ}\right|-20^{\circ} 43^{\prime}$

M41 (NGC 2287) is a magnitude 4.6 open cluster.
Angular size is $38^{\prime}$.
$4.0^{\circ} \mathrm{S}$ from mag.-1.58 Sirius.

Possibly known in ancient times, this very bright open cluster is graced with a number of prominent red stars. It is approximately 190 million years old and is 2,300 light-years from Earth.


RA: $109.7^{\circ} \mid 7 \mathrm{~h} 18.8^{\prime}-$ DEC: $-24.95^{\circ} \mid-24^{\circ} 56^{\prime}$
C64 (Pirate's Jewels Cluster, NGC 2362) is a magnitude 4.1 open cluster. Angular size is $8^{\prime}$.
$2.0^{\circ}$ NNE from mag. 3.83 ome CMa.


Discovered in 1654 , this beautiful and bright cluster is only 4.5 million years old with a collective size of around 500 solar masses.


Also visible:
(1) NGC 2354 ( $6.5_{m}$ open cluster)
(2) VY CMa ( $6.5_{m}$ variable star)
(3) VY CMa ( $8.8_{m}$ carbon star)


RA: $110.74^{\circ}$ | 7h $22.97^{\prime}-$ DEC: $-25.77^{\circ} \mid-25^{\circ} 45^{\prime}$
VY CMa (HD 58061) is a magnitude 6.5 variable star.
Magnitude ranges from 9.6 to 6.5 ( $\Delta$ mag. 3.1).
$2.0^{\circ}$ NEE from mag. 3.83 ome CMa.

Distinguished by a relatively long period, this red pulsating variable star is roughly 1,832 light years from Earth.

Also visible:
(1) C64 (4.1m open cluster)
(2) NGC 2354 ( $6.5_{m}$ open cluster)

## December: $\mathbf{- 4 5}{ }^{\circ}$ South


$\beta$ Dor: page 134


RA: $79.82^{\circ} \mid 5 \mathrm{~h} 19.28^{\prime}-$ DEC: $-33.71^{\circ} \mid-33^{\circ} 41^{\prime}$

T Col (HD 34897) is a magnitude 6.6 variable star.
Magnitude ranges from 12.7 to 6.6 ( $\Delta$ mag. 6.1 ) with a period of 226 d .
$3.0^{\circ} \mathrm{NW}$ from mag. 3.92 eps Col.


This orange-red pulsating variable star is roughly 2.038 light-years from Earth.

Also visible:
(1) NGC 1808 (10.74 $4_{m}$ barred spiral galaxy)
(2) NGC 1792 (10.87m barred spiral galaxy)


RA: $102.25^{\circ}\left|6 \mathrm{~h} 48.98^{\prime}-\mathrm{DEC}:-36.01^{\circ}\right|-36^{\circ} 0^{\prime}$

NGC 2298 (NGC) is a magnitude 9.29 globular cluster. Angular size is $5^{\prime}$.
$3.5^{\circ} \mathrm{S}$ from mag. 3.78 kap CMa .

This globular cluster lies 34,900 light-years from Earth, and is believed to originate from the Canis Major Dwarf galaxy. The cluster shows signs of disruption by the gravitational effects of the Milky way, an exhibits a tidal tail.


RA: $78.53^{\circ} \mid 5 \mathrm{~h} 14.1^{\prime}-$ DEC: $-40.05^{\circ} \mid-40^{\circ} 2^{\prime}$
C73 (NGC 1851) is a magnitude 7.3 globular cluster.
Angular size is $11^{\prime}$.
$5.6^{\circ} \mathrm{SW}$ from mag.3.92 eps Col.

O:
This bright globular cluster is interesting as it contain populations of stars with different ages, whereas globular clusters are usually composed exclusively of ancient stars.

Also visible:
(1) NGC 1808 (10.74m barred spiral galaxy)
(2) NGC 1792 (10.87m barred spiral galaxy)


RA: $71.54^{\circ} \mid 4 \mathrm{~h} 46.15^{\prime}-$ DEC: $-49.25^{\circ} \mid-49^{\circ} 14^{\prime}$

R Pic (HD 030551) is a magnitude 6.35 variable star. Magnitude ranges from 10.1 to 6.35 ( $\Delta$ mag. 3.8) with a period of 171 d .
6.0 ${ }^{\circ}$ NNE from mag. 3.47 alf Dor.
$\because$
Nearby SAO 217006 (HD 30850) lies half a degree to the southeast; with magnitude 7.45 it marks the midpoint of the brightness range of R Pic.


RA: $83.41^{\circ} \mid 5 \mathrm{~h} 33.62^{\prime}-$ DEC: $-62.49^{\circ} \mid-62^{\circ} 28^{\prime}$
$\beta$ Dor (HD 037350) is a magnitude 3.46 variable star. Magnitude ranges from 4.08 to 3.46 ( $\Delta$ mag. 0.6 ) with a period of 9.8426 d .
$\beta$ Dor i a prminent star ten degrees SW of Canopus.

At its faintest, $\beta$ Dor is the same brightness as $\delta$ Dor which lies three degrees to the south.

February: $\mathbf{4 5}^{\circ}$ North


C7: page 136
W UMa: page 136
RT UMa: page 137
C25: page 137
R LMi: page 138


RA: $114.23^{\circ} \mid 7 \mathrm{~h} 36.89^{\prime}-$ DEC: $65.6^{\circ} \mid 65^{\circ} 36^{\prime}$
C7 (NGC 2403) is a magnitude 8.9 spiral galaxy.
Angular size is $22 \times 12^{\prime}$.
7.7 ${ }^{\circ}$ NW from mag.3.47 omi UMa.

A bright, obliquely oriented spiral galaxy.


RA: $145.94^{\circ} \mid 9 \mathrm{~h} 43.75^{\prime}-$ DEC: $55.95^{\circ} \mid 55^{\circ} 57^{\prime}$
W UMa (HD 083950) is a magnitude 7.75 variable star. Magnitude ranges from 8.48 to 7.75 ( $\Delta$ mag. 0.7) with a period of 0.3336 d .
$3.2^{\circ}$ SSW from mag.3.89 ups UMa.

This star is the prototype of the W Ursae Majoris class of variables. These stars are contact binaries of spectral types F, G, or K. The two distorted stars exchange gas through a connecting "neck". The close pair are eclipsing over a very short period, so the moderate light curve can be observed in a period of hours. W UMa is surrounded by three stars of approximately mag. 9.0 , which can be used as a reference for the minima.

Also visible:
(1) NGC 2768 ( $10.84_{m}$ elliptical galaxy)
(2) RT UMa (8.6 $\sigma_{m}$ carbon star)


RA: $139.6^{\circ} \mid 9 \mathrm{~h} 18.39^{\prime}-$ DEC: $51.4^{\circ} \mid 51^{\circ} 24^{\prime}$
RT UMa (TYC 3431-229-1) is a magnitude 8.6 carbon star. Magnitude ranges from 9.6 to 8.6 ( $\Delta$ mag. $1.0)$.
2.2 ${ }^{\circ} \mathrm{W}$ from mag.3.26 the UMa.


This carbon star is deeply red with a B-V color index of 3.5 . The spectral class is $\mathrm{C} 4+, 4$.

Also visible:
(1) NGC 2841 (10.09 barred spiral galaxy)
(2) NGC 2681 (11.09 ${ }_{m}$ barred spiral galaxy)
(3) $\mathrm{W} \mathrm{UMa}\left(7.75_{m}\right.$ variable star)


RA: $114.53^{\circ} \mid 7 \mathrm{~h} 38.1^{\prime}-$ DEC: $38.88^{\circ} \mid 38^{\circ} 53^{\prime}$

C25 (Intergalactic Wanderer, NGC 2419) is a magnitude 10.4 globular cluster. Angular size is $6^{\prime}$.

Due north of Castor, roughly twice the distance between Castor and Pollux.

A very distant globular cluster, more distant than the Magellanic Clouds, this object may not have formed with the Milky Way, and may instead be a visitor from afar. Because of the extreme distance of the cluster, individual stars cannot be resolved with moderatelysized telescopes.


RA: $146.39^{\circ} \mid 9 \mathrm{~h} 45.57^{\prime}-$ DEC: $34.51^{\circ} \mid 34^{\circ} 31^{\prime}$

R LMi (HD 084346) is a magnitude 6.3 variable star. Magnitude ranges from 13.2 to 6.3 ( $\Delta$ mag. 6.9 ) with a period of 372 d .
$5.0^{\circ}$ E from mag.3.3 alf Lyn.

This orange variable at its brightest rivals nearby magnitude 6.1013 LMi , while at its faintest it disappears from view in small and moderate telescopes.

February: $\mathbf{1 5}^{\circ}$ North



RA: $130.03^{\circ} \mid 8 \mathrm{~h} 40.1^{\prime}-$ DEC: $19.98^{\circ} \mid 19^{\circ} 59^{\prime}$

M44 (Beehive Cluster, NGC 2632) is a magnitude 3.7 open cluster. Angular size is 95 '.

## 12.1${ }^{\circ}$ NNE from mag.3.76 Altarf.

Truly spectacular from a dark site, this large open cluster has around 1,000 member stars and has a mass of around 600 Suns. Only $2 \%$ of the stars in the cluster are bright A-type stars, and there are only five giant stars.

Also visible:
(1) S Cnc ( $8.29_{m}$ variable star)
(2) T Cnc ( $7.6_{m}$ variable star)
(3) X Cnc ( $5.6_{m}$ variable star)
(4) T Cnc (7.8m carbon star)


X Cnc (HD 76221) is a magnitude 5.6 variable star. Magnitude ranges from 7.5 to $5.6(\Delta$ mag. 1.9$)$ with a period of 195d.

Halfway between Regulus of Leo and Wasat in Gemini.

One of the brightest carbon stars in the sky, X Cnc has a B-V color index 2.96 so it is markedly red, and is still reasonably bright at its minimum, at which point it exhibits its deepest red. 4,600 times brighter than the Sun, X Cnc is 2,900 light-years from Earth.

Also visible:
(1) M44 (3.7m open cluster)
(2) $\mathrm{S} \mathrm{Cnc}\left(8.29_{m}\right.$ variable star)
(3) T Cnc (7. $6_{m}$ variable star)
(4) T Cnc ( $7.8_{m}$ carbon star)


RA: $124.14^{\circ} \mid 8 \mathrm{~h} 16.56^{\prime}-$ DEC: $11.73^{\circ} \mid 11^{\circ} 44^{\prime}$

R Cnc (HD 69243) is a magnitude 6.07 variable star. Magnitude ranges from 11.8 to 6.07 ( $\Delta$ mag. 5.7) with a period of 362 d .
$2.5^{\circ} \mathrm{N}$ from mag.3.76 Altarf.


R Cnc is a Mira-type variable 830 light-years from Earth.


RA: $132.6^{\circ} \mid 8 \mathrm{~h} 50.39^{\prime}-$ DEC: $11.82^{\circ} \mid 11^{\circ} 49^{\prime}$
M67 (NGC 2682) is a magnitude 6.1 open cluster. Angular size is 30 '.
5.4 ${ }^{\circ} \mathrm{N}$ from mag.3.48 eps Hya.

A truly ancient open cluster (by the standards of open clusters) at three to five billion years age. It has a mass in the region 1000 to 1,500 solar masses, and is rich in stars similar to the Sun. The cluster is a good example of mass segregation, whereby velocity is transferred preferentially to less massive members of the cluster which causes the cluster to lose its lighter members over time.


RA: $146.89^{\circ} \mid 9 \mathrm{~h} 47.55^{\prime}-$ DEC: $11.43^{\circ} \mid 11^{\circ} 26^{\prime}$
R Leo (HD 084748) is a magnitude 4.4 variable star.
Magnitude ranges from 11.3 to 4.4 ( $\Delta$ mag. 6.9) with a period of 310 d .

Travel from Eta Leonis to Regulus, turn left, and travel an equal distance onward.

A reddish Mira-type variable with a bright maximum and deep minimum.


RA: $113.18^{\circ} \mid 7 \mathrm{~h} 32.71^{\prime}-$ DEC: $8.32^{\circ} \mid 8^{\circ} 19^{\prime}$
S CMi (HD 59950) is a magnitude 6.6 variable star. Magnitude ranges from 13.2 to 6.6 ( $\Delta$ mag. 6.6) with a period of 333 d .
$1.3^{\circ} \mathrm{E}$ from mag.3.09 Gomeisa.

A Mira-type variable, S CMi is shedding significant quantities of silicate dust from its outer layers to surrounding space. It is 6,500 times brighter than the Sun.


RA: $137.57^{\circ} \mid 9 \mathrm{~h} 10.29^{\prime}-$ DEC: $7.03^{\circ} \mid 7^{\circ} 2^{\prime}$
C48 (NGC 2775) is a magnitude 10.3 spiral galaxy.
Angular size is $4.5 \times 3$ '.
Two-fifths of the distance from Regulus to Procyon, slightly south of this line.

Small telescopes can show the large yellow core of the galaxy, but even large telescopes struggle to reveal the surrounding spiral arms.


Also visible:
(1) S Hya ( $7.2_{m}$ variable star)


RA: $133.39^{\circ} \mid 8 \mathrm{~h} 53.56^{\prime}-$ DEC: $3.07^{\circ} \mid 3^{\circ} 4^{\prime}$

S Hya (HD 076011) is a magnitude 7.2 variable star. Magnitude ranges from 13.3 to 7.2 ( $\Delta$ mag. 6.1) with a period of 257 d .

One third of the distance from Alphard to Pollux.

This pulsating variable star is roughly 2,700 lightyears from Earth.

Also visible:
(1) C48 (10.3m spiral galaxy)

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## February: $\mathbf{- 1 5}^{\circ}$ South (1)



February: $\mathbf{- 1 5}^{\circ}$ South (2)



## M48

RA: $123.45^{\circ} \mid 8 \mathrm{~h} 13.8^{\prime}-$ DEC: $-5.8^{\circ} \mid-5^{\circ} 47^{\prime}$
M48 (NGC 2548) is a magnitude 5.5 open cluster.
Angular size is 54 '.
Pointed to by lines from Mirzam through Sirius and Adhara through Wezen.

This cluster was observed by Charles Messier in 1771, but he recorded a declination that was wrong by five degrees. As such the discovery is often credited to Caroline Herschel in 1783. This massive open cluster ( 2,300 solar masses) is in the process of fragmenting: it is divided into three distinct lobes with different proper motion. M48 is 2,500 light-years from Earth.


RA: $112.7^{\circ} \mid 7 \mathrm{~h} 30.79^{\prime}-$ DEC: $-9.78^{\circ} \mid-9^{\circ} 46^{\prime}$
U Mon (HD 059693) is a magnitude 6.1 variable star. Magnitude ranges from 8.8 to $6.1(\Delta$ mag. 2.7$)$ with a period of 91.3 d .
$13.0^{\circ}$ NEE from mag.-1.58 Sirius.

Located in a bright and colorful field, U Mon is neighbored by SAO 152986 to the south (magnitude 6.6), marking the brighter limit of this star's range. Half a degree to the east is SAO 134821 (magnitude 7.6) which acts as a guide to the middle of $U$ Mon's range.

Also visible:
(1) M47 (5.2m open cluster)
(2) NGC 2423 ( $6.7_{m}$ open cluster)


RA: $122.68^{\circ} \mid 8 \mathrm{~h} 10.69^{\prime}-$ DEC: $-12.83^{\circ} \mid-12^{\circ} 49^{\prime}$

NGC 2539 is a magnitude 6.5 open cluster. Angular size is $22^{\prime}$.
$9.6^{\circ}$ SSW from mag.3.95 C Hya.


This moderately rich cluster has around 50-100 stars.

Also visible:
(1) C54 (7.6 $m_{m}$ open cluster)


RA: $114.15^{\circ} \mid 7 \mathrm{~h} 36.6^{\prime}-$ DEC: $-14.5^{\circ} \mid-14^{\circ} 29^{\prime}$

M47 (NGC 2422) is a magnitude 5.2 open cluster.
Angular size is $30^{\prime}$.
Forms the sharp tip of a slightly extended isosceles triangle with Sirius and Wezen (the bright star in Canis Major's tail).

This cluster was discovered and rediscovered by Giovanni Battista Hodierna (1654), Charles Messier (1771) and Caroline Herschel. Messier mistakenly transposed the Right Ascension and Declination signs in his work, so for centuries M47 was considered a "lost Messier". M47 is 78 million years old, has the mass of 453 Suns, and spans 30 light years.

Also visible:
(1) C 58 ( $7.2_{m}$ open cluster)
(2) M46 ( $6.0_{m}$ open cluster)
(3) NGC 2423 ( $6.7_{m}$ open cluster)
(4) R CMa (5.7 $7_{m}$ variable star)
(5) U Mon ( $6.1_{m}$ variable star)


RA: $156.2^{\circ} \mid 10 \mathrm{~h} 24.79^{\prime}-$ DEC: $-18.63^{\circ} \mid-18^{\circ} 37^{\prime}$
C59 (Ghost of Jupiter, NGC 3242) is a magnitude 8.6 planetary nebula. Angular size is $0.4^{\prime}$.
6.4 ${ }^{\circ}$ SWW from mag. 3.32 nu. Hya.

A bright planetary nebula with rich structures that be picked out with larger telescopes. The nebula spans 2 light-years, and the inner shell of expelled gas was ejected from the central white dwarf 1,500 years ago.


RA: $121.33^{\circ} \mid 8 \mathrm{~h} 5.3^{\prime}-$ DEC: $-28.17^{\circ} \mid-28^{\circ} 9^{\prime}$

NGC 2527 is a magnitude 6.5 open cluster. Angular size is $22^{\prime}$.
$3.8^{\circ} \mathrm{S}$ from mag.2.88 rho Pup.

A loose scattering of stars, originally discovered by William Herschel on December 9, 1784. His son John rediscovered the cluster in 1837, giving it the duplicate identifier NGC 2520.

Also visible:
(1) M93 ( $6.0_{m}$ open cluster)
(2) NGC 2452 (12.5m planetary nebula)


RA: $151.3^{\circ}$ | 10h 5.2' - DEC: $-7.72^{\circ}$ | $-7^{\circ} 42^{\prime}$
C53 (Spindle Galaxy, NGC 3115) is a magnitude 9.1
elliptical galaxy. Angular size is $8 \times 3{ }^{\prime}$.
4.8응․


This edge-on lenticular galaxy is quite accessible to smaller telescopes, which can reveal the elongated shape and bright core.


C54 (NGC 2506) is a magnitude 7.6 open cluster.
Angular size is $7^{\prime}$.
Midway between Alphard and Sirius, bearing slightly toward Procyon.

A fine sight in binoculars and any telescope, the member stars of this cluster are quite faint but numerous. At an age of 1.1 billion years, this is an elderly open cluster.

Also visible:
(1) M46 ( $6.0_{m}$ open cluster)
(2) NGC 2539 ( $6.5_{m}$ open cluster)


RA: $114.28^{\circ} \mid 7 \mathrm{~h} 37.1^{\prime}-$ DEC: $-13.87^{\circ} \mid-13^{\circ} 51^{\prime}$
NGC 2423 is a magnitude 6.7 open cluster. Angular size is $1^{\prime}$.

Forms the sharp tip of a slightly extended isosceles triangle with Sirius and Wezen (the bright star in Canis Major's tail).

The finder circle is crowded with interesting objects

- Messier 47 is less than a degree to the south, while Messier 46 is nearly two degrees to the southeast. The faint cluster NGC 2414 is two degrees to the southwest. Messier 46 includes the faint planetary nebula NGC 2438.


Also visible:
(1) C58 (7.2m open cluster)
(2) M46 ( $6.0_{m}$ open cluster)
(3) M47 ( $5.2_{m}$ open cluster)
(4) R CMa (5.7m variable star)
(5) U Mon (6.1m variable star)


RA: $115.45^{\circ} \mid 7 \mathrm{~h} 41.8^{\prime}-$ DEC: $-14.82^{\circ} \mid-14^{\circ} 48^{\prime}$
M46 (NGC 2437) is a magnitude 6.0 open cluster.
Angular size is 27 '.
Forms the sharp tip of a slightly extended isosceles triangle with Sirius and Wezen (the bright star in Canis Major's tail).

A bright, rich and large cluster with a surprise on its northern edge - the planetary nebula NGC 2438 (magnitude 10.8, 1.1' diameter). The cluster has a mass of around 450 Suns and a radius of 37.8 lightyears. It is 250 million years old. It forms a "double cluster" with M47, one degree away, but this is a line-of-sight coincidence and the two strikingly different clusters are unrelated.

Also visible:
(1) C54 (7.6 $6_{m}$ open cluster)
(2) M47 ( $5.2_{m}$ open cluster)
(3) NGC 2423 ( $6.7_{m}$ open cluster)
(4) NGC 2440 (10.5m planetary nebula)


RA: $116.15^{\circ} \mid 7 \mathrm{~h} 44.6^{\prime}-$ DEC: $-23.87^{\circ} \mid-23^{\circ} 51^{\prime}$
M93 (NGC 2447) is a magnitude 6.0 open cluster.
Angular size is $22^{\prime}$.
1.4 ${ }^{\circ}$ NW from mag.3.47 Azmidiske.


This rich open cluster was discovered by Charles Messier in 1781, and by Caroline Herschel in 1783. In larger telescopes the cluster can resemble a starfish. It is about 390 million years old, and is 3,380 light-years from Earth.

Also visible:
(1) NGC 2527 ( $6.5_{m}$ open cluster)
(2) NGC 2452 (12.5m planetary nebula)

## February: $\mathbf{- 4 5}^{\circ}$ South (1)



NGC 2546: page 155
NGC 3228: page 157
NGC 2808: page 159

C71: page 155
IC 2581: page 157

C79: page 156
C96: page 158

NGC 2547: page 156
1 Car: page 158

## February: -45 ${ }^{\circ}$ South (2)




RA: $123.1^{\circ} \mid 8 \mathrm{~h} 12.39^{\prime}-$ DEC: $-37.63^{\circ} \mid-37^{\circ} 37^{\prime}$
NGC 2546 is a magnitude 6.3 open cluster. Angular size is $41^{\prime}$.
2.9 ${ }^{\circ}$ NE from mag.2.27 Suhail Hadar.

Discovered from South Africa in 1751-1752 by Abbe Lacaille, this sparse open cluster is 3,300 light years from Earth.


Also visible:
(1) C71 (5.8m open cluster)
(2) RS Pup ( $6.52_{m}$ variable star)
(3) RT Pup ( $8.5_{m}$ carbon star)


RA: $118.08^{\circ} \mid 7 h 52.3^{\prime}-$ DEC: $-38.55^{\circ} \mid-38^{\circ} 32^{\prime}$

C71 (NGC 2477) is a magnitude 5.8 open cluster.
Angular size is 27 '.
$1.4^{\circ}$ SEE from mag. 3.72 c Pup.

This massive and dense open cluster resembles a globular cluster with its spherical shape and numerous stars. The center of Caldwell 71 however lacks the brightness of a true globular cluster. Given its size and great brightness, this is one of the best clusters in the sky.
©
Also visible:
(1) NGC 2451 ( $2.8_{m}$ open cluster)
(2) NGC 2546 ( $6.3_{m}$ open cluster)
(3) RS Pup ( $6.52_{m}$ variable star)
(4) RT Pup ( $8.5_{m}$ carbon star)


RA: $154.4^{\circ} \mid 10 \mathrm{~h} 17.6^{\prime}-$ DEC: $-46.42^{\circ} \mid-46^{\circ} 24^{\prime}$
C79 (NGC 3201) is a magnitude 6.7 globular cluster.
Angular size is 18 '.
5.7 ${ }^{\circ}$ NWW from mag. 2.84 mu . Vel.

A brilliant globular cluster and easy binocular target, a small telescope can resolve the member stars well into the relatively diffuse core. Caldwell 79 orbits in the opposite direction to the Milky Way's rotation, suggesting it may be a captured object. Interestingly, stars in this cluster become redder and cooler toward the core of the cluster, an inhomogenous characteristic only shared by Messier 4.


RA: $122.68^{\circ}\left|8 \mathrm{~h} 10.69^{\prime}-\mathrm{DEC}:-49.27^{\circ}\right|-49^{\circ} 15^{\prime}$

NGC 2547 is a magnitude 4.7 open cluster. Angular size is $20^{\prime}$.
$1.9^{\circ} \mathrm{S}$ from mag.1.92 Regor.

This young cluster (less than 30 million years old) possesses multiple stars with excess infrared radiation, strongly suggesting planet-forming discs around these stars. The cluster is only 1,200 lightyears distant.


RA: $155.45^{\circ} \mid 10 \mathrm{~h} 21.79^{\prime}-$ DEC: $-51.72^{\circ} \mid-51^{\circ} 42^{\prime}$

NGC 3228 is a magnitude 6.0 open cluster. Angular size is $18^{\prime}$.
$4.5^{\circ}$ SWW from mag. 2.84 mu . Vel.

O"
This small, bright open cluster is roughly 260 million years old. 53 possible members have been ascribed to this cluster.


RA: $156.85^{\circ} \mid 10 \mathrm{~h} 27.39^{\prime}-$ DEC: $-57.63^{\circ} \mid-57^{\circ} 37^{\prime}$
IC 2581 is a magnitude 4.3 open cluster. Angular size is $8^{\prime}$.
$3.6^{\circ}$ NWW from mag.3.88 u Car.

This bright open cluster has about 120 members, dominated by the A7 supergiant HD 90772 near its center.

Also visible:
(1) C92 (6.2m bright nebula)
(2) NGC 3114 (4.2m open cluster)
(3) S Car (4.5m variable star)


RA: $119.58^{\circ} \mid 7 \mathrm{~h} 58.3^{\prime}-$ DEC: $-60.87^{\circ} \mid-60^{\circ} 51^{\prime}$

C96 (Southern Beehive, NGC 2516) is a magnitude 3.8 open cluster. Angular size is $30^{\prime}$.
3.3 ${ }^{\circ}$ SWW from mag.1.74 71130

The Southern Beehive is a broad swarm of stars, easily seen by the naked eye in a dark location. 1,000 light-years from Earth, the cluster has over 100 stars.


RA: $146.31^{\circ} \mid 9 \mathrm{~h} 45.24^{\prime}-$ DEC: $-62.51^{\circ} \mid-62^{\circ} 29^{\prime}$

1 Car (HD 84810) is a magnitude 3.28 variable star.
Magnitude ranges from 4.18 to 3.28 ( $\Delta$ mag. 0.9 ) with a period of 35.53584 d .
$2.5^{\circ} \mathrm{N}$ from mag. 3.15 ups Car.

This bright Cepheid variable is best compared with the unaided eye to the other bright stars of Carina.

Also visible
(1) C90 ( $9.7_{m}$ planetary nebula)
(2) NGC 3114 (4.2m open cluster)
(3) NGC 2808 ( $6.2_{m}$ globular cluster)
(4) R Car (3.9m variable star)
(5) S Car (4.5m variable star)


RA: $138.01^{\circ} \mid 9$ h $12.04^{\prime}-$ DEC: $-64.86^{\circ} \mid-64^{\circ} 51^{\prime}$

NGC 2808 (NGC) is a magnitude 6.2 globular cluster.
Angular size is $14^{\prime}$.
$3.7^{\circ} \mathrm{W}$ from mag. 3.15 ups Car.

One of the largest globular clusters in the Milky Way, NGC 2808 has over a million stars. Studies suggest that the stars in NGC 2808 formed over a period of 200 million years in three distinct generations.

Also visible:
(1) IC 2501 (11.0 mplanetary nebula)
(2) 1 Car ( $3.28_{m}$ variable star)
(3) R Car (3.9 $9_{m}$ variable star)


RA: $116.35^{\circ}\left|7 \mathrm{~h} 45.39^{\prime}-\mathrm{DEC}:-37.97^{\circ}\right|-37^{\circ} 57^{\prime}$

NGC 2451 is a magnitude 2.8 open cluster. Angular size is $45^{\prime}$.

Draw a line from Sirius through Wezen, and extend it slightly more than an equal distance.

Two for one! This bright open cluster is actually two open clusters in the same line of sight - NGC 2451A at 600 light-years, and NGC 2451B at 1,200 lightyears.

Also visible:
(1) C71 (5.8m open cluster)
(2) RT Pup ( $8.5_{m}$ carbon star)


RA: $151.93^{\circ} \mid 10 \mathrm{~h} 7.7^{\prime}-$ DEC: $-40.43^{\circ} \mid-40^{\circ} 25^{\prime}$
C74 (Eight Burst Nebula, NGC 3132) is a magnitude 8.2 planetary nebula. Angular size is $0.8^{\prime}$.
$7.0^{\circ} \mathrm{E}$ from mag.3.64 psi Vel.

This fainter planetary nebula is only 0.4 light-years wide, so it is relatively young. The tenth magnitude star within the ring of nebulosity is not the source of the gas cloud nor of the ultraviolet light that causes the gas to fluoresce. This honor falls to the sixteenth magnitude companion star nearby.

Also visible:
(1) $\mathrm{X} \mathrm{Vel} \mathrm{( } 8.4_{m}$ carbon star)


## IC 2395

RA: $130.28^{\circ}\left|8 \mathrm{~h} 41.1^{\prime}-\mathrm{DEC}:-48.2^{\circ}\right|-48^{\circ} 11^{\prime}$

IC 2395 is a magnitude 4.6 open cluster. Angular size is $8^{\prime}$.
4.0 ${ }^{\circ}$ SWW from mag. 3.69 c Vel.

This open cluster has around 45 members and is under 18 million years old and 2,300 light-years distant. Its brightest member is magnitude 5.5.


RA: $131.23^{\circ} \mid 8 \mathrm{~h} 44.89^{\prime}-$ DEC: $-52.97^{\circ} \mid-52^{\circ} 57^{\prime}$
NGC 2669 is a magnitude 6.1 open cluster. Angular size is $12^{\prime}$.
$0.6^{\circ}$ E from mag.3.68 omi Vel.


RA: $140.35^{\circ} \mid 9 \mathrm{~h} 21.39^{\prime}-$ DEC: $-58.32^{\circ} \mid-58^{\circ} 18^{\prime}$
C90 (NGC 2867) is a magnitude 9.7 planetary nebula.
Angular size is $0.2^{\prime}$.
$1.1^{\circ}$ NNE from mag.2.25 Tureis.

When John Herschel discovered this object on April Fool's Day in 1834, he at first thought he had discovered a new planet, but it was "only" a planetary nebula. Being quite faint and small, Herschel's brief confusion is quite excusable.

Also visible:
(1) IC 2501 (11.0 planetary nebula)
(2) NGC 2899 (12.0 $0_{m}$ planetary nebula)
(3) 1 Car ( $3.28_{m}$ variable star)
(4) R Car (3.9 $9_{m}$ variable star)


RA: $150.68^{\circ} \mid 10 \mathrm{~h} 2.7^{\prime}-$ DEC: $-60.12^{\circ} \mid-60^{\circ} 6^{\prime}$

NGC 3114 is a magnitude 4.2 open cluster. Angular size is 35 '.
$2.1^{\circ}$ NWW from mag. 3.44 q Car.


A sparse open cluster almost lost in a very rich star field.

Also visible:
(1) C102 (1.9m open cluster)
(2) IC 2581 ( $4.3_{m}$ open cluster)
(3) 1 Car ( $3.28_{m}$ variable star)
(4) R Car (3.9 $9_{m}$ variable star)
(5) S Car (4.5m variable star)


RA: $152.34^{\circ} \mid 10 \mathrm{~h} 9.36^{\prime}-$ DEC: $-61.55^{\circ} \mid-61^{\circ} 32^{\prime}$

S Car (HD 88366) is a magnitude 4.5 variable star.
Magnitude ranges from 9.9 to 4.5 ( $\Delta$ mag. 5.4 ) with a period of 149 d .
$0.9^{\circ}$ SWW from mag. 3.44 q Car.

A Mira-type variable, although it is unusually hot for a star of this class and also has an unusually short period.
Also visible:
(1) C102 (1.9m open cluster)
(2) NGC 3114 (4.2m open cluster)
(3) IC 2581 ( $4.3_{m}$ open cluster)
(4) 1 Car ( $3.28_{m}$ variable star)
(5) R Car (3.9m variable star)


RA: $143.06^{\circ} \mid 9 \mathrm{~h} 32.24^{\prime}-$ DEC: $-62.79^{\circ} \mid-62^{\circ} 46^{\prime}$
R Car (HD 82901) is a magnitude 3.9 variable star.
Magnitude ranges from 10.5 to 3.9 ( $\Delta$ mag. 6.6 ) with a period of 309 d .
2.7 ${ }^{\circ}$ NW from mag. 3.15 ups Car.


A double star where the primary is a Mira-type variable, and the faint secondary (magnitude 11.3) is separated by only $2.1^{\prime \prime}$ on a position angle of $132^{\circ}$.

Also visible:
(1) C90 (9.7m planetary nebula)
(2) NGC 3114 (4.2m open cluster)
(3) NGC 2808 ( $6.2_{m}$ globular cluster)
(4) 1 Car ( $3.28_{m}$ variable star)
(5) S Car (4.5m variable star)

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## April: $\mathbf{4 5}^{\circ}$ North (1)



## April: $45^{\circ}$ North (2)




RA: $194.1^{\circ} \mid 12 \mathrm{~h} 56.39^{\prime}-$ DEC: $66.0^{\circ} \mid 66^{\circ} 0^{\prime}$
RY Dra (HD 112559) is a magnitude 6.0 carbon star.
Magnitude ranges from 8.2 to 6.0 ( $\Delta$ mag. 2.2) with a period of 170 d .

Forms an equilateral triangle with Mizar and Phecda.


An extremely red star with a B-V color index of 3.28, RY Dra is classified as a C-N3III carbon star. It is roughly 1,254 light-years from Earth.


Also visible:
(1) C3 (9.7m barred spiral galaxy)
(2) NGC 4605 (10.89 $9_{m}$ barred spiral galaxy)


RA: $179.13^{\circ} \mid 11 \mathrm{~h} 56.5^{\prime}-$ DEC: $57.87^{\circ} \mid 57^{\circ} 52^{\prime}$
Z UMa (HD 103681) is a magnitude 6.2 variable star. Magnitude ranges from 9.4 to 6.2 ( $\Delta$ mag. 3.2) with a period of 196d.
Position Phecda at the southern edge of the finder view; Z UMa is on the northern edge.

This pulsating variable star is roughly 1,059 lightyears from Earth.

Also visible:
(1) M40 (8.4m double star)
(2) M109 (9.8m barred spiral galaxy)


RA: $168.7^{\circ} \mid 11 \mathrm{~h} 14.79^{\prime}-$ DEC: $55.02^{\circ} \mid 55^{\circ} 1^{\prime}$
M97 (Owl Nebula, NGC 3587) is a magnitude 9.9
planetary nebula. Angular size is $3 \times 3.3^{\prime}$.
2.2 ${ }^{\circ}$ SE from mag.2.44 Merakh.

This nebula is composed of three concentric shells of gas. This innermost shell is actually barrel-shaped, and inclined at an angle to us, so the two dark "eyes" of this nebula are the open ends of the innermost shell.

Also visible:
(1) M108 (10.0 $0_{m}$ spiral galaxy)
(2) NGC 3631 (11.01 $1_{m}$ spiral galaxy)


RA: $202.5^{\circ} \mid 13 \mathrm{~h} 30.0^{\prime}-$ DEC: $47.18^{\circ} \mid 47^{\circ} 11^{\prime}$

M51 (Whirlpool Galaxy, NGC 5194) is a magnitude 8.4 spiral galaxy. Angular size is $11 \times 7$ '.

A quarter of the way along a line from Alkaid to Cor Caroli.

O:
M51 is a perfect example of an interacting galaxy pair. The main galaxy (NGC 5194) is a face-on spiral galaxy with well defined spiral arms and a bright core, while the companion galaxy is a bright but unclassifiable mass connected to one of the spiral arms by a dust lane. M51 was the first galaxy to be identified as being spiral in structure, by William Parsons in 1845.


RA: $191.28^{\circ}$ | $12 \mathrm{~h} 45.1^{\prime}-$ DEC: $45.4^{\circ} \mid 45^{\circ} 24^{\prime}$
Y Cvn (La Superba, HD 110914) is a magnitude 5.0 carbon star. Magnitude ranges from 6.4 to 5.0 ( $\Delta \mathrm{mag}$. 1.4) with a period of 158 d .

Travel from Cor Caroli to Chara in Canes Venaticorum, turn left, and travel an equal distance.

An extremely red carbon star (also known as La Superba) with a B-V color index of 3.0, this star is classified as a C-N5 carbon star. It is only 756 lightyears from Earth and its relatively bright minimum magnitude allows better observation of its strong color.
Also visible:
(1) C 21 ( $9.4_{m}$ irregular galaxy)
(2) M94 (8.2m spiral galaxy)
(3) M106 (8.4 spiral galaxy)
(4) NGC 4490 ( $10.22_{m}$ barred spiral galaxy)
(5) NGC 4618 ( $11.22_{m}$ SBd $)$


RA: $198.95^{\circ} \mid 13 \mathrm{~h} 15.79^{\prime}-$ DEC: $42.03^{\circ} \mid 42^{\circ} 2^{\prime}$
M63 (Sunflower Galaxy, NGC 5055) is a magnitude 8.6 spiral galaxy. Angular size is $10 \times 6$ '.

Located a third away along a line from Cor Caroli to Alkaid.

This flocculant spiral galaxy is so named because it does not have well-defined spiral arms. This galaxy was discovered in 1779 by Pierre Méchain, and lies 27 million light-years from Earth.

Also visible:
(1) M94 (8.2m spiral galaxy)


RA: $197.73^{\circ} \mid 13 \mathrm{~h} 10.89^{\prime}-$ DEC: $37.05^{\circ} \mid 37^{\circ} 3^{\prime}$
C29 (NGC 5005) is a magnitude 9.8 barred spiral galaxy. Angular size is $5 \times 2^{\prime}$.
3.1 ${ }^{\circ}$ SEE from mag.2.9 Cor Caroli.

A small and relatively faint galaxy, Caldwell 29 is nevertheless easier to spot even under less-than-ideal conditions because it has a very bright, prominent core. It is a LINER galaxy, meaning that gas in the core is ionized, probably by a supermassive black hole.

Also visible:
(1) M94 (8.2 $2_{m}$ spiral galaxy)
(2) NGC 5033 (10.75m spiral galaxy)


C32 (Whale Galaxy, NGC 4631) is a magnitude 9.3 spiral galaxy. Angular size is $15 \times 3$ '.

Draw a line from Mizar to Cor Caroli, and extend it by a third. C32 is slightly west of this point.

This elongated, interacting galaxy is one of the brighter Caldwell galaxies and is a viable target for smaller telescopes and poorer skies. It lies 30 million light-years from Earth.

Also visible:
(1) C35 (11.4m elliptical galaxy)
(2) C36 (9.8m spiral galaxy)
(3) NGC 4395 (10.64 ${ }_{m}$ Magellanic barred spiral galaxy)
(4) NGC 4414 (10.96 ${ }_{m}$ spiral galaxy)
(5) NGC 4656 (10.96 ${ }_{m}$ Magellanic barred spiral galaxy)


RA: $185.6^{\circ} \mid 12 \mathrm{~h} 22.39^{\prime}-$ DEC: $58.08^{\circ} \mid 58^{\circ} 5^{\prime}$
M40 (Winnecke 4, Win 4) is a magnitude 8.4 double star. Angular size is $0.8^{\prime}$.
1.4 ${ }^{\circ}$ NE from mag.3.44 Megrez.

This double star was included by Charles Messier in his catalog as he had a report of a nebula at that location but could find only this faint and unimpressive double star. The separations is very wide at 53", and the two components are very dim at magnitude 9.74 and 10.18 . Position angle is $77^{\circ}$. In a larger telescope, the components are light yellow.

Also visible:
(1) M109 (9.8m barred spiral galaxy)
(2) NGC 4605 (10.89m barred spiral galaxy)
(3) T UMa ( $6.6_{m}$ variable star)
(4) Z UMa ( $6.2_{m}$ variable star)


RA: $167.88^{\circ} \mid 11 \mathrm{~h} 11.5^{\prime}-$ DEC: $55.67^{\circ} \mid 55^{\circ} 40^{\prime}$
M108 (NGC 3556) is a magnitude 10.0 spiral galaxy. Angular size is 8 x 1 '.
$1.5^{\circ}$ SEE from mag.2.44 Merakh.

This nearly edge-on barred spiral galaxy is a member of the Ursa Major Cluster.

Also visible:
(1) M97 (9.9m planetary nebula)
(2) NGC 3587 (11.0m planetary nebula)
(3) NGC 3631 (11.01m spiral galaxy)


RA: $179.4^{\circ} \mid 11 \mathrm{~h} 57.6^{\prime}-$ DEC: $53.38^{\circ} \mid 53^{\circ} 23^{\prime}$
M109 (NGC 3992) is a magnitude 9.8 barred spiral galaxy. Angular size is $7 \times 4^{\prime}$.
$0.6^{\circ}$ SEE from mag.2.54 Phecda.

This fine barred spiral galaxy is the leading galaxy of the M109 Group, which contains up to 50 galaxies.

Also visible:
(1) NGC 3953 (10.84m barred spiral galaxy)
(2) NGC 4088 (11.15m barred spiral galaxy)
(3) NGC 3893 (11.16 $\sigma_{m}$ spiral galaxy)
(4) Z UMa ( $6.2_{m}$ variable star)


RA: $184.73^{\circ} \mid 12 \mathrm{~h} 18.89^{\prime}-$ DEC: $47.32^{\circ} \mid 47^{\circ} 19^{\prime}$
M106 (NGC 4258) is a magnitude 8.4 spiral galaxy.
Angular size is $19 x 8$ '.
Draw a line from Dubhe through Phecda, and extend it an equal length to reach M106.

M106 resembles the much nearer Andromeda Galaxy in size and absolute brightness. The black hole at its core is nearly ten times the mass of the the Milky Way's supermassive black hole.

Also visible:
(1) C21 (9.4m irregular galaxy)
(2) NGC 4051 (10.83 ${ }_{m}$ barred spiral galaxy)
(3) NGC 3938 (10.9m spiral galaxy)
(4) NGC 4088 (11.15m barred spiral galaxy)
(5) Y Cvn (5.0 carbon star)


RA: $187.05^{\circ} \mid 12 \mathrm{~h} 28.2^{\prime}-$ DEC: $44.1^{\circ} \mid 44^{\circ} 6^{\prime}$
C21 (NCG 4449, NGC 4449) is a magnitude 9.4
irregular galaxy. Angular size is $6 \times 44^{\prime}$.
7.7 ${ }^{\circ}$ NW from mag.2.9 Cor Caroli.

An irregular dwarf galaxy, Caldwell 21 is a nearby starburst galaxy, speckled with star-forming regions. The galaxy has a central bar making it similar to the Large Magellanic Cloud (LMC), but with 15 billion stars it is only half the size. Caldwell 21 compensates by having twice the star formation rate of the LMC.

Also visible:
(1) M94 (8.2 spiral galaxy)
(2) M106 (8.4 $4_{m}$ spiral galaxy)
(3) NGC 4490 (10.22 barred spiral galaxy)
(4) NGC 4051 ( $10.83_{m}$ barred spiral galaxy)
(5) $\mathrm{Y} \operatorname{Cvn}\left(5.0_{m}\right.$ carbon star)


RA: $192.73^{\circ} \mid 12 \mathrm{~h} 50.89^{\prime}-$ DEC: $41.13^{\circ} \mid 41^{\circ} 8^{\prime}$

M94 (NGC 4736) is a magnitude 8.2 spiral galaxy.
Angular size is 7x3'.
2.9 ${ }^{\circ}$ NNW from mag.2.9 Cor Caroli.

This spiral galaxy is distinguished by the ring of star formation around the nucleus, and its curious lack of dark matter.

Also visible:
(1) C 21 (9.4 irregular galaxy)
(2) C 29 ( $9.8_{m}$ barred spiral galaxy)
(3) M63 (8.6 $6_{m}$ spiral galaxy)
(4) NGC 4490 (10.22 ${ }_{m}$ barred spiral galaxy)
(5) Y Cvn ( $5.0_{m}$ carbon star)


RA: $184.38^{\circ} \mid 12 \mathrm{~h} 17.5^{\prime}-$ DEC: $37.82^{\circ} \mid 37^{\circ} 49^{\prime}$
C26 (NGC 4244) is a magnitude 10.6 spiral galaxy.
Angular size is $17 \times 2^{\prime}$.
7.6 ${ }^{\circ} \mathrm{W}$ from mag.2.9 Cor Caroli.

O
An edge-on spiral galaxy, the Silver Needle is distinguished by lacking both a prominent core and a central line of dust clouds. Like many galaxies, the Silver Needle suffers badly under light-polluted skies.

Also visible:
(1) NGC 4490 (10.22 $2_{m}$ barred spiral galaxy)
(2) NGC 4214 ( $10.24_{m}$ irregular galaxy)
(3) NGC 4395 (10.64m Magellanic barred spiral galaxy)
(4) NGC 4618 (11.22m $S B d)$

## April: $\mathbf{1 5}^{\circ}$ North (1)



C35: page 177
M98: page 179
M86: page 181
M60: page 183

C38: page 177
M88: page 179
M89: page 181
M61: page 183

M53: page 178
M90: page 180
M96: page 182

C40: page 178
M105: page 180
M58: page 182

## April: $\mathbf{1 5}^{\circ}$ North (2)



C36: page 184
M91: page 186
M84: page 188
M49: page 190

M64: page 184
M99: page 186
M87: page 188
SS Vir: page 190

M85: page 185
M65: page 187
M95: page 189

M100: page 185
M66: page 187
M59: page 189


## C35

RA: $195.03^{\circ} \mid 13 \mathrm{~h} 0.1^{\prime}-$ DEC: $27.98^{\circ} \mid 27^{\circ} 59^{\prime}$

C35 (NGC 4889) is a magnitude 11.4 elliptical galaxy.
Angular size is $3 \times 2^{\prime}$.
$10.3^{\circ} \mathrm{S}$ from mag.2.9 Cor Caroli.

A massive galaxy over twice as large as the Milky Way, Caldwell 35 contains the largest known black hole, with a mass of 21 billion Suns. Over 300 million light-years from Earth, this object is extremely difficult for smaller telescopes or for light-polluted observing sites.

Also visible:
(1) C32 (9.3 $3_{m}$ spiral galaxy)
(2) NGC 4725 (10.11m barred spiral galaxy)
(3) NGC 4656 (10.96 ${ }_{m}$ Magellanic barred spiral galaxy)


RA: $189.08^{\circ} \mid 12 \mathrm{~h} 36.29^{\prime}-$ DEC: $25.98^{\circ} \mid 25^{\circ} 59^{\prime}$

C38 (Needle Galaxy, NGC 4565) is a magnitude 9.6 barred spiral galaxy. Angular size is $16 \times 3$ '.

South and slightly west of Cor Caroli; it is as distant from Cor Caroli as Alkaid at the end of the Dipper's handle.

O
Although a faint galaxy, the Needle Galaxy can be seen with smaller telescopes as the edge-on orientation of the galaxy enhances the surface brightness. In contrast a face-one galaxy has lower surface brightness as is is viewed though its thinnest dimension. Larger telescopes will reveal the needle to be bisected lengthwise by a dark dust lane.

Also visible:
(1) C36 (9.8m spiral galaxy)
(2) M64 (8.5m spiral galaxy)
(3) NGC 4725 (10.11m barred spiral galaxy)
(4) NGC 4494 ( $10.71_{m}$ elliptical galaxy)
(5) NGC 4278 (11.09 elliptical galaxy)


RA: $198.23^{\circ} \mid 13 \mathrm{~h} 12.89^{\prime}-$ DEC: $18.17^{\circ} \mid 18^{\circ} 10^{\prime}$

M53 (NGC 5024) is a magnitude 7.6 globular cluster.
Angular size is 13 '.
Two-fifths of the way from Arcturus to Denebola.

A very distant globular cluster (60,000 light-years from Earth). The member stars of this cluster are dimmed greatly by distance and require a larger telescope to resolve them. It is quite old for a globular cluster at 12.67 billion years, and its stars are largely first-generation stars (they formed from gas that has not been recycled through other stars). It is interacting with NGC 5053 (mag. 9.96, one degree SE), to which it is connected by a tidal stream.

Also visible:
(1) M64 (8.5m spiral galaxy)
(2) NGC 5053 ( $9.47_{m}$ globular cluster)


C40 (NGC 3626) is a magnitude 10.9 barred spiral galaxy. Angular size is $3 \times 2$ '.
2.8 ${ }^{\circ}$ SSE from mag.2.58 Zosma.

This is one of the smallest Caldwell objects (as viewed from Earth) and it is also not particularly bright. It has a relatively prominent core, making it more accessible to smaller telescopes.

Also visible:
(1) NGC 3628 (10.28m barred spiral galaxy)
(2) NGC 3607 (10.82m lenticular galaxy)


RA: $183.48^{\circ} \mid 12 \mathrm{~h} 13.89^{\prime}-$ DEC: $14.92^{\circ} \mid 14^{\circ} 55^{\prime}$
M98 (NGC 4192) is a magnitude 10.1 spiral galaxy.
Angular size is $10 \times 3$ '.
Slightly north of a point one quarter of the way from Denebola in Leo to Vindemiatrix in Virgo.

This intermediate spiral galaxy is a member of the Virgo Cluster, and is approaching us at $140 \mathrm{~km} / \mathrm{s}$.

Also visible:
(1) M84 (9.1 $1_{m}$ lenticular galaxy)
(2) M85 (9.1m lenticular galaxy)
(3) M86 (8.9m lenticular galaxy)
(4) M87 (8.6 elliptical galaxy)
(5) R Com (7.1m variable star)


RA: $188.03^{\circ} \mid 12 \mathrm{~h} 32.1^{\prime}-$ DEC: $14.43^{\circ} \mid 14^{\circ} 26^{\prime}$

M88 (NGC 4501) is a magnitude 9.6 spiral galaxy.
Angular size is 7 x 4 '.
Slightly north of a point halfway between the splendidly named Vindemiatrix in Virgo and Denebola in Leo.

This spiral galaxy is a member of the Virgo Cluster of galaxies. It is on course for the center of the cluster, and may make a close pass by Messier 87 in 200-300 million years.

Also visible:
(1) M60 (8.8m elliptical galaxy)
(2) M84 (9.1m lenticular galaxy)
(3) M85 (9.1m lenticular galaxy)
(4) M86 (8.9m lenticular galaxy)
(5) M87 (8.6m elliptical galaxy)


RA: $189.2^{\circ} \mid 12 \mathrm{~h} 36.79^{\prime}-$ DEC: $13.17^{\circ} \mid 13^{\circ} 10^{\prime}$
M90 (NGC 4569) is a magnitude 9.5 spiral galaxy.
Angular size is $10 \times 5{ }^{\prime}$.
Two thirds of the way from Denebola in Leo to Vindemiatrix in Virgo.

This spiral galaxy, like many other members of the massive Virgo Cluster, has lost much of its gas due to ram pressure due to the speed with which the galaxy is moving the the Virgo Cluster's intergalactic medium. The galaxy has a 260,000 light-year tail of ionized gas lost due to this ram pressure.

Also visible:
(1) M60 (8.8 $8_{m}$ elliptical galaxy)
(2) M84 (9.1m lenticular galaxy)
(3) M86 (8.9m lenticular galaxy)
(4) M87 (8.6 $\sigma_{m}$ elliptical galaxy)
(5) M100 (9.3 $3_{m}$ spiral galaxy)


RA: $161.95^{\circ} \mid 10 \mathrm{~h} 47.79^{\prime}-$ DEC: $12.58^{\circ} \mid 12^{\circ} 35^{\prime}$

M105 (NGC 3379) is a magnitude 9.3 elliptical galaxy. Angular size is $2^{\prime}$.

Equally distant from Zosma, Algeiba and Regulus.

This large elliptical galaxy is part of the M96 Group. There is some evidence of current star formation within this galaxy, suggesting that elliptical galaxies are capable of star formation at a reduced rate.

Also visible:
(1) M95 (9.7 $7_{m}$ barred spiral galaxy)
(2) M96 (9.2m spiral galaxy)
(3) NGC 3384 ( $10.85_{m}$ elliptical galaxy)
(4) NGC 3489 (11.12m lenticular galaxy)


RA: $186.55^{\circ} \mid 12 \mathrm{~h} 26.2^{\prime}-$ DEC: $12.95^{\circ} \mid 12^{\circ} 57^{\prime}$
M86 (NGC 4406) is a magnitude 8.9 lenticular galaxy. Angular size is $8 \times 6{ }^{\prime}$.

Draw a line from Chertan through Denebola in Leo and extend it an equal distance.

This galaxy has the highest blue shift of all Messier galaxies, as it is approaching the Milky Way at 244 $\mathrm{km} / \mathrm{s}$. It is moving so fast through the intergalactic medium of the Virgo Cluster, that ram pressure is stripping the galaxy of its gas, leaving behind a wake of ionized matter glowing in the X-ray spectrum.

Also visible:
(1) M60 (8.8 $8_{m}$ elliptical galaxy)
(2) M84 (9.1m lenticular galaxy)
(3) M87 (8.6 elliptical galaxy)
(4) M90 (9.5m spiral galaxy)
(5) M100 (9.3 $3_{m}$ spiral galaxy)


RA: $188.93^{\circ} \mid 12 \mathrm{~h} 35.7^{\prime}-$ DEC: $12.55^{\circ} \mid 12^{\circ} 33^{\prime}$

M89 (NGC 4552) is a magnitude 9.8 elliptical galaxy.
Angular size is $4^{\prime}$.
Two thirds of the way from Denebola in Leo to
Vindemiatrix in Virgo.
This elliptical galaxy may in fact be almost perfectly spherical. It has ten times as many globular clusters as the Milky Way.

Also visible:
(1) M49 (8.4m elliptical galaxy)
(2) M60 (8.8 $8_{m}$ elliptical galaxy)
(3) M84 (9.1 $1_{m}$ lenticular galaxy)
(4) M86 (8.9m lenticular galaxy)
(5) M87 (8.6m elliptical galaxy)


RA: $161.7^{\circ} \mid 10 \mathrm{~h} 46.79^{\prime}-$ DEC: $11.82^{\circ} \mid 11^{\circ} 49^{\prime}$
M96 (NGC 3368) is a magnitude 9.2 spiral galaxy.
Angular size is $6 \times 4$ '.
4.2 ${ }^{\circ}$ NE from mag.3.85 rho Leo.

The lead member of the eponymous M96 Group, this dim galaxy has very weak spiral arms and a relatively bright core. A unique feature of this galaxy is the core not quite at the center of the galaxy, a distortion doubtlessly caused by interaction with its neighboring galaxies.

Also visible:
(1) M95 (9.7 $7_{m}$ barred spiral galaxy)
(2) M105 (9.3m elliptical galaxy)
(3) NGC 3384 ( $10.85_{m}$ elliptical galaxy)
(4) NGC 3489 (11.12m lenticular galaxy)


RA: $189.43^{\circ} \mid 12 \mathrm{~h} 37.7^{\prime}-$ DEC: $11.82^{\circ} \mid 11^{\circ} 49^{\prime}$

M58 (NGC 4579) is a magnitude 9.7 barred spiral galaxy. Angular size is $6 \times 5{ }^{\prime}$.

Draw a line from Chertan in Leo through Denebola, and extend it one-and-a-half times.

The most distant Messier object at 62 million lightyears, and one of the brightest Virgo galaxies. This barred spiral has a bright core, which is all that is visible in smaller telescopes. Larger telescopes have a chance of seeing some structure in the spiral disk on dark nights.

Also visible:
(1) M49 (8.4m elliptical galaxy)
(2) M60 (8.8m elliptical galaxy)
(3) M84 (9.1m lenticular galaxy)
(4) M86 (8.9m lenticular galaxy)
(5) M87 (8.6 elliptical galaxy)


RA: $190.93^{\circ} \mid 12 \mathrm{~h} 43.7^{\prime}-$ DEC: $11.55^{\circ} \mid 11^{\circ} 33^{\prime}$
M60 (NGC 4649) is a magnitude 8.8 elliptical galaxy.
Angular size is 7 x 6 '.
Draw a line from Chertan in Leo through Denebola, and extend it one-and-a-half times.
$\because$
M60 has one of the largest supermassive black holes at its heart - approximately 4.5 billion solar masses. NGC 4647, an intermediate spiral galaxy (magnitude 11.94), overlaps with M60.

Also visible:
(1) M49 (8.4m elliptical galaxy)
(2) M84 (9.1m lenticular galaxy)
(3) M86 (8.9m lenticular galaxy)
(4) M87 (8.6 $\sigma_{m}$ elliptical galaxy)
(5) R Vir (6.1 $1_{m}$ variable star)


## M61

RA: $185.48^{\circ}\left|12 \mathrm{~h} 21.89^{\prime}-\mathrm{DEC}: 4.47^{\circ}\right| 4^{\circ} 28^{\prime}$

M61 (NGC 4303) is a magnitude 9.7 spiral galaxy.
Angular size is $6^{\prime}$.
Draw a line from Zosma in Leo through Denebola, and extend it one-and-a-half times. M61 is slightly south of this point.

Home to seven recent supernovae, this starburst galaxy is 55 million light-years from Earth. Small telescopes will struggle with this target.

Also visible:
(1) M49 (8.4m elliptical galaxy)
(2) NGC 4365 (10.52 $2_{m}$ elliptical galaxy)
(3) NGC 4535 (10.59m spiral galaxy)
(4) R Vir ( $6.1_{m}$ variable star)
(5) SS Vir ( $6.0_{m}$ carbon star)


RA: $189.0^{\circ} \mid 12 \mathrm{~h} 36.0^{\prime}-$ DEC: $27.97^{\circ} \mid 27^{\circ} 58^{\prime}$
C36 (Koi Fish Galaxy, NGC 4559) is a magnitude 9.8 spiral galaxy. Angular size is $10 \times 44^{\prime}$.
11.1 ${ }^{\circ}$ SSW from mag.2.9 Cor Caroli.

An almost edge-on spiral galaxy with a irregular texture. It is the site of an active and exceedingly rare Luminous Blue Variable star AT2016blu, which has erupted with supernova-like intensity in 2014, 2016, 2017, 2018, 2019, 2020 and 2021.

Also visible:
(1) C32 (9.3 $3_{m}$ spiral galaxy)
(2) C38 (9.6 barred spiral galaxy)
(3) NGC 4725 ( $10.11_{m}$ barred spiral galaxy)
(4) NGC 4494 (10.71 $1_{m}$ elliptical galaxy)
(5) NGC 4656 ( $10.96_{m}$ Magellanic barred spiral galaxy)

RA: $194.18^{\circ} \mid 12 \mathrm{~h} 56.7^{\prime}-$ DEC: $21.68^{\circ} \mid 21^{\circ} 41^{\prime}$

M64 (Black Eye Galaxy, NGC 4826) is a magnitude 8.5 spiral galaxy. Angular size is $9 \times 5$ '.

Slightly less than halfway from Muphrid to Zosma.

Distinguished by a dark lane of dust partially occluding the bright nucleus, Messier 64 is a Seyfert galaxy. Part of the gas in this galaxy orbits the core in a retrograde direction, possibly indicating a galaxy merger in the past, or continuing absorption of gas from the intergalactic medium.

Also visible:
(1) C38 (9.6 $6_{m}$ barred spiral galaxy)
(2) M53 (7.6m globular cluster)
(3) NGC 5053 ( $9.47_{m}$ globular cluster)
(4) NGC 4725 (10.11m barred spiral galaxy)


RA: $186.38^{\circ}$ | $12 \mathrm{~h} 25.5^{\prime}-$ DEC: $18.2^{\circ} \mid 18^{\circ} 12^{\prime}$
M85 (NGC 4382) is a magnitude 9.1 lenticular galaxy. Angular size is 7x5'.

Draw a line from Regulus through Chertan in Leo, and extend it an equal distance.

This featureless lenticular galaxy is the northernmost member of the Virgo Cluster, with a mass in the region of 100 billion solar masses.

Also visible:
(1) M88 ( $9.6_{m}$ spiral galaxy)
(2) M91 (10.2 $2_{m}$ barred spiral galaxy)
(3) M98 (10.1m spiral galaxy)
(4) M99 (9.9m spiral galaxy)
(5) M100 ( $9.3_{m}$ spiral galaxy)


RA: $185.75^{\circ} \mid 12 \mathrm{~h} 23.0^{\prime}-$ DEC: $15.83^{\circ} \mid 15^{\circ} 50^{\prime}$

M100 (NGC 4321) is a magnitude 9.3 spiral galaxy.
Angular size is 7x6'.
Slightly north of a point one third of the way from Denebola in Leo to Vindemiatrix in Virgo.

This grand-design spiral galaxy is 56 million lightyears from Earth. While M100 is very photogenic, larger telescopes and good skies are needed to see any of its beautiful details.

Also visible:
(1) M84 (9.1 $1_{m}$ lenticular galaxy)
(2) M85 (9.1 1 lenticular galaxy)
(3) M86 ( $8.9_{m}$ lenticular galaxy)
(4) M87 (8.6 $6_{m}$ elliptical galaxy)
(5) $\mathrm{R} \operatorname{Com}$ (7.1m variable star)


RA: $188.88^{\circ} \mid 12 \mathrm{~h} 35.5^{\prime}-$ DEC: $14.5^{\circ} \mid 14^{\circ} 30^{\prime}$
M91 (NGC 4548) is a magnitude 10.2 barred spiral galaxy. Angular size is $5^{\prime}$.

Slightly north of a point halfway between Vindemiatrix in Virgo and Denebola in Leo.

One of the faintest Messier objects, this object was "lost" until 1969 as Charles Messier made a mistake in logging the location (he measured the location relative to Messier 89, but logged the directions as being relative to Messier 58 by mistake). Like many galaxies in the Virgo Cluster, M91 is deficient in interstellar gas.

Also visible:
(1) M60 ( $8.8_{m}$ elliptical galaxy)
(2) M84 (9.1m lenticular galaxy)
(3) M85 (9.1m lenticular galaxy)
(4) M86 (8.9m lenticular galaxy)
(5) M87 (8.6 elliptical galaxy)


RA: $184.73^{\circ} \mid$ 12h $18.89^{\prime}-$ DEC: $14.43^{\circ} \mid 14^{\circ} 26^{\prime}$

M99 (NGC 4254) is a magnitude 9.9 spiral galaxy.
Angular size is $5^{\prime}$.
One third of the way from Denebola in Leo to
Vindemiatrix in Virgo.
This grand-design spiral galaxy is 55 million lightyears from Earth.

Also visible:
(1) M84 (9.1 $1_{m}$ lenticular galaxy)
(2) M85 (9.1m lenticular galaxy)
(3) M86 (8.9m lenticular galaxy)
(4) M87 (8.6 $\sigma_{m}$ elliptical galaxy)
(5) R Com (7.1 $1_{m}$ variable star)


RA: $169.73^{\circ} \mid 11 \mathrm{~h} 18.89^{\prime}-$ DEC: $13.08^{\circ} \mid 13^{\circ} 5^{\prime}$
M65 (NGC 3623) is a magnitude 9.3 spiral galaxy.
Angular size is $8 \times 1.5^{\prime}$.
2.6 ${ }^{\circ}$ SSE from mag.3.41 Chertan.

Part of the famous Leo Triplet of galaxies (with Messier 66 and NGC 3623). The galaxy is 35 million light-years from Earth. There is very little starforming activity in this galaxy and very little gas, leading to relatively featureless photographs.

Also visible:
(1) M66 ( $8.9_{m}$ spiral galaxy)
(2) NGC 3628 (10.28m barred spiral galaxy)
(3) NGC 3489 (11.12m lenticular galaxy)


RA: $170.05^{\circ} \mid 11 \mathrm{~h} 20.2^{\prime}-$ DEC: $12.98^{\circ} \mid 12^{\circ} 59^{\prime}$

M66 (NGC 3627) is a magnitude 8.9 spiral galaxy.
Angular size is $8 \times 2.5^{\prime}$.
2.8 ${ }^{\circ}$ SSE from mag.3.41 Chertan.

Part of the famous Leo Triplet of galaxies (with Messier 65 and NGC 3623). In his discovery notes, Messier described this galaxy as "very long and very faint". Messier 65 is only 20 ' away, making an attractive pair in dark conditions.

Also visible:
(1) M65 (9.3 $3_{m}$ spiral galaxy)
(2) NGC 3628 (10.28m barred spiral galaxy)


RA: $186.2^{\circ}$ | 12h $25.1^{\prime}-$ DEC: $12.88^{\circ} \mid 12^{\circ} 53^{\prime}$
M84 (NGC 4374) is a magnitude 9.1 lenticular galaxy. Angular size is $5^{\prime}$.

Draw a line from Chertan through Denebola in Leo and extend it an equal distance.

M84 is a giant elliptical or lenticular galaxy distinguished by a dark dust lane and two jets of matter shooting from its central black hole (which is an impressive 1.5 billion solar masses). The galaxy has nearly 2000 globular clusters, which is actually less than expected for an elliptical galaxy of this size. M84 has some young star clusters and has had at least two supernovae since 1957, so it clearly still has continuing star-forming activity.

Also visible:
(1) M60 ( $8.8_{m}$ elliptical galaxy)
(2) M86 ( $8.9_{m}$ lenticular galaxy)
(3) M87 ( $8.6_{m}$ elliptical galaxy)
(4) M90 ( $9.5_{m}$ spiral galaxy)
(5) M100 (9.3.3 spiral galaxy)


RA: $187.7^{\circ} \mid 12 \mathrm{~h} 30.79^{\prime}-$ DEC: $12.4^{\circ} \mid 12^{\circ} 24^{\prime}$

M87 (NGC 4486) is a magnitude 8.6 elliptical galaxy.
Angular size is 7'.
Draw a line from Chertan through Denebola in Leo and extend it an equal distance.

This supergiant elliptical galaxy comprising trillions of stars is one of the most massive galaxies in the sky. It has 15,000 globular clusters. This galaxy is famous for its iconic plasma jet that originate from its central black hole (the corresponding jet on the far side of the galaxy is hidden by relativistic effects). This black hole is believed to be 6,5 billion solar masses, amongst the largest known, and also the first to be directly imaged (April 2019).

Also visible:
(1) M49 (8.4m elliptical galaxy)
(2) M60 (8.8m elliptical galaxy)
(3) M84 (9.1m lenticular galaxy)
(4) M86 ( $8.9_{m}$ lenticular galaxy)
(5) M100 (9.3 $3_{m}$ spiral galaxy)


RA: $161.0^{\circ} \mid 10 \mathrm{~h} 44.0^{\prime}-$ DEC: $11.7^{\circ} \mid 11^{\circ} 42^{\prime}$
M95 (NGC 3351) is a magnitude 9.7 barred spiral galaxy. Angular size is $4 \times 3$ '.
$3.6^{\circ}$ NE from mag.3.85 rho Leo.

Like M94, M95 has a ring-shaped circumnuclear star-forming region, which include several large star clusters that may are massive enough to be potentially become globular clusters. This faint galaxy is a member of the M96 Group, which includes a third Messier object, M105.

Also visible:
(1) M96 (9.2 $2_{m}$ spiral galaxy)
(2) M105 (9.3m elliptical galaxy)
(3) NGC 3384 ( $10.85_{m}$ elliptical galaxy)
(4) NGC 3489 (11.12m lenticular galaxy)


RA: $190.5^{\circ} \mid 12 \mathrm{~h} 42.0^{\prime}-$ DEC: $11.65^{\circ} \mid 11^{\circ} 39^{\prime}$
M59 (NGC 4621) is a magnitude 9.6 elliptical galaxy.
Angular size is $5 \times 4$.
Draw a line from Chertan in Leo through Denebola, and extend it one-and-a-half times.

M59 and M60 were spotted and cataloged by Charles Messier as he searched for the newly discovered Bode's Comet in Spring 1779. As an elliptical galaxy, M59 has no structure and while a small telescope can reveal the elongated shape, a larger telescope will not show more.

Also visible:
(1) M49 (8.4m elliptical galaxy)
(2) M60 ( $8.8_{m}$ elliptical galaxy)
(3) M86 ( $8.9_{m}$ lenticular galaxy)
(4) M87 (8.6 $6_{m}$ elliptical galaxy)
(5) R Vir (6.1m variable star)


RA: $187.45^{\circ} \mid 12 \mathrm{~h} 29.79^{\prime}-$ DEC: $8.0^{\circ} \mid 8^{\circ} 0^{\prime}$
M49 (NGC 4472) is a magnitude 8.4 elliptical galaxy.
Angular size is $9 \times 7.5^{\prime}$.
Draw a line from Zosma through Denebola in Leo, and extend it one-and-a-half times.

This giant elliptical galaxy is believed to harbor a supermassive black hole 565 million times more massive than the Sun (to put it differently, 150 times more massive than the black hole in the center of the Milky Way). M49 has nearly 6,000 globular clusters, significantly more than the 200 that can be found in our own galaxy. M49 was the first Virgo galaxy to be discovered.

Also visible:
(1) M59 ( $9.6_{m}$ elliptical galaxy)
(2) M60 ( $8.8_{m}$ elliptical galaxy)
(3) M61 ( $9.7_{m}$ spiral galaxy)
(4) M87 (8.6 $6_{m}$ elliptical galaxy)
(5) R Vir (6.1m variable star)


RA: $186.33^{\circ} \mid$ 12h $25.29^{\prime}-$ DEC: $0.7^{\circ} \mid 0^{\circ} 42^{\prime}$

SS Vir (HD 108105) is a magnitude 6.0 carbon star.
Magnitude ranges from 9.6 to $6.0(\Delta$ mag. 3.6) with a period of 355 d .
4.6 ${ }^{\circ}$ NWW from mag. 3.68 gam Vir.

A deeply red (B-V color index of 3.5) carbon star classified as C-N4.5. Located 2,280 light-years from Earth.

Also visible:
(1) M61 (9.7 $7_{m}$ spiral galaxy)
(2) NGC 4636 (10.43 ${ }_{m}$ elliptical galaxy)
(3) NGC 4517 (11.1m spiral galaxy)
(4) NGC 4536 (11.16 $\sigma_{m}$ barred spiral galaxy)

## April: $\mathbf{- 1 5}^{\circ}$ South



C52: page 192
C61: page 194
M68: page 196

M104: page 192
R Crv: page 194

U Hya: page 193
V Hya: page 195

C60: page 193
R Hya: page 195


RA: $192.15^{\circ} \mid 12 \mathrm{~h} 48.6^{\prime}-$ DEC: $-5.8^{\circ} \mid-5^{\circ} 47^{\prime}$
C52 (NGC 4697) is a magnitude 9.3 elliptical galaxy.
Angular size is $6 \times 3$ '.
4.6 ${ }^{\circ}$ SSE from mag. 3.68 gam Vir.
$\because$
The bright core of this galaxy can be seen in lightpolluted skies with a small telescope. Larger telescopes only reveal a larger and somewhat elongated smudge.

Also visible:
(1) NGC 4776 (11.5m planetary nebula)
(2) NGC 4699 (10.41m barred spiral galaxy)
(3) NGC 4753 (10.85m irregular galaxy)


## M104

RA: $190.0^{\circ} \mid 12 \mathrm{~h} 40.0^{\prime}-$ DEC: $-11.62^{\circ} \mid-11^{\circ} 36^{\prime}$

M104 (Sombrero Galaxy, NGC 4594) is a magnitude 8.0 spiral galaxy. Angular size is $9 x 4$ '.
5.4 ${ }^{\circ}$ NNE from mag.3.11 Algorab.

The Sombrero Galaxy is so named because of its prominent outer dust ring and large central bulge give it a hat shape. This is a brighter galaxy with high contrast, so it is a good one for smaller telescopes.

Also visible:
(1) NGC 4776 (11.5m planetary nebula)
(2) NGC 4699 (10.41m barred spiral galaxy)


RA: $159.4^{\circ}$ | 10h $37.6^{\prime}-\mathrm{DEC}:-13.4^{\circ} \mid-13^{\circ} 23^{\prime}$
U Hya (HD 92055) is a magnitude 4.7 carbon star.
Magnitude ranges from 6.2 to 4.7 ( $\Delta$ mag. 1.5).
4.0 $0^{\circ} \mathrm{NW}$ from mag. 3.32 nu . Hya.

This carbon star has a B-V color index of 2.81 and spectral class of C-N5. SAO 150629 (HD 91120) is 1.5 degrees west and at magnitude 5.55 marks a midpoint in the brightness range of this star.

Also visible:
(1) U Hya (7.0 $0_{m}$ variable star)


RA: $180.48^{\circ} \mid 12 \mathrm{~h} 1.89^{\prime}-$ DEC: $-18.87^{\circ} \mid-18^{\circ} 51^{\prime}$
C60 (Antennae Galaxies, NGC 4038) is a magnitude
11.3 spiral galaxy. Angular size is $2.6 \times 1.8^{\prime}$.
3.5 ${ }^{\circ}$ SWW from mag.2.78 Dzhanakh.

An interesting pair of colliding galaxies riven by colossal star forming regions.

Also visible:
(1) R Crv (6.7m variable star)


RA: $180.48^{\circ} \mid 12 \mathrm{~h} 1.89^{\prime}-$ DEC: $-18.88^{\circ} \mid-18^{\circ} 52^{\prime}$
C61 (Antennae Galaxies, NGC 4039) is a magnitude
13.0 spiral galaxy. Angular size is $3.2 \times 2.2^{\prime}$.
3.5 ${ }^{\circ}$ SWW from mag.2.78 Dzhanakh.


An interesting pair of colliding galaxies riven by colossal star forming regions.

Also visible:
(1) R Crv ( $6.7_{m}$ variable star)


RA: $184.91^{\circ} \mid 12 \mathrm{~h} 19.63^{\prime}-$ DEC: $-19.26^{\circ} \mid-19^{\circ} 14^{\prime}$
R Crv (HD 107199) is a magnitude 6.7 variable star. Magnitude ranges from 14.4 to 6.7 ( $\Delta$ mag. 7.7) with a period of 317 d .
$1.9^{\circ} \mathrm{SSE}$ from mag.2.78 Dzhanakh.

Deeply red (B-V color index: 3.7), R Corvi has a very deep minimum brightness and only spends about a quarter of its cycle brighter than magnitude 9.0. R Corvi is a Mira-type variable.

Also visible:
(1) C60 (11.3 $3_{m}$ spiral galaxy)
(2) C61 (13.0 $\mathrm{m}_{\mathrm{m}}$ spiral galaxy)
(3) NGC 4361 (10.5m planetary nebula)


RA: $162.9^{\circ}$ | $10 \mathrm{~h} 51.6^{\prime}-\mathrm{DEC}:-21.3^{\circ} \mid-21^{\circ} 17^{\prime}$
V Hya (SAO 179278) is a magnitude 6.5 carbon star. Magnitude ranges from 12.0 to 6.5 ( $\Delta \mathrm{mag} .5 .5$ ) with a period of 533 d .
$5.1^{\circ} \mathrm{S}$ from mag. 3.32 nu . Hya.


An utterly unique star. A carbon star with a B-V color index of 2.59, it varies in magnitude from 6 to 8 over 530 days, but every 6160 days it drops down to magnitude 12. The star is believed to be transitioning to a planetary nebula. It is interacting with an unseen companion to eject jets of matter periodically.


RA: $202.43^{\circ}$ | 13h $29.71^{\prime}-$ DEC: $-23.28^{\circ} \mid-23^{\circ} 16^{\prime}$
R Hya (HD 117287) is a magnitude 3.5 variable star. Magnitude ranges from 10.9 to 3.5 ( $\Delta \mathrm{mag}$. 7.4) with a period of 389 d .
$2.4^{\circ} \mathrm{E}$ from mag.3.33 gam Hya.

A very bright Mira-type variable that spends almost half its life brighter than magnitude 6. It is thought to be between 500 to 1,000 million years old and possibly twice the mass of the Sun.
Also visible:
(1) NGC 5068 (10.7m barred spiral galaxy)


RA: $189.88^{\circ} \mid 12 \mathrm{~h} 39.5^{\prime}-$ DEC: $-26.75^{\circ} \mid-26^{\circ} 44^{\prime}$
M68 (NGC 4590) is a magnitude 7.8 globular cluster. Angular size is 12 .
$3.5^{\circ}$ SSE from mag.2.84 Kraz.
$\%$
William Herschel described this globular cluster as follows in 1780: "a beautiful cluster of stars, extremely rich, and so compressed that most of the stars are blended together". Messier 68 appears to be undergoing core-collapse, whereby momentum is transferred from the core of the cluster to its fringes, resulting in a very dense core.

## April: $\mathbf{- 4 5}^{\circ}$ South (1)



U Ant: page 199
C80: page 199
C91: page 200
C92: page 200
R Cru: page 201
C98: page 201
C102: page 202

## April: -45 ${ }^{\circ}$ South (2)



C77: page 202
C100: page 204

C83: page 203
C99: page 205

C94: page 203
C97: page 204


RA: $158.8^{\circ}$ | 10h $35.2^{\prime}-$ DEC: $-39.6^{\circ} \mid-39^{\circ} 35^{\prime}$
U Ant (HD 91793) is a magnitude 5.7 carbon star.
Magnitude ranges from 8.6 to 5.7 ( $\Delta$ mag. 2.9) with a period of 170 d .
$10.0^{\circ} \mathrm{NNW}$ from mag. 2.84 mu . Vel.

A deeply red carbon star (B-V color index: 3.03).
It emits 8,000 times more energy than the Sun, but mostly as infrared light; in the visible spectrum the star is only 500 times brighter than the Sun.


RA: $201.7^{\circ} \mid 13 \mathrm{~h} 26.79^{\prime}-$ DEC: $-47.48^{\circ} \mid-47^{\circ} 28^{\prime}$

C80 (Omega Centauri, NGC 5139) is a magnitude 3.6 globular cluster. Angular size is 36 '.
$4.8^{\circ} \mathrm{W}$ from mag. 3.06 zet Cen.

By far the brightest globular cluster in Earth's sky, Omega Centauri may well be the remnant core of a captured galaxy. It has a mass of four million Suns, seven time more than the most famous northern hemisphere globular, Messier 13.

Also visible:
(1) C 77 ( $7.0_{m}$ peculiar galaxy)
(2) C83 (9.5m spiral galaxy)
(3) C84 (7.6 $6_{m}$ globular cluster)
(4) NGC 5286 ( $7.34_{m}$ globular cluster)
(5) NGC 4976 (11.04m elliptical galaxy)


## C91

RA: $166.6^{\circ} \mid 11 \mathrm{~h} 6.39^{\prime}-$ DEC: $-58.67^{\circ} \mid-58^{\circ} 39^{\prime}$

C91 (Wishing Well Cluster, NGC 3532) is a magnitude 3.0 open cluster. Angular size is $55^{\prime}$.
$1.6^{\circ} \mathrm{E}$ from mag. 3.88 u Car.

This brilliant, naked-eye cluster is twice the size of the full moon. It was the first target observed by the Hubble Space Telescope (20 May 1990). It is the subject of the photograph taken at the greatest distance from the Earth, by New Horizons on December 5, 2017.

Also visible:
(1) C92 ( $6.2_{m}$ bright nebula)
(2) C97 (5.3 $3_{m}$ open cluster)
(3) C100 (4.5m open cluster)
(4) NGC 3699 (11.0 m planetary nebula)


RA: $160.95^{\circ} \mid 10 \mathrm{~h} 43.79^{\prime}-$ DEC: $-59.87^{\circ} \mid-59^{\circ} 51^{\prime}$
C92 (Eta Carinae Nebula, NGC 3372) is a magnitude 6.2 bright nebula. Angular size is $120 \times 120^{\prime}$.
$1.5^{\circ} \mathrm{SW}$ from mag. 3.88 u Car.

This naked-eye emission nebula is home to Eta Carina, one of the most massive stars known with a mass of 100 Suns and a luminosity five million times greater than the Sun. Given its huge extent, this object is best viewed in small wide-field telescopes or binoculars.

Also visible:
(1) C91 (3.0 open cluster)
(2) C102 (1.9m open cluster)
(3) IC 2581 (4.3 $3_{m}$ open cluster)
(4) NGC 3211 (11.5m planetary nebula)
(5) S Car (4.5m variable star)


RA: $185.91^{\circ} \mid 12 \mathrm{~h} 23.62^{\prime}-$ DEC: $-61.63^{\circ} \mid-61^{\circ} 37^{\prime}$

R Cru (HD 107805) is a magnitude 6.4 variable star. Magnitude ranges from 7.23 to 6.4 ( $\Delta$ mag. 0.8 ) with a period of 5.82575 d .
$1.2^{\circ}$ SSE from mag.3.57 eps Cru.


A Cepheid variable, this yellow supergiant pulsates with a radius varying between 42 and 47 times the radius of the Sun. It is approximately 1,600 lightyears distant.


Also visible:
(1) C94 (4.2m open cluster)
(2) C98 ( $6.9_{m}$ open cluster)
(3) C99 (20.0 dark nebula)
(4) NGC 3918 ( $8.0_{m}$ Type)


RA: $190.58^{\circ} \mid 12 \mathrm{~h} 42.29^{\prime}-$ DEC: $-62.97^{\circ} \mid-62^{\circ} 57^{\prime}$
C98 (Coalsack Cluster, NGC 4609) is a magnitude 6.9 open cluster. Angular size is $5^{\prime}$.

## $1.7^{\circ} \mathrm{E}$ from mag.2.09 alf02 Cru.

In comparison to other southern hemisphere Caldwell clusters that are prominent naked-eye objects, Caldwell 98 is somewhat more pedestrian, only visible with the aid of binoculars at least, and being relatively sparse.

Also visible:
(1) C94 (4.2m open cluster)
(2) C99 (20.0 $0_{m}$ dark nebula)
(3) R Cru ( $6.4_{m}$ variable star)


## C102

RA: $160.8^{\circ} \mid 10 \mathrm{~h} 43.2^{\prime}-$ DEC: $-64.4^{\circ} \mid-64^{\circ} 23^{\prime}$
C102 (Theta Car Cluster, IC 2602) is a magnitude 1.9 open cluster. Angular size is $50^{\prime}$.
$0.0^{\circ}$ SEE from mag. 3.03 the Car.

The Theta Carinae cluster (also known as the Southern Pleiades) is the third brightest open cluster in Earth's skies, outshone only by the Hyades and Pleiades.

Also visible:
(1) C92 (6.2m bright nebula)
(2) IC 2553 (11.5m planetary nebula)
(3) NGC 3211 (11.5m planetary nebula)
(4) IC 2621 (11.5m planetary nebula)
(5) $\mathrm{S} \mathrm{Car}\left(4.5_{m}\right.$ variable star)


## C77

RA: $201.38^{\circ} \mid 13 \mathrm{~h} 25.5^{\prime}-$ DEC: $-43.02^{\circ} \mid-43^{\circ} 0^{\prime}$
C77 (Centaurus A, NGC 5128) is a magnitude 7.0 peculiar galaxy. Angular size is $18 \times 14^{\prime}$.
$3.7^{\circ}$ SSW from mag. 3.96 d Cen.

This peculiar galaxy is the fifth-brightest galaxy in the sky, and is relatively close at $10-16$ million lightyears. It has an active galactic nucleus that strongly emits in the radio spectrum (hence its radio source name Centaurus A). It is believed that the galaxy is a merger between an elliptical galaxy and a smaller spiral galaxy.

Also visible:
(1) C80 (3.6 m $_{m}$ globular cluster)


RA: $196.35^{\circ} \mid 13 \mathrm{~h} 5.39^{\prime}-$ DEC: $-49.47^{\circ} \mid-49^{\circ} 27^{\prime}$
C83 (NGC 4945) is a magnitude 9.5 spiral galaxy.
Angular size is $20 \times 4$ '.
$3.9^{\circ}$ E from mag. 2.38 gam Cen.

This edge-on barred spiral galaxy is home to a water megamaser (a vast structure emitting coherent microwave radiation, similar to a laser).


Also visible:
(1) C80 (3.6 $6_{m}$ globular cluster)
(2) Ruprecht 106 (10.9m globular cluster)
(3) NGC 4976 (11.04m elliptical galaxy)


C94 (Jewel Box, NGC 4755) is a magnitude 4.2 open cluster. Angular size is 10'.
$0.9^{\circ} \mathrm{SE}$ from mag. 1.5 Becrux.

This bright naked-eye cluster can be resolved into its stars with binoculars. It is quite distant at 6,500 lightyears, and consists of 100 stars.

Also visible:
(1) C98 ( $6.9_{m}$ open cluster)
(2) C99 (20.0 $0_{m}$ dark nebula)
(3) R Cru ( $6.4_{m}$ variable star)


RA: $174.03^{\circ} \mid 11 \mathrm{~h} 36.1^{\prime}-$ DEC: $-61.62^{\circ} \mid-61^{\circ} 36^{\prime}$
C97 (Pearl Cluster, NGC 3766) is a magnitude 5.3 open cluster. Angular size is 12 '.
$1.4^{\circ} \mathrm{N}$ from mag.3.34 lam Cen.

Located in the Carina molecular cloud, this open cluster contains over 100 stars within a 10 light-year radius.


Also visible:
(1) C91 (3.0 $0_{m}$ open cluster)
(2) C100 (4.5m open cluster)
(3) NGC 3918 ( $8.0_{m}$ Type)
(4) NGC 3699 (11.0m planetary nebula)
(5) IC 2621 (11.5m planetary nebula)


RA: $174.15^{\circ} \mid 11 \mathrm{~h} 36.6^{\prime}-$ DEC: $-63.03^{\circ} \mid-63^{\circ} 1^{\prime}$

C100 (Lambda Centauri Nebula, IC 2944) is a magnitude 4.5 open cluster. Angular size is $15^{\prime}$.
$0.0^{\circ} \mathrm{E}$ from mag.3.34 lam Cen.

A combination of a very bright nebula and an open cluster with many prominent members, Caldwell 100 is a good binocular target that benefits greatly from dark skies.

Also visible:
(1) C91 (3.0 $0_{m}$ open cluster)
(2) C97 (5.3 open cluster)
(3) NGC 3699 (11.0m planetary nebula)
(4) IC 2621 (11.5m planetary nebula)


RA: $193.25^{\circ} \mid 12 \mathrm{~h} 53.0^{\prime}-$ DEC: $-63.0^{\circ} \mid-63^{\circ} 0^{\prime}$
C99 (Coalsack Nebula, TGU H1867) is a magnitude 20.0 dark nebula. Angular size is $400 \times 300^{\prime}$.
$2.9^{\circ} \mathrm{E}$ from mag.2.09 alf02 Cru.
$\circ$
This dark cloud of dust is only 600 light-years from Earth, obscuring the light of everything behind it. As it is relatively close, few stars shine in front of the Coalsack, making this dark patch stand out prominently under dark skies, either with binoculars or the unaided naked eye.

Also visible:
(1) C94 (4.2m open cluster)
(2) NGC 5189 (10.0 m planetary nebula)
(3) IC 4191 (11.5m planetary nebula)
(4) R Cru ( $6.4_{m}$ variable star)

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## May: $\mathbf{4 5}^{\circ}$ North




RA: $226.63^{\circ}\left|15 \mathrm{~h} 6.5^{\prime}-\mathrm{DEC}: 55.77^{\circ}\right| 55^{\circ} 46^{\prime}$
M102 (NGC 5866) is a magnitude 9.9 lenticular galaxy. Angular size is $5 \times 2^{\prime}$.

Two fifths of the way from Pherkad in Ursa Minor to Izar in Bootes.

Two fifths of the way from Pherkad in Ursa Minor to Izar in Bootes.

One of the more doubtful Messiers, M102 was disowned by its discoverer Pierre Méchain as a duplicate discovery of M101. However the modern consensus is that M102 was a genuine discovery corresponding to the edge-on spiral galaxy NGC 5866.


Also visible:
(1) NGC 5907 (11.12m spiral galaxy)


RA: $210.8^{\circ} \mid 14 \mathrm{~h} 3.2^{\prime}-$ DEC: $54.35^{\circ} \mid 54^{\circ} 21^{\prime}$

M101 (Pinwheel Galaxy, NGC 5457) is a magnitude 7.9 spiral galaxy. Angular size is $22^{\prime}$.

Forms an equilateral triangle with triple star Mizar and Alkaid.

Pierre Méchain described this galaxy as "very obscure and pretty large, $6^{\prime}$ to 7 ' in diameter, between the left hand of Bootes and the tail of the great Bear." In 1784 William Herschel could make out the mottled appearance of the object, and even thought he would be able to resolve the stars of it is composed. William Parsons was the first to detect the spiral shape of the galaxy in the latter half of the 19th century. To repeat Parson's feat requires dark skies and a substantial telescope.

Also visible:
(1) NGC 5585 (11.2m spiral galaxy)


RA: $240.66^{\circ} \mid 16 \mathrm{~h} 2.65^{\prime}-$ DEC: $47.24^{\circ} \mid 47^{\circ} 14^{\prime}$

X Her (HD 144205) is a magnitude 6.0 variable star.
Magnitude ranges from 7.0 to $6.0(\Delta$ mag. 1.0) with a period of 95.0d.
3.0 ${ }^{\circ}$ NWW from mag. 3.91 tau Her.

This red giant radiates 680 times the energy of the
Sun and is 449 light-years distant. Just less than a degree to the northeast, SAO 45895 (HD 145082) is magnitude 6.65 and marks a rough midpoint in the range of X Her.


## g Her

RA: $247.16^{\circ} \mid 16 \mathrm{~h} 28.64^{\prime}-$ DEC: $41.88^{\circ} \mid 41^{\circ} 53^{\prime}$
$\mathrm{g} \operatorname{Her}$ (HD 148783) is a magnitude 4.3 variable star.
Magnitude ranges from 6.3 to 4.3 ( $\Delta$ mag. 2.0) with a period of 89.2d.
$4.0^{\circ}$ NW from mag.3.61 eta Her.

This pulsating variable star is 354 light years from Earth. 35 Herculis is one degree to the northeast, and at magnitude 4.2 marks the bright end of $g$ Herculis' range.
Also visible:
(1) NGC 6058 (13.0 planetary nebula)
(2) $\mathrm{W} \mathrm{CrB}\left(7.8_{m}\right.$ variable star)


RA: $237.38^{\circ} \mid 15 \mathrm{~h} 49.52^{\prime}-$ DEC: $39.57^{\circ} \mid 39^{\circ} 34^{\prime}$
V CrB (HD 141826) is a magnitude 6.9 variable star.
Magnitude ranges from 12.6 to 6.9 ( $\Delta$ mag. 5.7 ) with a period of 358 d .
8.7 ${ }^{\circ}$ SW from mag.3.91 tau Her.

This carbon star is extremely red with a B-V color index of 3.5. The spectral class is C6,2eMS3. This star is 2,740 light-years from Earth.

Also visible:
(1) NGC 6058 (13.0 m planetary nebula)


RA: $243.85^{\circ} \mid 16 \mathrm{~h} 15.4^{\prime}-$ DEC: $37.8^{\circ} \mid 37^{\circ} 48^{\prime}$
W CrB (HD 146560) is a magnitude 7.8 variable star.
Magnitude ranges from 14.3 to 7.8 ( $\Delta$ mag. 6.5 ) with a period of 238 d .

Halfway between Chi and Zeta Herculis. Also, one finder circle west of M13.

This star is roughly 7,800 light-years from Earth.

Also visible:
(1) NGC 6058 (13.0 $0_{m}$ planetary nebula)
(2) $\mathrm{g} \operatorname{Her}\left(4.3_{m}\right.$ variable star)


RA: $229.55^{\circ} \mid 15 \mathrm{~h} 18.18^{\prime}-$ DEC: $31.65^{\circ} \mid 31^{\circ} 39^{\prime}$
U CrB (HD 136175) is a magnitude 7.66 variable star. Magnitude ranges from 8.79 to 7.66 ( $\Delta$ mag. 1.1) with a period of 3.45220 d .

Two degrees south and very slightly east of Thiba (Delta Bootis).

This spectroscopic pair, consisting of a bluish B6V main-sequence star and a yellowish F8III-IV subgiant, is an Algol-type eclipsing binary.

Also visible:
(1) S CrB ( $5.8_{m}$ variable star)


RA: $230.35^{\circ} \mid 15 \mathrm{~h} 21.39^{\prime}-$ DEC: $31.37^{\circ} \mid 31^{\circ} 22^{\prime}$

S CrB (HD 136753) is a magnitude 5.8 variable star. Magnitude ranges from 14.1 to 5.8 ( $\Delta$ mag. 8.3) with a period of 360 d .
2.3 ${ }^{\circ}$ SSE from mag.3.54 Printseps.

The Mira-type variable has a period of almost exactly a year. It is roughly 1,300 light-years distant and has a radius 308 times greater than the Sun, despite only being 1.34 times more massive.

Also visible:
(1) U CrB ( $7.66_{m}$ variable star)

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## May: $\mathbf{1 5}^{\circ}$ North



M3: page 214
U Her: page 216
M5: page 218

R CrB: page 214
R Ser: page 216

R Boo: page 215
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T CrB: page 215
C45: page 217


M3 (NGC 5272) is a magnitude 6.2 globular cluster.
Angular size is $16.2^{\prime}$.
Just short of halfway from Arcturus to Cor Caroli.

A naked-eye object under perfect conditions, M3 was discovered by Charles Messier in 1764. It is an atypical globular cluster as it is quite far from the galactic center and the galactic plane. it also has an unusually large number of variable stars.


R CrB (HD 141527) is a magnitude 5.71 variable star. Magnitude ranges from 14.8 to 5.71 ( $\Delta$ mag. 9.1).

Halfway between Izar in Bootes and Zeta Herculis, slightly closer to Zeta.

The prototype of R Cor Bor variables, this yellow supergiant is usually 6th magnitude, but fades by several magnitudes at irregular intervals. It is believed the star is emitting carbon that condenses around the star, temporarily obscuring it from sight.

Also visible:
(1) T CrB ( $2.0_{m}$ variable star)
(2) T CrB ( $2.3_{m}$ carbon star)


RA: $219.3^{\circ} \mid 14 \mathrm{~h} 37.19^{\prime}-$ DEC: $26.74^{\circ} \mid 26^{\circ} 44^{\prime}$

R Boo (HD 128609) is a magnitude 6.2 variable star.
Magnitude ranges from 13.1 to 6.2 ( $\Delta$ mag. 6.9 ) with a period of 223 d .
$1.7^{\circ} \mathrm{W}$ from mag.2.7 eps Boo.

This Mira-type variable is nearly 6,000 times brighter than the Sun, and has a radius 475 times larger. For comparison, half a degree east is tiny SAO 83440 (HD 128609), which is magnitude 8.55 .


RA: $239.88^{\circ} \mid 15 \mathrm{~h} 59.5^{\prime}-$ DEC: $25.92^{\circ} \mid 25^{\circ} 55^{\prime}$

T CrB (HD 143454) is a magnitude 2.0 variable star. Magnitude ranges from 10.8 to 2.0 ( $\Delta$ mag. 8.8) with a period of 80 y .
$3.7^{\circ} \mathrm{E}$ from mag. 3.93 gam CrB .

This cataclysmic variable is a recurring nova, occasionally surging to magnitude 2.0. Two outbursts are recorded, in 1866 and 1946.

Also visible:
(1) R CrB (5.71 $1_{m}$ variable star)


RA: $246.45^{\circ} \mid 16 \mathrm{~h} 25.79^{\prime}-$ DEC: $18.89^{\circ} \mid 18^{\circ} 54^{\prime}$
U Her (HD 148206) is a magnitude 6.4 variable star. Magnitude ranges from 13.4 to 6.4 ( $\Delta$ mag. 7.0 ) with a period of 406 d .
$0.9^{\circ}$ SEE from mag. 3.79 gam Her.


This pulsating variable star is orange-red and is roughly 1,860 light-years from Earth. Compare this star to magnitude 7.0 SAO 102152 (HD 148128), half a degree north.


Also visible:
(1) GCl 38 (14.74m globular cluster)


RA: $237.67^{\circ} \mid 15 \mathrm{~h} 50.69^{\prime}-$ DEC: $15.13^{\circ} \mid 15^{\circ} 8^{\prime}$

R Ser (HD 141850) is a magnitude 5.16 variable star. Magnitude ranges from 14.4 to 5.16 ( $\Delta$ mag. 9.2) with a period of 356 d .
$1.1^{\circ}$ SEE from mag. 3.74 bet Ser.

This Mira-type star was first shown to be variable by Karl Ludwig Harding in 1826. It spends only one third of its cycle brighter than magnitude 8.0. It is roughly 930 light-years distant.


RA: $230.41^{\circ} \mid 15 \mathrm{~h} 21.65^{\prime}-$ DEC: $14.31^{\circ} \mid 14^{\circ} 19^{\prime}$

S Ser (HD 136695) is a magnitude 7.0 variable star.
Magnitude ranges from 14.1 to 7.0 ( $\Delta$ mag. 7.1) with a period of 372 d .
$6.0^{\circ} \mathrm{W}$ from mag. 3.74 bet Ser.


RA: $204.38^{\circ} \mid 13 \mathrm{~h} 37.5^{\prime}-$ DEC: $8.88^{\circ} \mid 8^{\circ} 53^{\prime}$
C45 (NGC 5248) is a magnitude 10.2 spiral galaxy. Angular size is $6 \times 4$ '.

Slightly north-west of a point midway between Spica and Arcturus.

The arms of this galaxy are rich with young star clusters, while deep photographic images reveal delicate dust lanes woven deep into the core. This galaxy is a challenging target for small telescopes.

Also visible:
(1) NGC 5363 (11.05m irregular galaxy)
(2) NGC 5364 (11.17m barred spiral galaxy)


RA: $229.65^{\circ} \mid 15 \mathrm{~h} 18.6^{\prime}-$ DEC: $2.08^{\circ} \mid 2^{\circ} 5^{\prime}$
M5 (NGC 5904) is a magnitude 5.6 globular cluster.
Angular size is $17.4^{\prime}$.
7.7 ${ }^{\circ}$ SW from mag.2.75 Cor Serpentis.

This globular cluster was discovered by Gottfried Kirch in 1702. Its member stars are as bright as magnitude 10.6 , so they can be resolved by smaller telescopes. William Herschel counted roughly 200 stars in 1791. Over 200 years later, how many can you count?

Also visible:
(1) Palomar 5 ( $11.75_{m}$ globular cluster)
(2) NGC 5846 (11.05m elliptical galaxy)

## May: $-15^{\circ}$ South




## S Vir

RA: $203.25^{\circ} \mid 13 \mathrm{~h} 33.0^{\prime}-$ DEC: $-7.19^{\circ} \mid-7^{\circ} 11^{\prime}$

S Vir (HD 117833) is a magnitude 6.3 variable star.
Magnitude ranges from 13.2 to 6.3 ( $\Delta$ mag. 6.9 ) with a period of 375 d .
4.4 ${ }^{\circ}$ NNE from mag.1.21 Spica.

This moderately red pulsating variable star is roughly 2,380 light-years from Earth.


RA: $246.68^{\circ}\left|16 \mathrm{~h} 26.72^{\prime}-\mathrm{DEC}:-12.43^{\circ}\right|-12^{\circ} 25^{\prime}$

V Oph (HD 148182) is a magnitude 7.3 variable star. Magnitude ranges from 11.6 to 7.3 ( $\Delta$ mag. 4.3 ) with a period of 297 d .
3.1 ${ }^{\circ}$ SW from mag.2.7 Khan.

This deeply red carbon star (B-V color index: 3.41)
has spectral type C-N4. Messier 107 is two degrees to the southwest.

Also visible:
(1) M107 (7.9m globular cluster)


RA: $244.25^{\circ} \mid 16$ h $17.0^{\prime}-$ DEC: $-22.98^{\circ} \mid-22^{\circ} 58^{\prime}$
M80 (NGC 6093) is a magnitude 7.3 globular cluster.
Angular size is $10^{\prime}$.
2.7 ${ }^{\circ}$ NNW from mag.3.08 Alniyat.


This is one of the densest globular clusters in our galaxy, and is particularly rich in "blue stragglers", old stars that have been made hotter and bluer through collisions or close interactions with other cluster members.

Also visible:
(1) M4 (5.6 $6_{m}$ globular cluster)
(2) NGC 6144 ( $9.01_{m}$ globular cluster)
(3) Antares $\left(0.9_{m}\right.$ carbon star)


RA: $219.9^{\circ} \mid 14 \mathrm{~h} 39.6^{\prime}-$ DEC: $-26.53^{\circ} \mid-26^{\circ} 31^{\prime}$

C66 (NGC 5694) is a magnitude 10.2 globular cluster.
Angular size is $3.6^{\prime}$.
5.6 ${ }^{\circ}$ SWW from mag. 3.41 sig Lib .

One of the most distant globular clusters associated with our galaxy, Caldwell 66 may be on a hyperbolic path that will cause it to escape the gravity well of the Milky Way.


RA: $245.9^{\circ} \mid 16 \mathrm{~h} 23.6^{\prime}-$ DEC: $-26.53^{\circ} \mid-26^{\circ} 31^{\prime}$
M4 (NGC 6121) is a magnitude 5.6 globular cluster.
Angular size is $26.3^{\prime}$.
$1.0^{\circ} \mathrm{SSE}$ from mag.3.08 Alniyat.

This globular cluster was the first to be resolved into individual stars (the brightest being magnitude 10.8). It is also bright enough to be seen by the naked eye under ideal conditions.

Also visible:
(1) M80 (7.3m globular cluster)
(2) NGC 6144 (9.01m globular cluster)
(3) ESO452-SC11 (12.0 ${ }_{m}$ globular cluster)
(4) Antares ( $0.9_{m}$ carbon star)


RA: $247.35^{\circ} \mid 16 \mathrm{~h} 29.39^{\prime}-$ DEC: $-26.43^{\circ} \mid-26^{\circ} 25^{\prime}$
A fiercely brilliant and intensely red primary with a very close, bright, blue or green companion.
0.0 $0^{\circ}$ NWW from mag.1.22 148479

Antares means the "rival of Mars" as this star's color competes with that of the Red Planet. This is a gravitationally bound binary system about 550 lightyears from Earth. Antares is fairly red (B-V color index: 1.86).

Also visible:
(1) M4 (5.6 globular cluster)
(2) M80 (7.3 ${ }^{m}$ globular cluster)
(3) NGC 6144 ( $9.01_{m}$ globular cluster)
(4) ESO452-SC11 (12.0m globular cluster)

# May: $-45^{\circ}$ South (1) 



M83: page 225
C88: page 227

C75: page 225
C89: page 227

R Nor: page 226
NGC 5617: page 228

T Nor: page 226
NGC 5316: page 228

May: -45 ${ }^{\circ}$ South (2)



RA: $204.25^{\circ}\left|13 \mathrm{~h} 37.0^{\prime}-\mathrm{DEC}:-29.87^{\circ}\right|-29^{\circ} 51^{\prime}$

M83 (Southern Pinwheel, NGC 5236) is a magnitude 7.6 spiral galaxy. Angular size is $11 \times 10$ '.
$7.2^{\circ}$ SWW from mag. 3.48 pi. Hya.

This southerly barred spiral galaxy was discovered by Nicolas-Louis de Lacaille in 1752 from a site in South Africa. Only 15 million light-years away, it is one of the best examples of a barred spiral galaxy in our skies.

Also visible:
(1) NGC 5253 ( $10.87_{m}$ irregular galaxy)
(2) T Cen ( $5.5_{m}$ variable star)


C75 (NGC 6124) is a magnitude 5.8 open cluster.
Angular size is 29 '.
$5.4^{\circ}$ SEE from mag.3.61 eta Lup.

A bright and rich open cluster, with many blue and red members.

Also visible:
(1) NGC 6153 (11.0m planetary nebula)
(2) NGC 6139 ( $8.99_{m}$ globular cluster)


RA: $233.99^{\circ} \mid 15 \mathrm{~h} 35.95^{\prime}-$ DEC: $-49.51^{\circ} \mid-49^{\circ} 29^{\prime}$
R Nor (HD 138743) is a magnitude 6.5 variable star. Magnitude ranges from 13.9 to 6.5 ( $\Delta$ mag. 7.4) with a period of 508 d .

Position Zet Lup on the south-western edge of the finder circle; R Nor is on the north-eastern edge.

A long period variable located 1,896 light-years from Earth with a stately period of over 500 days.

Also visible:
(1) NGC 5882 (10.0 m planetary nebula)
(2) NGC 5927 (8.01m globular cluster)
(3) NGC 5946 ( $9.61_{m}$ globular cluster)
(4) ESO224-8 (14.0 $0_{m}$ globular cluster)


RA: $236.02^{\circ} \mid 15 \mathrm{~h} 44.06^{\prime}-$ DEC: $-54.99^{\circ} \mid-54^{\circ} 58^{\prime}$
T Nor (HD 140041) is a magnitude 6.2 variable star. Magnitude ranges from 13.6 to $6.2(\Delta$ mag. 7.4$)$ with a period of 241 d .

Just over one finder circle south east of Zet Lup.

A long-period pulsating variable, the star is 3,623 light-years from Earth.

Also visible:
(1) C89 ( $5.4_{m}$ open cluster)
(2) NGC 6067 (5.6 $6_{m}$ open cluster)
(3) NGC 5927 ( $8.01_{m}$ globular cluster)
(4) NGC 5946 ( $9.61_{m}$ globular cluster)


RA: $226.43^{\circ} \mid 15 \mathrm{~h} 5.7^{\prime}-$ DEC: $-55.6^{\circ} \mid-55^{\circ} 35^{\prime}$
C88 (NGC 5823) is a magnitude 7.9 open cluster.
Angular size is 10 .
$3.6^{\circ}$ SSW from mag. 3.5 zet Lup.


Discovered by James Dunlop in 1826.

Also visible:
(1) NGC 5662 ( $5.5_{m}$ open cluster)
(2) NGC 5946 ( $9.61_{m}$ globular cluster)


RA: $244.73^{\circ} \mid 16 \mathrm{~h} 18.89^{\prime}-$ DEC: $-57.9^{\circ} \mid-57^{\circ} 53^{\prime}$

C89 (S Norma Cluster, NGC 6087) is a magnitude 5.4 open cluster. Angular size is $12^{\prime}$.
4.1 ${ }^{\circ}$ NWW from mag.3.68 eta Ara.

A naked-eye open cluster composed of about 40 stars.
Very similar to the nearby Caldwell 95.
( Also visible:
(1) C95 (5.1 $1_{m}$ open cluster)
(2) NGC 6067 (5.6m open cluster)
(3) NGC 5979 (12.0m planetary nebula)
(4) NGC 6221 (10.66 $m_{m}$ barred spiral galaxy)


RA: $217.45^{\circ} \mid 14 \mathrm{~h} 29.79^{\prime}-$ DEC: $-60.72^{\circ} \mid-60^{\circ} 42^{\prime}$

NGC 5617 is a magnitude 6.3 open cluster. Angular size is 10 '.
$1.3^{\circ} \mathrm{W}$ from mag.1.7 alf Cen.


At low powers, this bright open cluster is in the same field as the famous double star Alpha Centauri, which lies one degree to the southeast. The cluster is estimated to be 82 million years old.

Also visible:
(1) NGC 5316 ( $6.0_{m}$ open cluster)
(2) NGC 5662 ( $5.5_{m}$ open cluster)
(3) R Cen ( $5.3_{m}$ variable star)


RA: $208.48^{\circ} \mid 13 \mathrm{~h} 53.89^{\prime}-$ DEC: $-61.87^{\circ} \mid-61^{\circ} 51^{\prime}$

NGC 5316 is a magnitude 6.0 open cluster. Angular size is $14^{\prime}$.
$1.9^{\circ}$ SW from mag.0.86 Khadar.

This bright and rich cluster shares the finder circle with Beta Centauri, and open cluster NGC 5281 (magnitude 5.9) and IC 4291 (magnitude 9.7). Its brightest star is magnitude 9.4 , and the cluster is around 100 million years old.

Also visible:
(1) NGC 5281 (5.9m open cluster)
(2) NGC 5617 ( $6.3_{m}$ open cluster)
(3) NGC 5189 (10.0 m planetary nebula)
(4) R Cen ( $5.3_{m}$ variable star)


RA: $205.44^{\circ} \mid 13 \mathrm{~h} 41.75^{\prime}-$ DEC: $-33.6^{\circ} \mid-33^{\circ} 35^{\prime}$
T Cen (HD 119090) is a magnitude 5.5 variable star.
Magnitude ranges from 9.0 to 5.5 ( $\Delta$ mag. 3.5) with a period of 90.44 d .
$5.3^{\circ} \mathrm{NE}$ from mag.2.91 iot Cen.

Also visible:
(1) M83 (7.6 $m_{m}$ spiral galaxy)
(2) NGC 5102 (10.35m lenticular galaxy)
(3) NGC 5253 ( $10.87_{m}$ irregular galaxy)


RA: $211.9^{\circ}\left|14 \mathrm{~h} 7.6^{\prime}-\mathrm{DEC}:-48.32^{\circ}\right|-48^{\circ} 18^{\prime}$

NGC 5460 is a magnitude 5.6 open cluster. Angular size is $25^{\prime}$.
$2.2^{\circ}$ SEE from mag.3.06 zet Cen.

A bright, loose cluster with an estimated age of 160 million years. It is roughly 2,350 light-years distant.
( Also visible:
(1) C84 (7.6 $6_{m}$ globular cluster)
(2) IC 4406 (10.5m planetary nebula)
(3) NGC 5307 (11.5m planetary nebula)
(4) NGC 5286 ( $7.34_{m}$ globular cluster)
(5) NGC 5643 (10.74m spiral galaxy)


RA: $206.6^{\circ} \mid 13 \mathrm{~h} 46.39^{\prime}-$ DEC: $-51.37^{\circ} \mid-51^{\circ} 21^{\prime}$
C84 (NGC 5286) is a magnitude 7.6 globular cluster.
Angular size is 9 '.
$2.3^{\circ}$ NNE from mag. 2.56 eps Cen.

This impressive globular cluster has been identified recently as being part of the Gaia Sausage - an arc of stars and globular clusters with anomalous orbits. This grouping of orbits was discovered by the Gaia space telescope and it is surmised that 8-11 billion years ago a substantial galaxy of 50 billion solar masses merged with the Milky Way.

Also visible:
(1) C80 (3.6m globular cluster)
(2) NGC 5460 ( $5.6_{m}$ open cluster)
(3) NGC 5307 (11.5m planetary nebula)


RA: $243.3^{\circ} \mid 16 \mathrm{~h} 13.2^{\prime}-$ DEC: $-54.22^{\circ} \mid-54^{\circ} 12^{\prime}$

NGC 6067 is a magnitude 5.6 open cluster. Angular size is 13 '.
6.7 ${ }^{\circ} \mathrm{W}$ from mag. 3.06 zet Ara.

Described by John Herschel as "a most superbly rich and large cluster" is around 102 million years old with a mass of 893 solar masses.

Also visible:
(1) C89 ( $5.4_{m}$ open cluster)
(2) NGC 6167 ( $6.7_{m}$ open cluster)
(3) T Nor ( $6.2_{m}$ variable star)


RA: $218.8^{\circ} \mid 14 \mathrm{~h} 35.2^{\prime}-$ DEC: $-56.55^{\circ} \mid-56^{\circ} 32^{\prime}$
NGC 5662 is a magnitude 5.5 open cluster. Angular size is $12^{\prime}$.
$4.3^{\circ} \mathrm{N}$ from mag.1.7 alf Cen.


93 million years old, this bright open cluster comprises 348 solar masses.

Also visible:
(1) C88 (7.9m open cluster)
(2) NGC 5617 ( $6.3_{m}$ open cluster)
(3) R Cen ( $5.3_{m}$ variable star)


RA: $214.14^{\circ} \mid 14 \mathrm{~h} 16.57^{\prime}-$ DEC: $-59.91^{\circ} \mid-59^{\circ} 54^{\prime}$
R Cen (HD 124601) is a magnitude 5.3 variable star. Magnitude ranges from 11.8 to 5.3 ( $\Delta$ mag. 6.5) with a period of 546 d .
$1.6^{\circ}$ NEE from mag. 0.86 Khadar.

A Mira-type variable with an unusual light curve: until recently the curve was double-peaked but this has now reverted to a normal single peak. The period altered around 1950 from 550 days to approximately 500 days. It is theorized a helium flash (in which the helium layer ignites to form carbon) might have caused the abnormalities.

Also visible:
(1) NGC 5281 (5.9m open cluster)
(2) NGC 5316 ( $6.0_{m}$ open cluster)
(3) NGC 5617 ( $6.3_{m}$ open cluster)
(4) NGC 5662 ( $5.5_{m}$ open cluster)


RA: $240.93^{\circ} \mid 16 \mathrm{~h} 3.7^{\prime}-$ DEC: $-60.5^{\circ} \mid-60^{\circ} 29^{\prime}$
C95 (NGC 6025) is a magnitude 5.1 open cluster.
Angular size is 12 .
Draw a line from Hadar through Rigil Kentaurus, and extend it two further lengths.

Another of the bright southern hemisphere naked-eye clusters, Caldwell 95 is 2,700 light-years from Earth.

Also visible:
(1) C89 ( $5.4_{m}$ open cluster)
(2) NGC 5979 (12.0m planetary nebula)


RA: $206.65^{\circ} \mid 13 \mathrm{~h} 46.6^{\prime}-$ DEC: $-62.9^{\circ} \mid-62^{\circ} 53^{\prime}$

NGC 5281 is a magnitude 5.9 open cluster. Angular size is $5^{\prime}$.
$3.2^{\circ}$ SW from mag.0.86 Khadar.

A small open cluster dominated by a few bright stars. The cluster is 45 million years old, and is 4,200 lightyears distant.

Also visible:
(1) NGC 5316 ( $6.0_{m}$ open cluster)
(2) NGC 5189 (10.0 planetary nebula)
(3) NGC 5315 (11.5m planetary nebula)
(4) R Cen (5.3 $3_{m}$ variable star)

## July: $\mathbf{4 5}^{\circ}$ North



C6: page 234
M13: page 236

CH Cyg: page 234
M57: page 236

R Lyr: page 235
u Her: page 237

M92: page 235
M56: page 237


RA: $269.65^{\circ} \mid 17 \mathrm{~h} 58.59^{\prime}-$ DEC: $66.63^{\circ} \mid 66^{\circ} 38^{\prime}$
C6 (Cat's Eye Nebula, NGC 6543) is a magnitude 8.8 planetary nebula. Angular size is $0.3^{\prime}$.

Draw a line between Delta and Zeta Draconis (Altais and Aldhibah respectively), and another line from Polaris to Gamma Draconis (Eltanin, the brightest star in Draco); C6 is at the intersection of these lines.


A fiercely bright and well-defined planetary nebula, unmistakable once found. The central star of the nebula may be a very close binary system. The complex structures in the nebula may have been caused by jets emanating from an accretion disk arising from mass transfer between the binary components.

Also visible:
(1) NGC 6503 (10.91 $1_{m}$ spiral galaxy)


## CH Cyg

RA: $291.14^{\circ} \mid 19 \mathrm{~h} 24.55^{\prime}-$ DEC: $50.24^{\circ} \mid 50^{\circ} 14^{\prime}$
CH Cyg (HD 182917) is a magnitude 5.6 variable star. Magnitude ranges from 8.49 to 5.6 ( $\Delta$ mag. 2.9).
$1.6^{\circ}$ SSW from mag.3.94 iot Cyg.

The closest symbiotic binary system to Earth (roughly 957 light-years distant), where a white dwarf siphons off the outer layers of a nearby red giant. The white dwarf and its accretion disk and jets are the brighter component of the system, but the two components cannot be resolved visually.

Also visible:
(1) C15 (9.8m planetary nebula)
(2) NGC 6811 ( $6.8_{m}$ open cluster)
(3) RT Cyg ( $6.0_{m}$ variable star)


RA: $283.83^{\circ} \mid 18 \mathrm{~h} 55.33^{\prime}-$ DEC: $43.95^{\circ} \mid 43^{\circ} 57^{\prime}$
R Lyr (HD 175865) is a magnitude 3.88 variable star. Magnitude ranges from 5.0 to 3.88 ( $\Delta \mathrm{mag}$. 1.1) with a period of 46 d .
6.2 ${ }^{\circ}$ NNE from mag. 0.14 Vega.

This brilliant red giant shines brilliantly in the near infrared, where it outshines nearby Vega. R Lyrae is only 298 light-years from Earth.


RA: $259.27^{\circ} \mid 17 \mathrm{~h} 17.09^{\prime}-$ DEC: $43.13^{\circ} \mid 43^{\circ} 8^{\prime}$
M92 (NGC 6341) is a magnitude 6.4 globular cluster.
Angular size is 11 '.
Take the northern edge of the keystone asterism in Hercules, and focus on a point to the north forming an equilateral triangle with this edge. M92 is slightly to the east of this point.

One of the brighter northern globular clusters, this ancient cluster contains about 330,000 tightly packed stars. The cluster contains very few elements heavier than helium, indicating the cluster formed from gas that had not been modified greatly by previous generations of stars. The age is estimated at around 11 billion years.


RA: $250.43^{\circ}$ | 16h $41.7^{\prime}-$ DEC: $36.47^{\circ} \mid 36^{\circ} 28^{\prime}$
M13 (Great Hercules Globular, NGC 6205) is a magnitude 5.8 globular cluster. Angular size is $17^{\prime}$.

Located on the western edge of the distinctive quad of four stars at the heart of Hercules, known as the keystone asterism.

The Great Globular Cluster in Hercules is far behind the brighter globulars of the southern hemisphere, but it is bright enough to be barely visible to the naked eye from a dark site. With a mass of 600,000 Suns it is quite large. Telescopes with an aperture of 100 mm can begin to resolve the stars in this cluster.


RA: $283.4^{\circ} \mid 18 \mathrm{~h} 53.59^{\prime}-$ DEC: $33.03^{\circ} \mid 33^{\circ} 2^{\prime}$
M57 (Ring Nebula, NGC 6720) is a magnitude 8.8 planetary nebula. Angular size is $1.4 \times 1.0^{\prime}$.

Roughly a degree south and east of the lovely double and variable star Sheliak.

Formed from the gas ejected from a red giant as it progressed to its white dwarf stage of life, M57 forms a perfect smoke ring even in a smaller telescope. It has a high surface brightness and stand up well to light pollution.

Also visible:
(1) $\mathrm{T} \operatorname{Lyr}\left(7.5_{m}\right.$ carbon star)
(2) $\mathrm{HK} \mathrm{Lyr}\left(8.5_{m}\right.$ carbon star)


RA: $259.33^{\circ} \mid 17 \mathrm{~h} 17.32^{\prime} —$ DEC: $33.1^{\circ} \mid 33^{\circ} 6^{\prime}$
u Her (HD 156633) is a magnitude 4.69 variable star. Magnitude ranges from 5.37 to 4.69 ( $\Delta$ mag. 0.7 ) with a period of 2.05103 d .
$3.7^{\circ} \mathrm{S}$ from mag. 3.36 pi. Her.


RA: $289.15^{\circ} \mid 19 \mathrm{~h} 16.59^{\prime}-$ DEC: $30.18^{\circ} \mid 30^{\circ} 11^{\prime}$
M56 (NGC 6779) is a magnitude 8.3 globular cluster.
Angular size is 7'.
You can find M56 midway between the famous double star Albireo (the head of Cygnus the swan) and Sulafat (the star of the Lyra quadrilateral that lies furthest from Vega). Alternately, put Albireo on the SE edge of a finder circle, and M56 is on the NW edge.

With a distance of 32,900 light-years, the stars in this globular cluster can only be resolved with telescopes with at least 200 mm aperture. The brightest stars in this cluster are 13th magnitude.

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## July: $\mathbf{1 5}^{\circ}$ North



U Sge: page 240
R Aql: page 242

S Her: page 240
NGC 6633: page 242

NGC 6709: page 241
IC 4665: page 243

X Oph: page 241
IC 4756: page 243


RA: $289.7^{\circ} \mid 19 \mathrm{~h} 18.8^{\prime}-$ DEC: $19.61^{\circ} \mid 19^{\circ} 37^{\prime}$
U Sge (HD 181182) is a magnitude 6.45 variable star. Magnitude ranges from 9.28 to 6.45 ( $\Delta$ mag. 2.8 ) with a period of 3.38062 d .
$6.5^{\circ} \mathrm{NNE}$ from mag. 3.02 zet Aql.

Located within two degrees of the Coathanger asterism, this eclipsing binary system consists of a blue B-type main sequence star and a yellow G-type giant. The blue primary used to be the smaller and fainter component, but it has absorbed the outer layers of its companion as it expanded into its giant phase.


Also visible:
(1) Palomar 10 ( $13.22_{m}$ globular cluster)


RA: $252.97^{\circ} \mid 16 \mathrm{~h} 51.89^{\prime}-$ DEC: $14.94^{\circ} \mid 14^{\circ} 57^{\prime}$
S Her (HD 152276) is a magnitude 6.4 variable star. Magnitude ranges from 13.8 to 6.4 ( $\Delta$ mag. 7.4) with a period of 307 d .
5.5 ${ }^{\circ} \mathrm{W}$ from mag.3.48 Rasalgethi.

This pulsating red giant star is roughly 1,553 light years from Earth.


RA: $282.88^{\circ} \mid 18 \mathrm{~h} 51.5^{\prime}-$ DEC: $10.35^{\circ} \mid 10^{\circ} 21^{\prime}$
NGC 6709 is a magnitude 6.7 open cluster. Angular size is 13 '.
4.8 ${ }^{\circ} \mathrm{SW}$ from mag. 3.02 zet Aql.


This moderately rich cluster it is 141 million years old and is approximately 3,510 light-years distant.

Also visible:
(1) $\mathrm{R} \mathrm{Aql} \mathrm{( } 5.5_{m}$ variable star)
(2) X Oph (5.9 $9_{m}$ variable star)
(3) UV Aql (8.6 $6_{m}$ carbon star)


RA: $279.59^{\circ} \mid 18 \mathrm{~h} 38.35^{\prime}-$ DEC: $8.83^{\circ} \mid 8^{\circ} 50^{\prime}$

X Oph (HD 172171) is a magnitude 5.9 variable star. Magnitude ranges from 9.2 to 5.9 ( $\Delta$ mag. 3.3) with a period of 329 d .
$7.6^{\circ} \mathrm{E}$ from mag. 3.73 72 Oph.

This Mira-type variable has a smaller range in brightness and a relatively luminous minimum brightness because it is actually a binary system and the light of the non-variable component dominates.
The pair are separated by $0.5^{\prime \prime}$ so splitting them is almost impossible, and was first resolved with the $36^{\prime \prime}$ Lick Refractor in 1900.

Also visible:
(1) NGC 6633 (4. $6_{m}$ open cluster)
(2) IC 4756 ( $5.0_{m}$ open cluster)
(3) NGC 6709 ( $6.7_{m}$ open cluster)
(4) DR Ser (8.4m carbon star)


RA: $286.59^{\circ} \mid 19 \mathrm{~h} 6.37^{\prime}-$ DEC: $8.23^{\circ} \mid 8^{\circ} 14^{\prime}$

R Aql (HD 177940) is a magnitude 5.5 variable star. Magnitude ranges from 12.0 to 5.5 ( $\Delta$ mag. 6.5) with a period of 284 d .
$5.6^{\circ} \mathrm{S}$ from mag. 3.02 zet Aql.

This Mira-type variable has a faint magnitude 11.45 optical companion (separation 78.2"; position angle $300^{\circ}$ ). R Aql is roughly 1376 light-years distant.

Also visible:
(1) NGC 6709 ( $6.7_{m}$ open cluster)
(2) NGC 6781 (12.0 m planetary nebula)
(3) DR Ser ( $8.4_{m}$ carbon star)


RA: $276.93^{\circ} \mid 18 \mathrm{~h} 27.7^{\prime}-$ DEC: $6.57^{\circ} \mid 6^{\circ} 34^{\prime}$

NGC 6633 is a magnitude 4.6 open cluster. Angular size is $27^{\prime}$.
5.8 ${ }^{\circ}$ SEE from mag.3.73 72 Oph .

This bright cluster is roughly as large as the Moon. It is relatively old at 660 million years and is about 1,000 light years distant.

Also visible:
(1) IC 4756 ( $5.0_{m}$ open cluster)
(2) NGC 6572 ( $8.5_{m}$ planetary nebula)
(3) X Oph (5.9m variable star)


RA: $266.58^{\circ} \mid 17 \mathrm{~h} 46.29^{\prime}-$ DEC: $5.72^{\circ} \mid 5^{\circ} 43^{\prime}$
IC 4665 is a magnitude 4.2 open cluster. Angular size is $41^{\prime}$.
1.3 ${ }^{\circ}$ NNE from mag.2.94 Cebalrai.

A very large and scattered cluster. Probably due its disorganized and scattered nature, it was overlooked in the NGC catalog.

Also visible:
(1) NGC 6426 (11.01m globular cluster)
(2) NGC 6384 (11.14 $m_{m}$ barred spiral galaxy)


RA: $279.75^{\circ} \mid 18 \mathrm{~h} 39.0^{\prime}-$ DEC: $5.45^{\circ} \mid 5^{\circ} 27^{\prime}$
IC 4756 is a magnitude 5.0 open cluster. Angular size is $52^{\prime}$.
8.8 ${ }^{\circ}$ SEE from mag.3.73 72 Oph .

Also known as Graff's Cluster, this large and bright cluster is roughly 1,300 light-years from the Earth.

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Also visible:
(1) NGC 6633 (4.6m open cluster)
(2) X Oph (5.9m variable star)
(3) DR Ser (8.4m carbon star)

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## July: $-15^{\circ}$ South (1)



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## July: - $\mathbf{1 5}^{\circ}$ South (2)



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RS Oph: page 253
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M12 (NGC 6218) is a magnitude 6.7 globular cluster. Angular size is $15^{\prime}$.

Somewhat west of the midpoint between Rasalhague and Antares.

The stars in this globular cluster can only be resolved in larger telescopes (from around 200 mm aperture) as the brightest stars are 12th magnitude. M12 is quite loosely packed for a globular cluster and is also lacking in low mass stars - it is believed these may have been stripped away by the cluster having moved through a denser region of the Milky Way.

Also visible:
(1) M10 ( $6.6_{m}$ globular cluster)
(2) GCl 50 (14.0m globular cluster)


RA: $254.28^{\circ}\left|16 \mathrm{~h} 57.1^{\prime}-\mathrm{DEC}:-4.1^{\circ}\right|-4^{\circ} 5^{\prime}$

M10 (NGC 6254) is a magnitude 6.6 globular cluster.
Angular size is $15^{\prime}$.
Halfway between Rasalhague and Antares.

A bright globular cluster, relatively close at a distance of 14,300 light-years. Because of its relative closeness, stars in this cluster can be resolved in smaller instruments. The cluster is not particularly large for a globular: it has a mass of 225,000 Suns.

Also visible:
(1) M12 ( $6.7_{m}$ globular cluster)
(2) GCl 50 (14.0 $\mathrm{O}_{\mathrm{m}}$ globular cluster)


RA: $282.77^{\circ} \mid 18 \mathrm{~h} 51.09^{\prime}-$ DEC: $-6.27^{\circ} \mid-6^{\circ} 15^{\prime}$
M11 (Wild Duck Cluster, NGC 6705) is a magnitude 6.3 open cluster. Angular size is $14^{\prime}$.

One third of the distance from Altair to Antares.


Amongst the most massive open clusters known with 11,000 solar masses and a large radius of 95 lightyears. From a dark site this object can be seen by the naked eye, but it is also in a thick part of the Milky Way, so the background can be confusing.

Also visible:
(1) M26 (8.0 m open cluster)
(2) V Aql ( $6.6_{m}$ variable star)
(3) R Sct (4.2 $2_{m}$ variable star)
(4) V Aql ( $6.6_{m}$ carbon star)
(5) S Sct ( $7.3_{m}$ carbon star)


## M26

RA: $281.3^{\circ}\left|18 \mathrm{~h} 45.2^{\prime}-\mathrm{DEC}:-9.4^{\circ}\right|-9^{\circ} 23^{\prime}$

M26 (NGC 6694) is a magnitude 8.0 open cluster.
Angular size is $15^{\prime}$.
6.8 ${ }^{\circ}$ SW from mag.3.55 lam Aql.

A fairly faint and sparse cluster with a busy background, this cluster can be a challenge to pick out. It is distinguished by a drop in stellar density in its core - the opposite of normal clusters, which have their greatest density at the center.

Also visible:
(1) M11 (6.3m open cluster)
(2) NGC 6712 ( $8.1_{m}$ globular cluster)
(3) R Sct (4.2 $2_{m}$ variable star)
(4) S Sct (7.3m carbon star)
(5) T Sct (8.9m carbon star)


RA: $248.13^{\circ}\left|16 \mathrm{~h} 32.5^{\prime}-\mathrm{DEC}:-13.05^{\circ}\right|-13^{\circ} 2^{\prime}$

M107 (NGC 6171) is a magnitude 7.9 globular cluster. Angular size is $10^{\prime}$.
2.7 ${ }^{\circ}$ SSW from mag.2.7 Khan.

This is a sparser and looser globular cluster, first observed by Pierre Méchain in April 1782. It lies at a distance of 21,000 light years. M107 is the lastdiscovered object in the Messier catalog.

Also visible:
(1) V Oph (7.3 $3_{m}$ variable star)


## M17

RA: $275.2^{\circ} \mid 18 \mathrm{~h} 20.79^{\prime}-$ DEC: $-16.18^{\circ} \mid-16^{\circ} 10^{\prime}$

M17 (Omega Nebula, NGC 6618) is a magnitude 7.0 diffuse nebula. Angular size is 11 .

Halfway between Altair and Antares (slightly closer to Antares).

A bright, rich star-forming region nestled in the richest regions of the Milky Way, the Omega Nebula is estimated to have a mass of 800 Suns. The cloud is illuminated by the one million year young open cluster NGC 6618.

Also visible:
(1) M16 ( $6.4_{m}$ open cluster)
(2) M24 (4.6 $6_{m}$ star cloud)
(3) M25 ( $6.5_{m}$ open cluster)
(4) NGC 6605 ( $6.0_{m}$ open cluster)
(5) U Sgr ( $6.28_{m}$ variable star)


RA: $274.23^{\circ}\left|18 \mathrm{~h} 16.9^{\prime}-\mathrm{DEC}:-18.5^{\circ}\right|-18^{\circ} 29^{\prime}$

M24 (Sagittarius Star Cloud, IC 4715) is a magnitude 4.6 star cloud. Angular size is 90 '.

Halfway between Altair and Antares (slightly closer to Antares).

This Messier object is simply a dense region of the Milky Way. Best viewed with binoculars at a dark site, which enable a good overview of the many thousands of stars.

Also visible:
(1) M16 (6.4m open cluster)
(2) M21 ( $6.5_{m}$ open cluster)
(3) M25 ( $6.5_{m}$ open cluster)
(4) NGC 6605 ( $6.0_{m}$ open cluster)
(5) U Sgr ( $6.28_{m}$ variable star)


M23 (NGC 6494) is a magnitude 6.9 open cluster.
Angular size is 27 '.
Slightly south of a point one third of the distance from Antares to Altair.

Consisting of many tiny points of light scattered over the area of a full Moon, it has a mass of 1,200 Suns or more. It is quite old for an open cluster at 330 million years.

Also visible:
(1) M20 (9.0 $0_{m}$ diffuse nebula)
(2) M21 ( $6.5_{m}$ open cluster)
(3) NGC 6567 (11.5m planetary nebula)
(4) NGC 6440 ( $9.2_{m}$ globular cluster)
(5) VX Sgr (6.52 $2_{m}$ variable star)


RA: $271.15^{\circ} \mid 18 \mathrm{~h} 4.59^{\prime}-$ DEC: $-22.5^{\circ} \mid-22^{\circ} 29^{\prime}$
M21 (NGC 6531) is a magnitude 6.5 open cluster.
Angular size is 13 '.
Forms an equilateral triangle with Antares and Sargas
in Scorpius's tail.
This reasonably bright open cluster has the mass of nearly 800 Suns but only has a few bright members with the remaining stars much smaller and dimmer, resulting in a relatively sparse but quite tightly packed cluster.

Also visible:
(1) M8 ( $6.0_{m}$ diffuse nebula)
(2) M23 ( $6.9_{m}$ open cluster)
(3) M24 (4.6 $m_{m}$ star cloud)
(4) M28 ( $6.8_{m}$ globular cluster)
(5) VX Sgr ( $6.52_{m}$ variable star)


M22 (Sagittarius Cluster, NGC 6656) is a magnitude
5.1 globular cluster. Angular size is 24 '.
2.4 ${ }^{\circ}$ NE from mag.2.94 Kaus Borealis.

This very bright globular cluster is an easy spot with the naked eye for those with dark skies. Its stars are also quite bright at 11th magnitude, due to its relatively close distance of only 10,600 light-years. It was discovered in 1665 by Abraham Ihle.

Also visible:
(1) M25 ( $6.5_{m}$ open cluster)
(2) M28 ( $6.8_{m}$ globular cluster)
(3) NGC 6716 (6.9m open cluster)
(4) NGC 6638 ( $9.02_{m}$ globular cluster)
(5) NGC 6642 ( $9.13_{m}$ globular cluster)


RA: $276.13^{\circ} \mid 18 \mathrm{~h} 24.5^{\prime}-$ DEC: $-24.87^{\circ} \mid-24^{\circ} 51^{\prime}$
M28 (NGC 6626) is a magnitude 6.8 globular cluster.
Angular size is 11 '.
0.9 ${ }^{\circ}$ NW from mag.2.94 Kaus Borealis.

The stars in the globular cluster begin to be resolved with a telescope aperture of around 150 mm . M28 contains 11 millisecond pulsars. These are old neutron stars that have been given fresh impetus from infalling matter from a companion star.

Also visible:
(1) M8 ( $6.0_{m}$ diffuse nebula)
(2) M21 ( $6.5_{m}$ open cluster)
(3) M22 (5.1 $1_{m}$ globular cluster)
(4) NGC 6544 ( $7.77_{m}$ globular cluster)
(5) VX Sgr ( $6.52_{m}$ variable star)


M14 (NGC 6402) is a magnitude 7.6 globular cluster.
Angular size is 12 .
Draw a line from Antares through Sabik, and extend it an equal distance.

Although nearly twice the mass of M13, M14 is also $50 \%$ more distant and closer to the galactic plane and its obscuring dust, so overall it is much fainter. The brightest stars in this cluster only shine at magnitude 14. Three degrees southwest of M13 is a faint globular cluster, NGC 6366 (magnitude 9.5).

Also visible:
(1) NGC 6366 (9.2m globular cluster)
(2) RS Oph (4.3m variable star)


RA: $281.87^{\circ} \mid 18 \mathrm{~h} 47.48^{\prime}-$ DEC: $-5.71^{\circ} \mid-5^{\circ} 41^{\prime}$
R Sct (HD 173819) is a magnitude 4.2 variable star.
Magnitude ranges from 8.6 to 4.2 ( $\Delta$ mag. 4.4) with a period of 146.5 d .
4.7 ${ }^{\circ} \mathrm{W}$ from mag.3.55 lam Aql.


The brightest of the RV Tauri class of variables, R Sct is marked by a long, highly irregular period with extreme minima. It is about 4,000 light-years distant.
Also visible:
(1) M11 ( $6.3_{m}$ open cluster)
(2) M26 (8.0 m open cluster)
(3) NGC 6712 ( $8.1_{m}$ globular cluster)
(4) V Aql ( $6.6_{m}$ carbon star)
(5) S Sct ( $7.3_{m}$ carbon star)


RA: $267.55^{\circ} \mid 17 \mathrm{~h} 50.21^{\prime}-$ DEC: $-6.71^{\circ} \mid-6^{\circ} 41^{\prime}$

RS Oph (HD 162214) is a magnitude 4.3 variable star. Magnitude ranges from 12.5 to 4.3 ( $\Delta$ mag. 8.2).
3.7 ${ }^{\circ}$ NW from mag.3.5 Sinistra.

A recurrent nova system, RS Oph is usually magnitude 12.5 but has erupted to magnitude 5 in 1898, 1933, 1958, 1967, 1985, 2006 and 2021. It appears the system is binary consisting of a white dwarf accreting matter from a nearby red giant.
Approximately every 15 years sufficient matter builds up on the white dwarf to cause a runaway fusion reaction.

Also visible:
(1) M14 (7.6 m $_{m}$ globular cluster)
(2) NGC 6517 (10.23 $3_{m}$ globular cluster)
(3) NGC 6539 ( $9.33_{m}$ globular cluster)


RA: $274.7^{\circ} \mid 18 \mathrm{~h} 18.79^{\prime}-$ DEC: $-13.78^{\circ} \mid-13^{\circ} 46^{\prime}$
M16 (Eagle Nebula, NGC 6611) is a magnitude 6.4 open cluster. Angular size is $7^{\prime}$.

Halfway between Altair and Antares (slightly closer to Antares).

This bright, rich star-forming region is the site of the Hubble Space Telescope's iconic Pillars of Creation photograph. The open cluster illuminating the gas cloud is NGC 6611. The brightest member of this extremely young cluster is HD 168076, a binary system with a total mass of 100 Suns and the luminosity of a million Suns.

Also visible:
(1) M17 (7.0 $0_{m}$ diffuse nebula)
(2) M18 (7.5m open cluster)
(3) M24 (4.6m star cloud)
(4) NGC 6605 ( $6.0_{m}$ open cluster)
(5) $\mathrm{SS} \operatorname{Sgr}\left(9.0_{m}\right.$ carbon star)


RA: $274.27^{\circ} \mid 18 \mathrm{~h} 17.09^{\prime}-$ DEC: $-14.97^{\circ} \mid-14^{\circ} 57^{\prime}$

NGC 6605 is a magnitude 6.0 open cluster. Angular size is 29 .
6.8 ${ }^{\circ}$ SE from mag.3.5 Sinistra.

Located between the Eagle and Omega Nebulae, this cluster is lost within a crowded field of stars.

Also visible:
(1) M16 ( $6.4_{m}$ open cluster)
(2) M17 (7.0 $0_{m}$ diffuse nebula)
(3) M24 (4.6m star cloud)
(4) M25 ( $6.5_{m}$ open cluster)
(5) U Sgr (6.28 $8_{m}$ variable star)


RA: $274.98^{\circ} \mid 18 \mathrm{~h} 19.9^{\prime}-$ DEC: $-17.13^{\circ} \mid-17^{\circ} 7^{\prime}$
M18 (NGC 6613) is a magnitude 7.5 open cluster.
Angular size is $9^{\prime}$.
Halfway between Altair and Antares (slightly closer to Antares).


This cluster is 4,250 light-years from Earth, with an estimated mass of 188 Suns and an age of 33 million years.


Also visible:
(1) M16 (6.4m open cluster)
(2) M24 (4.6 $\mathrm{m}_{\mathrm{m}}$ star cloud)
(3) M25 ( $6.5_{m}$ open cluster)
(4) NGC 6605 ( $6.0_{m}$ open cluster)
(5) U Sgr ( $6.28_{m}$ variable star)


## M9

RA: $259.8^{\circ}\left|17 \mathrm{~h} 19.2^{\prime}-\mathrm{DEC}:-18.52^{\circ}\right|-18^{\circ} 30^{\prime}$

M9 (NGC 6333) is a magnitude 7.7 globular cluster.
Angular size is 9.3'.
$3.5^{\circ} \mathrm{SE}$ from mag.2.63 Sabik.

A reasonably bright globular cluster accompanied by two other globular clusters: 1.3 degrees to the northeast is NGC 6356 and 1.3 degrees to the southeast is NGC 6342. The fact other globulars are in the same field of view is no coincidence: M9 is close to the galactic core so we view it through a thick section of the galactic halo. The mass of this cluster is 420,000 Suns.

Also visible:
(1) IC 4634 (11.0 $0_{m}$ planetary nebula)
(2) NGC 6287 ( $9.35_{m}$ globular cluster)
(3) NGC 6342 ( $9.66_{m}$ globular cluster)
(4) NGC 6356 ( $8.25_{m}$ globular cluster)


RA: $277.9^{\circ}\left|18 \mathrm{~h} 31.59^{\prime}-\mathrm{DEC}:-19.25^{\circ}\right|-19^{\circ} 14^{\prime}$
M25 (IC 4725) is a magnitude 6.5 open cluster.
Angular size is $40^{\prime}$.
6.2 ${ }^{\circ} \mathrm{N}$ from mag.2.94 Kaus Borealis.

Thirteen light years across and with a mass of nearly 2,000 Suns, this cluster is 2,000 light years from Earth and partially obscured by dust clouds.


Also visible:
(1) M17 (7.0 $0_{m}$ diffuse nebula)
(2) M22 (5.1 $1_{m}$ globular cluster)
(3) M24 (4.6 $6_{m}$ star cloud)
(4) NGC 6605 ( $6.0_{m}$ open cluster)
(5) $\mathrm{U} \mathrm{Sgr}\left(6.28_{m}\right.$ variable star)


M20 (Trifid Nebula, NGC 6514) is a magnitude 9.0 diffuse nebula. Angular size is $28^{\prime}$.

Forms an equilateral triangle with Antares and Sargas in Scorpius's tail.

The Trifid Nebula is divided into three apparent lobes by dark dust lanes. M20 combines an emission nebula, a reflection nebula, a star cluster and a dark nebula, resulting in a complex and fascinating object.

Also visible:
(1) M8 ( $6.0_{m}$ diffuse nebula)
(2) M21 (6.5m open cluster)
(3) M23 ( $6.9_{m}$ open cluster)
(4) M24 (4.6 m $_{m}$ star cloud)
(5) VX Sgr ( $6.52_{m}$ variable star)


RA: $270.95^{\circ} \mid 18 \mathrm{~h} 3.79^{\prime}-$ DEC: $-24.38^{\circ} \mid-24^{\circ} 22^{\prime}$
M8 (Lagoon Nebula, NGC 6523) is a magnitude 6.0 diffuse nebula. Angular size is $90 \times 40^{\prime}$.

Travel from the scorpion's sting Shaula to Kaus Australis, turn right and continue an equal distance.

This star-forming region is one of only two such objects visible to the naked eye from northern latitudes. It spans 110 by 50 light-years. At its heart is the star cluster NGC 6530 (a dozen stars are visible with a small telescope, but it is believed that the cluster contains 2,768 stars. Of these over 300 are T Tauri variables, namely very young stars typically with an accretion disk. Most of the stars formed in the past two million years.

Also visible:
(1) M21 ( $6.5_{m}$ open cluster)
(2) M28 ( $6.8_{m}$ globular cluster)
(3) NGC 6544 ( $7.77_{m}$ globular cluster)
(4) NGC 6553 ( $8.06_{m}$ globular cluster)
(5) VX Sgr ( $6.52_{m}$ variable star)


M19 (NGC 6273) is a magnitude 6.8 globular cluster.
Angular size is $14^{\prime}$.
$4.5^{\circ}$ SWW from mag. 3.37 the Oph.

This globular appears to be slightly flattened, but this is only due to intervening dust obscuring one edge. In infrared images the cluster is perfectly spherical (infrared is not absorbed by dust).
( Also visible:
(1) M62 ( $6.5_{m}$ globular cluster)
(2) NGC 6293 ( $8.22_{m}$ globular cluster)
(3) NGC 6304 ( $8.22_{m}$ globular cluster)
(4) BF Oph ( $6.93_{m}$ variable star)
(5) RR Sco (5.0 ${ }_{m}$ variable star)

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## July: $\mathbf{- 4 5}{ }^{\circ}$ South (1)



## July: -45 ${ }^{\circ}$ South (2)




RA: $255.3^{\circ} \mid 17 \mathrm{~h} 1.2^{\prime}-$ DEC: $-30.12^{\circ} \mid-30^{\circ} 6^{\prime}$
M62 (NGC 6266) is a magnitude 6.5 globular cluster.
Angular size is 14 '.
4.7 ${ }^{\circ}$ NNE from mag. 2.36 eps Sco.

This southern Messier is a relatively bright globular cluster, and is quite massive at 1.22 million solar masses. It is suspected to harbor an intermediate mass black hole, in the region of several thousand solar masses.

Also visible:
(1) M19 ( $6.8_{m}$ globular cluster)
(2) NGC 6304 ( $8.22_{m}$ globular cluster)
(3) BF Oph ( $6.93_{m}$ variable star)
(4) RR Sco ( $5.0_{m}$ variable star)
(5) SU Sco ( $8.0_{m}$ carbon star)


RA: $254.16^{\circ} \mid 16 \mathrm{~h} 56.63^{\prime}-$ DEC: $-30.58^{\circ} \mid-30^{\circ} 34^{\prime}$
RR Sco (HD 152783) is a magnitude 5.0 variable star. Magnitude ranges from 12.4 to 5.0 ( $\Delta$ mag. 7.4) with a period of 281 d .
$3.9^{\circ}$ NNE from mag. 2.36 eps Sco.

Just over a degree west of M62, this Mira-type variable is roughly 903 light-years from Earth.

Also visible:
(1) M19 ( $6.8_{m}$ globular cluster)
(2) M62 ( $6.5_{m}$ globular cluster)
(3) NGC 6304 ( $8.22_{m}$ globular cluster)
(4) BF Oph ( $6.93_{m}$ variable star)
(5) SU Sco ( $8.0_{m}$ carbon star)


RA: $266.1^{\circ} \mid 17 \mathrm{~h} 44.4^{\prime}-$ DEC: $-32.35^{\circ} \mid-32^{\circ} 20^{\prime}$

NGC 6416 is a magnitude 5.7 open cluster. Angular size is $18^{\prime}$.
4.8 ${ }^{\circ}$ NNW from mag.3.25 G Sco.


Dwarfed by glittering M6, NGC 6416 is a bright cluster in its own right.

Also visible:
(1) M6 (4.2m open cluster)
(2) M7 (3.3m open cluster)
(3) NGC 6441 ( $7.15_{m}$ globular cluster)
(4) NGC 6522 ( $8.27_{m}$ globular cluster)
(5) SX Sco ( $8.5_{m}$ carbon star)


RA: $280.8^{\circ} \mid 18 \mathrm{~h} 43.2^{\prime}-$ DEC: $-32.3^{\circ} \mid-32^{\circ} 17^{\prime}$

M70 (NGC 6681) is a magnitude 7.9 globular cluster.
Angular size is $8^{\prime}$.
4.4 ${ }^{\circ}$ NEE from mag.1.95 Kaus Australis.

This cluster is close to M69 in size, luminosity and distance. In contrast to M69, M70 has a very dense core, having undergone "core collapse", where the more massive stars of the cluster have migrated closer to the center of the cluster. The cluster is approximately 29,400 light-years from Earth.
(1) Also visible:
(1) M54 (7.6 m $_{m}$ globular cluster)
(2) M69 (7.6m globular cluster)
(3) NGC 6624 ( $7.87_{m}$ globular cluster)
(4) NGC 6652 ( $8.62_{m}$ globular cluster)
(5) NGC 6723 ( $7.01_{m}$ globular cluster)


RA: $289.14^{\circ} \mid 19 \mathrm{~h} 16.54^{\prime}-$ DEC: $-33.52^{\circ} \mid-33^{\circ} 30^{\prime}$
RY $\operatorname{Sgr}$ (HD 180093) is a magnitude 5.8 variable star. Magnitude ranges from 14.0 to 5.8 ( $\Delta$ mag. 8.2) with a period of Period.
4.6 ${ }^{\circ}$ SE from mag.2.71 Ascella.

A member of the exceptionally rare R Corona Borealis class of variables, this star shows irregular, sudden and extreme drops in brightness. These stars are very deficient in hydrogen, and are believed to be the result of the merger of two white dwarfs.

Also visible:
(1) C68 (9.7m bright nebula)
(2) M54 (7.6 $6_{m}$ globular cluster)
(3) NGC 6723 ( $7.01_{m}$ globular cluster)


## C68

RA: $285.48^{\circ} \mid$ 19h $1.9^{\prime}-$ DEC: $-36.95^{\circ} \mid-36^{\circ} 56^{\prime}$

C68 (NGC 6729) is a magnitude 9.7 bright nebula.
Angular size is $1.0^{\prime}$.
7.0 ${ }^{\circ} \mathrm{S}$ from mag.2.71 Ascella.

This bright reflection nebula is one of the closest starforming regions to the Earth being only 400 lightyears distant.

Also visible:
(1) M70 (7.9 ${ }^{m}$ globular cluster)
(2) IC 1297 ( $9.8_{m}$ planetary nebula)
(3) NGC 6723 ( $7.01_{m}$ globular cluster)
(4) RY Sgr (5.8m variable star)


RA: $259.63^{\circ} \mid 17 \mathrm{~h} 18.5^{\prime}-$ DEC: $-42.95^{\circ} \mid-42^{\circ} 56^{\prime}$
NGC 6322 is a magnitude 6.0 open cluster. Angular size is 10 .
1.1 ${ }^{\circ}$ NEE from mag.3.44 eta Sco.


Also visible:
(1) C76 (2.6 $\mathrm{o}_{\mathrm{m}}$ open cluster)
(2) NGC 6242 ( $6.4_{m}$ open cluster)
(3) NGC 6250 ( $5.9_{m}$ open cluster)
(4) NGC 6388 ( $6.72_{m}$ globular cluster)
(5) RS Sco ( $6.2_{m}$ variable star)


RA: $254.5^{\circ} \mid 16 \mathrm{~h} 58.0^{\prime}-$ DEC: $-45.8^{\circ} \mid-45^{\circ} 47^{\prime}$
NGC 6250 is a magnitude 5.9 open cluster. Angular size is $8^{\prime}$.
$3.4^{\circ} \mathrm{S}$ from mag. 3.75 zet02 Sco.

Aged at 14 million years, this bright, young cluster is about 2,820 light-years from Earth.
( Also visible:
(1) C76 (2.6 ${ }_{m}$ open cluster)
(2) C82 (5.2m open cluster)
(3) NGC 6167 ( $6.7_{m}$ open cluster)
(4) NGC 6322 ( $6.0_{m}$ open cluster)
(5) RS Sco ( $6.2_{m}$ variable star)


## C82

RA: $250.33^{\circ} \mid 16 \mathrm{~h} 41.29^{\prime}-$ DEC: $-48.77^{\circ} \mid-48^{\circ} 45^{\prime}$
C82 (NGC 6193) is a magnitude 5.2 open cluster.
Angular size is $15^{\prime}$.
$6.8^{\circ}$ SSW from mag. 3.75 zet02 Sco.

A bright open cluster, Caldwell 82 contains two Otype stars. These rare, immense stars have very short lives ending in supernovae, indicating that this cluster is very young.

Also visible:
(1) NGC 6167 ( $6.7_{m}$ open cluster)
(2) NGC 6250 ( $5.9_{m}$ open cluster)
(3) RS Sco ( $6.2_{m}$ variable star)


RA: $287.73^{\circ} \mid 19 \mathrm{~h} 10.9^{\prime}-$ DEC: $-59.98^{\circ} \mid-59^{\circ} 58^{\prime}$
C93 (Pavo Globular Cluster, NGC 6752) is a magnitude 5.4 globular cluster. Angular size is 20'.
8.9 ${ }^{\circ}$ NW from mag.3.64 del Pav.

This globular cluster holds over 100,000 stars in a 100 light-year sphere.

Also visible:
(1) C101 (9.0 m barred spiral galaxy)


RA: $283.77^{\circ} \mid 18 \mathrm{~h} 55.09^{\prime}-$ DEC: $-30.48^{\circ} \mid-30^{\circ} 28^{\prime}$
M54 (NGC 6715) is a magnitude 7.6 globular cluster. Angular size is $9^{\prime}$.
1.7${ }^{\circ}$ SWW from mag.2.71 Ascella.

This large globular cluster is very distant at over 87,000 light-years, and it belongs to a different galaxy - the Sagittarius Dwarf Elliptical Galaxy (SagDEG). At its great distance, it is not possible to resolve the stars in this cluster. The cluster is exceptionally dense, with enhanced surface brightness in its core. There is debate as to whether M54 is a true globular cluster or is in fact SagDEG's core.

Also visible:
(1) M70 (7.9 $9_{m}$ globular cluster)
(2) NGC 6652 ( $8.62_{m}$ globular cluster)
(3) RY $\operatorname{Sgr}\left(5.8_{m}\right.$ variable star)


RA: $265.02^{\circ} \mid 17 \mathrm{~h} 40.09^{\prime}-$ DEC: $-32.22^{\circ} \mid-32^{\circ} 12^{\prime}$
M6 (Butterfly Cluster, NGC 6405) is a magnitude 4.2 open cluster. Angular size is $25^{\prime}$.
5.0 ${ }^{\circ}$ NNE from mag.1.71 Shaula.

A bright, naked-eye cluster rich with many stars of varying brightness. The brightest star in this cluster is BM Scorpii, varying semi-regularly between magnitude 5.5 and 7.0 over a period of roughly 800 days.

Also visible:
(1) M7 (3.3m open cluster)
(2) NGC 6416 (5.7m open cluster)
(3) NGC 6453 ( $10.08_{m}$ globular cluster)
(4) Terzan 6 ( $13.85_{m}$ globular cluster)
(5) SX Sco ( $8.5_{m}$ carbon star)


RA: $277.85^{\circ} \mid 18 \mathrm{~h} 31.4^{\prime}-$ DEC: $-32.35^{\circ} \mid-32^{\circ} 20^{\prime}$

M69 (NGC 6637) is a magnitude 7.6 globular cluster. Angular size is $7^{\prime}$.
$2.5^{\circ}$ NE from mag.1.95 Kaus Australis.

This is a very metal-rich globular cluster meaning the gas that formed the cluster was enriched by heavier elements ejected by previous generations of stars. With Messier 69 centered in the finder scope, another globular is just over a degree to the southeast (NGC 6637 , mag. 8.3), while Messier 70 is 2.5 degrees due east.

Also visible:
(1) M70 (7.9 ${ }_{m}$ globular cluster)
(2) NGC 6558 ( $9.26_{m}$ globular cluster)
(3) NGC 6569 ( $8.55_{m}$ globular cluster)
(4) NGC 6624 ( $7.87_{m}$ globular cluster)
(5) NGC 6652 (8.62m globular cluster)


## M7

RA: $268.48^{\circ} \mid 17 \mathrm{~h} 53.9^{\prime}-$ DEC: $-34.82^{\circ} \mid-34^{\circ} 48^{\prime}$
M7 (Ptolemy's Cluster, NGC 6475) is a magnitude 3.3 open cluster. Angular size is $80^{\prime}$.
$2.3^{\circ}$ NNE from mag.3.25 G Sco.

The Messier object furthest south, M7 was first described 130 AD by Ptolemy. Giovanni Batista Hodierna counted 30 stars in 1654. The cluster spans 25 light-years and lies at a distance of nearly 1,000 light-years.

Also visible:
(1) M6 (4.2m open cluster)
(2) NGC 6416 (5.7m open cluster)
(3) NGC 6441 ( $7.15_{m}$ globular cluster)
(4) NGC 6569 ( $8.55_{m}$ globular cluster)
(5) SX Sco ( $8.5_{m}$ carbon star)


RA: $258.43^{\circ} \mid 17 \mathrm{~h} 13.7^{\prime}-$ DEC: $-37.1^{\circ} \mid-37^{\circ} 5^{\prime}$
C69 (Bug Nebula, NGC 6302) is a magnitude 12.8
planetary nebula. Angular size is $0.8^{\prime}$.
3.4 ${ }^{\circ} \mathrm{W}$ from mag.2.8 Lesuth.


This complex planetary nebula was formed in several outbursts from the central star over the past 5000 years. The central star is extremely faint and was first observed with the Hubble Space Telescope in 2009.

Also visible:
(1) NGC 6242 ( $6.4_{m}$ open cluster)
(2) NGC 6337 (12.0 m planetary nebula)
(3) NGC 6256 (11.29m globular cluster)
(4) NGC 6380 (11.31m globular cluster)
(5) RT Sco (7.0 $0_{m}$ variable star)


RA: $253.5^{\circ} \mid 16 \mathrm{~h} 54.0^{\prime}-$ DEC: $-41.8^{\circ} \mid-41^{\circ} 47^{\prime}$
C76 (Northern Jewel Box, NGC 6231) is a magnitude 2.6 open cluster. Angular size is 15 '.
$0.5^{\circ}$ NNW from mag. 3.75 zet02 Sco.

This brilliant naked-eye cluster can be easily admired under any conditions with any telescope. Under dark skies, this cluster appears to form the head of a "False Comet", with the tail formed by the clusters Collinder 316 and Trumpler 24.

Also visible:
(1) NGC 6242 ( $6.4_{m}$ open cluster)
(2) NGC 6250 ( $5.9_{m}$ open cluster)
(3) NGC 6322 ( $6.0_{m}$ open cluster)
(4) NGC 6153 (11.0 $0_{m}$ planetary nebula)
(5) RS Sco ( $6.2_{m}$ variable star)


## C78

RA: $272.0^{\circ}$ | $18 \mathrm{~h} 8.0^{\prime}-$ DEC: $-43.7^{\circ} \mid-43^{\circ} 41^{\prime}$
C78 (Cacciatore Cluster, NGC 6541) is a magnitude 6.6 globular cluster. Angular size is 13 '.
$4.0^{\circ} \mathrm{NWW}$ from mag.3.76 alf Tel.


A bright globular cluster with the mass of half a million Suns, it is possible to resolve the stars in this cluster with a moderately-sized telescope.
Also visible:
(1) IC 4663 ( $12.5_{m}$ planetary nebula)
(2) IC 4699 ( $12.5_{m}$ planetary nebula)
(3) NGC 6496 ( $8.54_{m}$ globular cluster)


C81 (NGC 6352) is a magnitude 8.1 globular cluster.
Angular size is $7^{\prime}$.
$1.7^{\circ} \mathrm{NW}$ from mag.2.97 alf Ara.

A loose globular cluster. The member stars can be resolved with moderately sized telescopes.

Also visible:
(1) IC 4651 ( $6.9_{m}$ open cluster)
(2) NGC 6326 (11.5m planetary nebula)
(3) IC 4663 ( $12.5_{m}$ planetary nebula)
(4) NGC 6388 ( $6.72_{m}$ globular cluster)
(5) U Ara (7.7m variable star)


RA: $265.18^{\circ} \mid 17 \mathrm{~h} 40.7^{\prime}-$ DEC: $-53.67^{\circ} \mid-53^{\circ} 39^{\prime}$
C86 (NGC 6397) is a magnitude 5.6 globular cluster.
Angular size is 26 '.
$2.9^{\circ} \mathrm{NE}$ from mag. 2.8 bet Ara.

Relatively close at only 7,800 light-years, Caldwell 86 was close enough to have it's distance accurately measured by parallax by the Hubble Space Telescope. This in turn allowed an accurate estimate of age at 13.4 billion years, very nearly as old at the universe. Being relatively close, amateur instruments can more easily resolve its member stars.

Also visible:
(1) IC 4651 ( $6.9_{m}$ open cluster)
(2) NGC 6326 (11.5m planetary nebula)
(3) U Ara ( $7.7_{m}$ variable star)


RA: $287.45^{\circ}\left|19 \mathrm{~h} 9.79^{\prime}-\mathrm{DEC}:-63.85^{\circ}\right|-63^{\circ} 50^{\prime}$

C101 (NGC 6744) is a magnitude 9.0 barred spiral galaxy. Angular size is $16 \times 10^{\prime}$.
$6.5^{\circ}$ NWW from mag.3.64 del Pav.

Discovered in 1826 by James Dunlop, this galaxy is relatively close (at 30 million light-years distance) and the bright core is easily seen in smaller telescopes. Dark skies and larger telescopes can reveal the dim extended disk.

Also visible:
(1) C93 (5.4m globular cluster)

## August: $\mathbf{4 5}^{\circ}$ North (1)



## August: $\mathbf{4 5}^{\circ}$ North (2)




RA: $328.43^{\circ} \mid 21 \mathrm{~h} 53.7^{\prime}-$ DEC: $62.6^{\circ} \mid 62^{\circ} 36^{\prime}$
NGC 7160 is a magnitude 6.1 open cluster. Angular size is 7 .
$4.0^{\circ} \mathrm{E}$ from mag.2.6 Alderamin.

This young cluster is aged less than 19 million years old and is about 2,570 light-years from Earth. The cluster is within the Cepheus bubble, a 10 degree wide shell of dust.

Also visible:
(1) mu Cep ( $3.7_{m}$ carbon star)


## mu Cep

RA: $325.88^{\circ} \mid 21 \mathrm{~h} 43.5^{\prime}-$ DEC: $58.8^{\circ} \mid 58^{\circ} 48^{\prime}$
mu Cep (Garnet Star, NGC 206936) is a magnitude 3.7 carbon star. Magnitude ranges from 5.0 to 3.7 ( $\Delta$ mag. 1.3).
$3.6^{\circ}$ NWW from mag. 3.62 zet Cep.

The famed Garnet Star was one of the first carbon stars to be observed, and was named by William Herschel. This red hypergiant (B-V color index: 2.24) is so vast that it would encompass Jupiter's orbit if it were our sun. The absolute magnitude of this star is -7.6 . Recent Gaia data places this star 3,060 light years from Earth.

Also visible:
(1) NGC 7160 ( $6.1_{m}$ open cluster)


RA: $296.2^{\circ} \mid 19 \mathrm{~h} 44.79^{\prime}-$ DEC: $50.52^{\circ} \mid 50^{\circ} 31^{\prime}$
C15 (Blinking Planetary, NGC 6826) is a magnitude
9.8 planetary nebula. Angular size is $0.5^{\prime}$.
$2.6^{\circ}$ SEE from mag. 3.94 iot Cyg.

This bright planetary nebula is most easily seen using averted vision, leading to the illusion that the object blinks into view.


Also visible:
(1) NGC 6811 ( $6.8_{m}$ open cluster)
(2) NGC 6884 (12.5m planetary nebula)
(3) R Cyg ( $6.1_{m}$ variable star)
(4) RT Cyg ( $6.0_{m}$ variable star)
(5) CH Cyg (5. $6_{m}$ variable star)


RA: $295.91^{\circ} \mid 19 \mathrm{~h} 43.62^{\prime}-$ DEC: $48.78^{\circ} \mid 48^{\circ} 47^{\prime}$

RT Cyg (HD 186686) is a magnitude 6.0 variable star. Magnitude ranges from 13.1 to 6.0 ( $\Delta$ mag. 7.1) with a period of 190 d .
$3.6^{\circ} \mathrm{N}$ from mag. 2.97 del Cyg.

This moderately red pulsating giant shares the finder circle with two planetary nebulae. NGC 6833 is one degree to the east, while Caldwell 15 is two degrees to the north.

Also visible:
(1) C 15 ( $9.8_{m}$ planetary nebula)
(2) NGC 6811 ( $6.8_{m}$ open cluster)
(3) NGC 6884 (12.5m planetary nebula)
(4) R Cyg ( $6.1_{m}$ variable star)
(5) CH Cyg (5. $\sigma_{m}$ variable star)


RA: $304.9^{\circ} \mid 20 \mathrm{~h} 19.6^{\prime}-$ DEC: $47.89^{\circ} \mid 47^{\circ} 54^{\prime}$
U Cyg (HD 193680) is a magnitude 5.9 variable star. Magnitude ranges from 12.1 to 5.9 ( $\Delta$ mag. 6.2) with a period of 463d.
$1.5^{\circ} \mathrm{NE}$ from mag.3.95 192578.


This extremely red carbon star has a B-V color index of 3.31 and a spectral type of C9,2e.

Also visible:
(1) NGC 6884 (12.5m planetary nebula)
(2) V Cyg ( $7.8_{m}$ carbon star)


## C20

RA: $314.7^{\circ} \mid 20 \mathrm{~h} 58.79^{\prime}-$ DEC: $44.33^{\circ} \mid 44^{\circ} 20^{\prime}$
C20 (North America Nebula, NGC 7000) is a magnitude 6.0 bright nebula. Angular size is $120 \times 100^{\prime}$.
$1.1^{\circ}$ NWW from mag. 3.92 ksi Cyg.

A very difficult, if not entirely impossible, target for visual observers in light-polluted surroundings. Described by both William Herschel and his son John as "faint", the resemblance to the North American continent was a photographic discovery.
Also visible:
(1) IC 1369 ( $6.8_{m}$ open cluster)
(2) NGC 7027 ( $9.0_{m}$ planetary nebula)
(3) NGC 7048 (11.5m planetary nebula)
(4) NGC 7026 ( $12.0_{m}$ planetary nebula)
(5) V Cyg (7.8m carbon star)


RA: $305.98^{\circ} \mid 20 \mathrm{~h} 23.9^{\prime}-$ DEC: $38.53^{\circ} \mid 38^{\circ} 32^{\prime}$
M29 (NGC 6913) is a magnitude 7.1 open cluster.
Angular size is $7^{\prime}$.
$1.7^{\circ} \mathrm{S}$ from mag. 2.32 Sadr.


A fairly bright open cluster with perhaps 20 bright member stars, this cluster is around 13 million years old.


Also visible:
(1) C27 ( $7.5_{m}$ bright nebula)
(2) NGC 6871 ( $5.2_{m}$ open cluster)
(3) X Cyg ( $5.85_{m}$ variable star)
(4) RS Cyg ( $6.6_{m}$ carbon star)


RA: $310.85^{\circ} \mid 20 \mathrm{~h} 43.4^{\prime}-$ DEC: $35.59^{\circ} \mid 35^{\circ} 35^{\prime}$
X Cyg (HD 197572) is a magnitude 5.85 variable star. Magnitude ranges from 6.91 to 5.85 ( $\Delta$ mag. 1.1) with a period of 16.38633 d .
1.7 ${ }^{\circ}$ NNW from mag.2.64 Gienakh.

This bright Cepheid variable has a moderate range with a bright minimum over a relatively short period. The extremely regular change in brightness can be followed with a pair of binoculars. Its distance is estimated 3,623 light-years.
Also visible:
(1) C33 (8.0 m supernova remnant)
(2) M29 (7.1m open cluster)


RA: $314.1^{\circ} \mid 20 \mathrm{~h} 56.4^{\prime}-$ DEC: $31.72^{\circ} \mid 31^{\circ} 43^{\prime}$
C33 (East Veil Nebula, NGC 6992/5) is a magnitude 8.0 supernova remnant. Angular size is $60 \times 8^{\prime}$.
3.1 ${ }^{\circ}$ SE from mag.2.64 Gienakh.

The Eastern Veil nebula is an extensive fragment of a supernova remnant. Although fairly bright, it is a widely extended object so the best chance of viewing it is with a nebula filter such as an OIII filter, which will block all light except the wavelengths emitted by the glowing nebula. Because of its size, the nebula is best viewed in very small, widefield telescopes.

Also visible:
(1) C34 (8.0 mupernova remnant)
(2) $\mathrm{X} \operatorname{Cyg}$ ( $5.85_{m}$ variable star)


RA: $308.7^{\circ} \mid 20 \mathrm{~h} 34.79^{\prime}-$ DEC: $60.15^{\circ} \mid 60^{\circ} 9^{\prime}$
C12 (NGC 6946) is a magnitude 9.7 spiral galaxy.
Angular size is $12 \times 10^{\prime}$.
$2.0^{\circ}$ SW from mag.3.59 eta Cep.

A bright face-on spiral galaxy. Because the galaxy is seen face on, the spiral arms are dimmer, but the core remains bright. The galaxy is known as the Fireworks Galaxy as it frequently hosts supernovae (1917, 1939, 1948, 1968, 1969, 1980, 2002, 2004, 2008, and 2017).


RA: $294.21^{\circ} \mid 19 \mathrm{~h} 36.82^{\prime}-$ DEC: $50.2^{\circ} \mid 50^{\circ} 12^{\prime}$
R Cyg (HD 185456) is a magnitude 6.1 variable star. Magnitude ranges from 14.4 to 6.1 ( $\Delta$ mag. 8.3) with a period of 426 d .
$1.8^{\circ}$ SE from mag.3.94 iot Cyg.


This star shows a strong phenomenon of period doubling, where the maxima alternate between roughly magnitude 7 and 8 , so the period of the star can be considered twice the stated period.


Also visible:
(1) C 15 ( $9.8_{m}$ planetary nebula)
(2) NGC 6811 ( $6.8_{m}$ open cluster)
(3) RT Cyg ( $6.0_{m}$ variable star)
(4) CH Cyg ( $5.6_{m}$ variable star)


RA: $333.83^{\circ} \mid 22$ h $15.29^{\prime}-$ DEC: $49.88^{\circ} \mid 49^{\circ} 53^{\prime}$

C16 (NGC 7243) is a magnitude 6.4 open cluster.
Angular size is $21^{\prime}$.
Roughly halfway between Caph and Deneb, the cluster is pointed to by the line from Navi through Caph.

A dispersed cluster composed of many relatively faint blue and white stars. It is approximately 100 million years old.

Also visible:
(1) C 19 (10.0 $\mathrm{o}_{\mathrm{m}}$ bright nebula)
(2) NGC 7209 ( $6.7_{m}$ open cluster)


RA: $323.05^{\circ} \mid 21 \mathrm{~h} 32.2^{\prime}-$ DEC: $48.43^{\circ} \mid 48^{\circ} 26^{\prime}$
M39 (NGC 7092) is a magnitude 4.6 open cluster.
Angular size is $32^{\prime}$.
$6.5^{\circ}$ NE from mag. 3.92 ksi Cyg.

This bright cluster is a naked-eye target in dark conditions. It is a small cluster with an estimated mass of only around 230 Suns.

Also visible:
(1) C 19 (10.0 $\mathrm{o}_{\mathrm{m}}$ bright nebula)
(2) IC 1369 ( $6.8_{m}$ open cluster)
(3) NGC 7048 (11.5m planetary nebula)
(4) NGC 7026 (12.0 m planetary nebula)
(5) W Cyg ( $6.8_{m}$ variable star)


RA: $328.38^{\circ} \mid 21 \mathrm{~h} 53.5^{\prime}-$ DEC: $47.27^{\circ} \mid 47^{\circ} 16^{\prime}$
C19 (Cocoon Nebula, IC 5146) is a magnitude 10.0 bright nebula. Angular size is 12 '.

Pointed at by lines from Navi through Caph in Cassiopeia, and from Vega through Deneb.

A relatively compact nebula, connected with a spectacular region of dark nebulae $1.5^{\circ}$ to the northwest.

Also visible:
(1) C 16 ( $6.4_{m}$ open cluster)
(2) M39 (4.6 $\mathrm{K}_{\mathrm{m}}$ open cluster)
(3) NGC 7209 ( $6.7_{m}$ open cluster)
(4) W Cyg ( $6.8_{m}$ variable star)
(5) SS Cyg ( $7.7_{m}$ variable star)


RA: $303.0^{\circ} \mid 20 \mathrm{~h} 12.0^{\prime}-$ DEC: $38.35^{\circ} \mid 38^{\circ} 21^{\prime}$
C27 (Crescent Nebula, NGC 6888) is a magnitude 7.5 bright nebula. Angular size is $20 \times 10^{\prime}$.
2.7 ${ }^{\circ}$ SW from mag.2.32 Sadr.


The Crescent Nebula is formed by the ejected matter of WR 136, a massive and unstable Wolf-Rayet star. A quarter of million years ago, the star expelled a large quantity of gas when it was a red supergiant, forming a shell. Having now evolved into a WolfRayet, matter is ejected at a higher velocity from the star, which is colliding with the earlier shell and leading to the intricate structures revealed in photographs.

Also visible:
(1) M29 (7.1m open cluster)
(2) NGC 6871 (5.2m open cluster)
(3) RS Cyg ( $6.6_{m}$ carbon star)


## NGC 6871

RA: $301.48^{\circ} \mid 20 \mathrm{~h} 5.9^{\prime}-$ DEC: $35.78^{\circ} \mid 35^{\circ} 47^{\prime}$

NGC 6871 is a magnitude 5.2 open cluster. Angular size is $20^{\prime}$.
$5.5^{\circ} \mathrm{SW}$ from mag.2.32 Sadr.

This very bright cluster is estimated at less than 10 million years age and the blue and white stars that are visible are all hot type O and B stars. The cluster is quite distant at 5,133 light-years.

Also visible:
(1) C27 (7.5m bright nebula)
(2) M29 (7.1 $1_{m}$ open cluster)
(3) $\chi$ Cyg ( $3.3_{m}$ variable star)
(4) RS Cyg ( $6.6_{m}$ carbon star)


RA: $297.64^{\circ} \mid 19 \mathrm{~h} 50.56^{\prime}-$ DEC: $32.91^{\circ} \mid 32^{\circ} 55^{\prime}$
$\chi$ Cyg (HD 187796) is a magnitude 3.3 variable star. Magnitude ranges from 14.2 to 3.3 ( $\Delta$ mag. 10.9) with a period of 408 d .
6.5 ${ }^{\circ} \mathrm{NE}$ from mag.3.24 183913.


This Mira variable shows one of the largest variations in visual brightness of stars of its type, varying by a factor of 10,000 . Distance is calculated at 553 light years. The star's radius appears to range between 348 and 480 solar radii.


Also visible:
(1) NGC 6871 (5.2m open cluster)
(2) NGC 6842 (13.0 $0_{m}$ planetary nebula)
(3) SU Cyg ( $6.44_{m}$ variable star)


RA: $311.43^{\circ} \mid 20 \mathrm{~h} 45.7^{\prime}-$ DEC: $30.72^{\circ} \mid 30^{\circ} 43^{\prime}$

C34 (West Veil Nebula, NGC 6960) is a magnitude 8.0 supernova remnant. Angular size is $70 \times 6$ '.
$3.2^{\circ} \mathrm{S}$ from mag.2.64 Gienakh.

The Western Veil nebula resembles a witch's broom. Like the Eastern Veil (Caldwell 33), it is best viewed with a nebula filter at very low magnification. The Veil complex was formed by a supernova explosion between 10 and 20 thousand years ago.

Also visible:
(1) C33 (8.0 $0_{m}$ supernova remnant)
(2) NGC 6940 ( $6.3_{m}$ open cluster)
(3) $\mathrm{BD} \operatorname{Vul}\left(9.3_{m}\right.$ carbon star)

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## August: $\mathbf{1 5}^{\circ}$ North (1)



TW Peg: page 285
C37: page 285
M71: page 286
WZ Sge: page 286
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## August: $\mathbf{1 5}^{\circ}$ North (2)



NGC 6940: page 287
M27: page 288
EU Del: page 288
C42: page 289
C47: page 289

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RA: $331.0^{\circ} \mid 22 \mathrm{~h} 4.0^{\prime}-$ DEC: $28.3^{\circ} \mid 28^{\circ} 18^{\prime}$
TW Peg (HD 209598) is a magnitude 7.0 carbon star.
Magnitude ranges from 9.2 to 7.0 ( $\Delta$ mag. 2.2) with a period of 956d.
$3.0^{\circ}$ NNW from mag. 3.96 iot Peg.

A pulsating variable star with a small range in brightness but with a glacially long period of 929 days, TW Peg is located only 796 light-years from Earth.

Also visible:
(1) NGC 7217 (11.02 $2_{m}$ barred spiral galaxy)


RA: $303.0^{\circ} \mid 20 \mathrm{~h} 12.0^{\prime}-$ DEC: $26.48^{\circ} \mid 26^{\circ} 29^{\prime}$

C37 (20 Vulpeculae Cluster, NGC 6885) is a magnitude 5.7 open cluster. Angular size is $7^{\prime}$.

Roughly halfway between Deneb and Altair, slightly closer to Deneb.

Open cluster NGC 6940 (mag. 6.3) is located just over a finder circle to the north east. A finder circle to the south west lies the famous Dumbbell Nebula (Messier 27, mag. 7.4).

Also visible:
(1) M27 (7.4m planetary nebula)
(2) NGC 6842 (13.0 planetary nebula)


RA: $298.45^{\circ} \mid 19 \mathrm{~h} 53.79^{\prime}-$ DEC: $18.78^{\circ} \mid 18^{\circ} 47^{\prime}$
M71 (NGC 6838) is a magnitude 8.2 globular cluster.
Angular size is $7^{\prime}$.
$1.3^{\circ}$ SWW from mag. 3.71 gam Sge.


For a long time this globular cluster was misclassified as an open cluster, but now proven to be a globular cluster without significant central compression. M71 is relatively young ( $9-10$ billion years) and 13,000 light-years distant.

Also visible:
(1) M27 (7.4m planetary nebula)
(2) NGC 6886 (12.0 m planetary nebula)
(3) WZ Sge ( $7.0_{m}$ variable star)
(4) BF Sge ( $8.5_{m}$ carbon star)
(5) X Sge ( $8.7_{m}$ carbon star)


RA: $301.9^{\circ} \mid 20 \mathrm{~h} 7.6^{\prime}-$ DEC: $17.7^{\circ} \mid 17^{\circ} 42^{\prime}$
WZ Sge (HV 03518) is a magnitude 7.0 variable star. Magnitude ranges from 15.53 to 7.0 ( $\Delta$ mag. 8.5) with a period of 33 y .
2.7 ${ }^{\circ}$ SE from mag. 3.71 gam Sge.

A cataclysmic dwarf nova, with eruptions in 1913, 1946, 1978 and 2001, this binary system consists of a white dwarf and a low mass red or brown dwarf. The red dwarf edges across the Roche limit of the white dwarf in its rapid 1.361 hour orbit, and material is torn from it into an accretion disk. An outburst occurs when this disk collapses onto the white dwarf, releasing gravitational potential energy. 147 lightyears distant.

Also visible:
(1) M71 (8.2m globular cluster)
(2) IC 4997 (11.5m planetary nebula)
(3) NGC 6905 (11.5m planetary nebula)
(4) BF Sge ( $8.5_{m}$ carbon star)
(5) X Sge (8.7m carbon star)


RA: $322.5^{\circ} \mid 21 \mathrm{~h} 30.0^{\prime}-$ DEC: $12.17^{\circ} \mid 12^{\circ} 10^{\prime}$
M15 (Great Pegasus Globular, NGC 7078) is a magnitude 6.2 globular cluster. Angular size is $12^{\prime}$.

Midway between Altair and Markab (the corner of the square of Pegasus closest to Altair).

This globular is old even by the standards of these typically ancient objects, with an estimated age of 12.5 billion years. It is slightly too faint to be seen by the naked eye even from a dark site. It contains many interesting objects, including 8 pulsars and a planetary nebula (Pease 1 which, with magnitude 15.4 and angular size of $3^{\prime \prime}$, is best seen with a camera from a mountain-top!).

Also visible:
(1) X Peg ( $8.8_{m}$ variable star)


## NGC 6940

RA: $308.65^{\circ} \mid 20 \mathrm{~h} 34.59^{\prime}-$ DEC: $28.3^{\circ} \mid 28^{\circ} 18^{\prime}$

NGC 6940 is a magnitude 6.3 open cluster. Angular size is $31^{\prime}$.
6.1 ${ }^{\circ}$ SSW from mag.2.64 Gienakh.

Relatively ancient for an open cluster, NGC 6940 is estimated to be 720 million years old. The scattered cluster is dominated by 20 red giants. The distance to NGC 6940 is roughly 2,510 light-years.

Also visible:
(1) C33 ( $8.0_{m}$ supernova remnant)
(2) C34 (8.0 $0_{m}$ supernova remnant)
(3) $\mathrm{BD} \operatorname{Vul}\left(9.3_{m}\right.$ carbon star)


RA: $299.9^{\circ} \mid 19 \mathrm{~h} 59.59^{\prime}-$ DEC: $22.72^{\circ} \mid 22^{\circ} 43^{\prime}$
M27 (Dumbbell Nebula, NGC 6853) is a magnitude 7.4 planetary nebula. Angular size is $8 \times 6$ '.

One third of the distance from Altair to Deneb.

The Dumbbell Nebula is one of the most famous of all the Messiers and rightly so. Bright enough to be appreciated in urban skies, especially with a nebula filter, it is also generously proportioned with an intricate shape. The nebula is roughly 10,000 years old and at its center lies a white dwarf with the mass of half the Sun.

Also visible:
(1) C37 ( $5.7_{m}$ open cluster)
(2) M71 (8.2m globular cluster)
(3) NGC 6886 ( $12.0_{m}$ planetary nebula)
(4) BF Sge $\left(8.5_{m}\right.$ carbon star)
(5) X Sge (8.7m carbon star)


RA: $309.48^{\circ} \mid 20 \mathrm{~h} 37.91^{\prime}-$ DEC: $18.27^{\circ} \mid 18^{\circ} 16^{\prime}$
EU Del (HD 196610) is a magnitude 5.79 variable star. Magnitude ranges from 6.9 to 5.79 ( $\Delta$ mag. 1.1) with a period of 59.7 d .
$2.3^{\circ} \mathrm{N}$ from mag. 3.86 Svalocin.


This pulsating red giant is very close at only 359 light-years.

Also visible:
(1) IC 4997 ( $11.5_{m}$ planetary nebula)
(2) NGC 6905 (11.5m planetary nebula)
(3) U Del (7.6m variable star)


RA: $315.38^{\circ} \mid 21 \mathrm{~h} 1.5^{\prime}-$ DEC: $16.18^{\circ} \mid 16^{\circ} 11^{\prime}$
C42 (NGC 7006) is a magnitude 10.6 globular cluster. Angular size is $2.8^{\prime}$.
$5.2^{\circ}$ E from mag.3.86 Svalocin.

This globular cluster is somewhat attenuated by galactic dust. As this scatters away shorter wavelengths, the cluster is significantly brighter in infrared wavelengths. The cluster has a mass equivalent to 3.03 million Suns, or three quarters of the mass of the supermassive black hole at the center of the Milky Way.

Also visible:
(1) Segue 3 ( $14.9_{m}$ globular cluster)
(2) $\mathrm{U} \operatorname{Del}\left(7 . \sigma_{m}\right.$ variable star)
(3) $\mathrm{X} \mathrm{Peg} \mathrm{( } 8.8_{m}$ variable star)


RA: $308.55^{\circ} \mid 20 \mathrm{~h} 34.2^{\prime}-$ DEC: $7.4^{\circ} \mid 7^{\circ} 24^{\prime}$

C47 (NGC 6934) is a magnitude 8.9 globular cluster.
Angular size is $5^{\prime} 9^{\prime}$.
Two and a half finder circles east of Altair.

Roughly 50,000 light-years from Earth, the stars of this globular cluster can only be resolved with larger telescopes.

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## August: $\mathbf{- 1 5}^{\circ}$ South (1)



M2: page 293
M72: page 293
C57: page 294
RT Cap: page 294

## August: $-15^{\circ}$ South (2)




## M2

RA: $323.38^{\circ} \mid 21 \mathrm{~h} 33.5^{\prime}-$ DEC: $-0.82^{\circ} \mid 0^{\circ} 48^{\prime}$
M2 (NGC 7089) is a magnitude 6.5 globular cluster.
Angular size is $12.9^{\prime}$.
4.7 ${ }^{\circ} \mathrm{N}$ from mag.3.07 Sadalsud.
":
One of the largest globular clusters, M2 is quite distant (55,000 light-years). As such its stars are dimmed by the distance and the brightest only shine at magnitude 13.1.


M72 (NGC 6981) is a magnitude 9.3 globular cluster. Angular size is $6^{\prime}$.
3.3 ${ }^{\circ}$ SSE from mag.3.83 Albali.

This globular cluster is quite distant at over 54,000 light-years. A telescope of at least 150 mm aperture is needed to resolve some of its stars.

Also visible:
(1) C55 (8.3 ${ }_{m}$ planetary nebula)
(2) M73 (9.0 $0_{m}$ asterism)


RA: $296.23^{\circ} \mid 19 \mathrm{~h} 44.9^{\prime}-$ DEC: $-14.8^{\circ} \mid-14^{\circ} 47^{\prime}$
C57 (Barnard's Galaxy, NGC 6822) is a magnitude 9.3 irregular galaxy. Angular size is $10 \mathrm{x} 9^{\prime}$.

Due south of Altair, at the same declination as Dabikh ( 2 finder circles to the east).

This member of the Local Group of galaxies is a barred irregular galaxy discovered by E E Barnard in 1884 using a 6 inch refractor. This was the first galaxy beyond the Magellanic Clouds to have its distance measured, and was the first strong indication that the universe was substantially larger than the Milky Way. This object is difficult to spot in its busy star field, even with larger telescopes.

Also visible:
(1) NGC 6818 ( $9.3_{m}$ planetary nebula)
(2) AQ Sgr ( $6.6_{m}$ carbon star)


RA: $304.25^{\circ}\left|20 \mathrm{~h} 17.0^{\prime}-\mathrm{DEC}:-21.3^{\circ}\right|-21^{\circ} 17^{\prime}$

RT Cap (HD 192737) is a magnitude 6.5 carbon star. Magnitude ranges from 8.1 to 6.5 ( $\Delta$ mag. 1.6) with a period of 395d.
$6.5^{\circ} \mathrm{S}$ from mag.3.25 193496.

A carbon star with B-V color index 2.46 and spectral type C6,4, RT Cap also has a faint magnitude 13 companion separated by $25.7^{\prime \prime}$ at position angle $230^{\circ}$.

Also visible:
(1) M75 (8.5m globular cluster)


RA: $301.52^{\circ}\left|20 \mathrm{~h} 6.09^{\prime}-\mathrm{DEC}:-21.92^{\circ}\right|-21^{\circ} 54^{\prime}$
M75 (NGC 6864) is a magnitude 8.5 globular cluster.
Angular size is $6^{\prime}$.



This giant globular cluster appears to be gently rotating. It is over 67,000 light years from Earth, and so telescopes of at least 250 mm aperture are required to resolve the stars in this cluster. As the cluster's concentrated center is quite bright, it is relatively easy to spot with binoculars. M75 is part of the Gaia Sausage, a grouping of objects that may be the remnants of a galaxy absorbed by the Milky Way.

Also visible:
(1) RT Cap ( $6.5_{m}$ carbon star)


RA: $316.05^{\circ}\left|21 \mathrm{~h} 4.2^{\prime}-\mathrm{DEC}:-11.37^{\circ}\right|-11^{\circ} 21^{\prime}$

C55 (Saturn Nebula, NGC 7009) is a magnitude 8.3 planetary nebula. Angular size is $0.7^{\prime}$.

Midway between Fomalhaut and Altair - slightly closer to Altair.

Somewhat resembling Saturn's rings, this superb bright planetary nebula was first observed by William Herschel in 1782. It was only in the 1840s that it acquired the Saturn Nebula moniker.

Also visible:
(1) M72 (9.3m globular cluster)
(2) M73 (9.0 $0_{m}$ asterism)


RA: $314.75^{\circ}\left|20 \mathrm{~h} 59.0^{\prime}-\mathrm{DEC}:-12.63^{\circ}\right|-12^{\circ} 37^{\prime}$
M73 (NGC 6994) is a magnitude 9.0 asterism.
Angular size is $3^{\prime}$.
4. $1^{\circ} \mathrm{SE}$ from mag. 3.83 Albali.

Messier 73 is an asterism, a chance alignment of otherwise unrelated stars. Although classified as a cluster by Charles Messier, this classification was first questioned by John Herschel.


Also visible:
(1) C55 ( $8.3_{m}$ planetary nebula)
(2) M72 ( $9.3_{m}$ globular cluster)


RA: $293.58^{\circ} \mid$ 19h $34.29^{\prime}-$ DEC: $-16.4^{\circ} \mid-16^{\circ} 23^{\prime}$
$\mathrm{AQ} \operatorname{Sgr}(\mathrm{HD} 184283)$ is a magnitude 6.6 carbon star.
Magnitude ranges from 7.7 to 6.6 ( $\Delta$ mag. 1.1) with a period of 200 d .
$3.3^{\circ}$ NEE from mag. 3.95 rho01 Sgr.

A carbon star with B-V color index 3.0 and spectral type C-N5, AQ Sgr forms a wide optical double and nice color contrast with magnitude 7.55 SAO 162744 , a white main sequence star about 4 ' to the north.
Also visible:
(1) C57 (9.3 im irregular galaxy)
(2) NGC 6818 ( $9.3_{m}$ planetary nebula)
(3) R Sgr ( $6.7_{m}$ variable star)
(4) V1942 Sgr ( $6.7_{m}$ carbon star)


C63 (Helix Nebula, NGC 7293) is a magnitude 6.5
planetary nebula. Angular size is $13^{\prime}$.
7.7${ }^{\circ} \mathrm{SW}$ from mag. 3.51 Scheat.

This fabulous planetary nebula spans over three lightyears and is only 650 light-years from Earth.


RA: $325.1^{\circ} \mid 21 \mathrm{~h} 40.4^{\prime}-$ DEC: $-23.18^{\circ} \mid-23^{\circ} 10^{\prime}$
M30 (NGC 7099) is a magnitude 7.2 globular cluster. Angular size is 11 '.
$3.2^{\circ}$ SEE from mag. 3.86 zet Cap.

This globular cluster was described by William Herschel as a "remarkable globular, bright, large, slightly oval". M30 has a retrograde orbit, indicating it was in all likelihood captured from a smaller galaxy devoured by the Milky Way in the distant past.

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## August: -45 ${ }^{\circ}$ South




RA: $298.99^{\circ} \mid 19 \mathrm{~h} 55.94^{\prime}-$ DEC: $-29.19^{\circ} \mid-29^{\circ} 10^{\prime}$
RR $\operatorname{Sgr}$ (HD 188378) is a magnitude 5.4 variable star.
Magnitude ranges from 14.0 to 5.4 ( $\Delta$ mag. 8.6) with a period of 336 d .
$10.8^{\circ}$ E from mag.3.42 tau Sgr.

A moderately red pulsating variable star, RR Sgr is a respectable 1,260 light-years from Earth.


Also visible:
(1) M55 ( $6.3_{m}$ globular cluster)


RA: $295.0^{\circ} \mid 19 \mathrm{~h} 40.0^{\prime}-$ DEC: $-30.97^{\circ} \mid-30^{\circ} 57^{\prime}$
M55 (NGC 6809) is a magnitude 6.3 globular cluster. Angular size is $19^{\prime}$.
$7.9^{\circ}$ SEE from mag. 3.42 tau Sgr.

This globular cluster is relatively closer to us at 17,600 light-years, so its individual stars burn a little brighter in our telescopes. The cluster has been aged at 12.3 billion years.

Also visible:
(1) $\mathrm{RR} \operatorname{Sgr}$ (5.4m variable star)


RA: $304.43^{\circ}\left|20 \mathrm{~h} 17.72^{\prime}-\mathrm{DEC}:-39.11^{\circ}\right|-39^{\circ} 6^{\prime}$
RT Sgr (HD 192702) is a magnitude 6.0 variable star.
Magnitude ranges from 14.1 to 6.0 ( $\Delta$ mag. 8.1) with a period of 306 d .
8.9 ${ }^{\circ}$ NNW from mag. 3.21 alf Ind.

This pulsating variable star is moderately red and is about 1,870 light-years from Earth.

Also visible:
(1) U Mic (7.0 ${ }_{m}$ variable star)
(2) RU $\operatorname{Sgr}$ ( $6.0_{m}$ variable star)


## U Mic

RA: $307.32^{\circ} \mid 20 \mathrm{~h} 29.26^{\prime}-$ DEC: $-40.42^{\circ} \mid-40^{\circ} 24^{\prime}$
U Mic (HD 194814) is a magnitude 7.0 variable star.
Magnitude ranges from 14.4 to 7.0 ( $\Delta$ mag. 7.4) with a period of 334 d .
$7.0^{\circ}$ NNW from mag.3.21 alf Ind.

This pulsating variable star is moderately red and is about 2,100 light-years from Earth.

Also visible:
(1) RT Sgr ( $6.0_{m}$ variable star)


RA: $299.68^{\circ} \mid 19 \mathrm{~h} 58.71^{\prime}-$ DEC: $-41.85^{\circ} \mid-41^{\circ} 50^{\prime}$
RU Sgr (HD 188813) is a magnitude 6.0 variable star. Magnitude ranges from 13.8 to 6.0 ( $\Delta$ mag. 7.8) with a period of 240 d .
8.7 ${ }^{\circ}$ NW from mag.3.21 alf Ind.

This orange-red pulsating variable star is about 5,200 light-years from Earth. The absolute magnitude of this star is around -3.4.

Also visible:
(1) $\mathrm{RT} \operatorname{Sgr}$ ( $6.0_{m}$ variable star)


## S Gru

RA: $336.52^{\circ} \mid 22 \mathrm{~h} 26.09^{\prime}-$ DEC: $-48.44^{\circ} \mid-48^{\circ} 25^{\prime}$

S Gru (HD 212539) is a magnitude 6.0 variable star.
Magnitude ranges from 15.0 to 6.0 ( $\Delta$ mag. 9.0) with a period of 402 d .
$3.1^{\circ}$ SWW from mag. 2.24 bet Gru.

This moderately red pulsating variable star is about 2,190 light-years from Earth.

Also visible:
(1) NGC 7213 (11.01m barred spiral galaxy)


RA: $298.81^{\circ}\left|19 \mathrm{~h} 55.23^{\prime}-\mathrm{DEC}:-59.2^{\circ}\right|-59^{\circ} 11^{\prime}$
S Pav (HD 187835) is a magnitude 6.6 variable star.
Magnitude ranges from 10.4 to 6.6 ( $\Delta$ mag. 3.8) with a period of 381 d .
4.7 ${ }^{\circ}$ SW from mag.2.12 alf Pav.

Only 620 light-years from Earth, this red giant (spectral type M8III) is small in comparison to many red giants, being only 14 times the the radius of the Sun.

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## Southern Circumpolar Sky (1)



## Southern Circumpolar Sky (2)




RA: $15.8^{\circ} \mid$ 1h $3.2^{\prime}-$ DEC: $-70.85^{\circ} \mid-70^{\circ} 50^{\prime}$
C104 (NGC 362) is a magnitude 6.6 globular cluster.
Angular size is 13 '.
$6.9^{\circ}$ NNE from mag. 2.9 bet Hyi.


This moderately bright object is two or three billion years younger than the the majority of other globular clusters.


Also visible:
(1) C106 (4.0 $0_{m}$ globular cluster)
(2) SMC (2.7m Magellanic barred spiral galaxy)


RA: $13.15^{\circ} \mid 0 \mathrm{~h} 52.6^{\prime}-$ DEC: $-72.8^{\circ} \mid-72^{\circ} 47^{\prime}$

SMC is a magnitude 2.7 Magellanic barred spiral galaxy. Angular size is $315^{\prime}$.
4.8 ${ }^{\circ}$ NNE from mag. 2.9 bet Hyi.

The Small Magellanic Cloud has the mass of 7 billion Suns and spans about 7,000 light years. It is currently a satellite galaxy of the Milky Way but may formerly have been a satellite of the Large Magellanic Cloud. It is about 200,000 light years from Earth and spans over 4 degrees in our sky.

Also visible:
(1) C104 ( $6.6_{m}$ globular cluster)
(2) C106 (4.0m globular cluster)


RA: $84.68^{\circ} \mid 5 \mathrm{~h} 38.69^{\prime}-$ DEC: $-69.1^{\circ} \mid-69^{\circ} 5^{\prime}$
C103 (Tarantula Nebula, NGC 2070) is a magnitude 4.0 bright nebula. Angular size is $40 \times 25{ }^{\prime}$.
$6.6^{\circ} \mathrm{S}$ from mag. 3.81 bet Dor.

This colossal nebula shines at fourth magnitude despite being at a distance of 170,000 light-years. This is not an ordinary star-forming region like the little Orion Nebula, this complex is the bow shock of Large Magellanic Cloud as is crashes through the outer fringes of the Milky Way. The clusters forming in the Tarantula Nebula are also massive, like R136, which has the mass of half a million Suns and is quite possibly a newly formed globular cluster.

Also visible:
(1) LMC (0.91m Magellanic barred spiral galaxy)
(2) S Dor (8.6 $\sigma_{m}$ variable star)


RA: $246.45^{\circ} \mid 16 \mathrm{~h} 25.79^{\prime}-$ DEC: $-72.2^{\circ} \mid-72^{\circ} 11^{\prime}$
C107 (NGC 6101) is a magnitude 9.3 globular cluster. Angular size is $11^{\prime}$.
$3.6^{\circ} \mathrm{SSW}$ from mag.1.88 alf TrA.

This distant globular cluster (50,000 light-years from Earth) is quite dim, requiring dark skies to appreciate it.


RA: $186.45^{\circ} \mid 12 \mathrm{~h} 25.79^{\prime}-$ DEC: $-72.67^{\circ} \mid-72^{\circ} 39^{\prime}$
C108 (NGC 4372) is a magnitude 7.8 globular cluster.
Angular size is $19^{\prime}$.
3.0 ${ }^{\circ}$ SWW from mag.3.63 del Mus.

A relatively dim globular cluster, it is 19,000 lightyears away, so it can be possible to resolve some stars in the cluster with a moderately sized telescope.


Also visible:
(1) C105 (7.3 $3_{m}$ globular cluster)
(2) IC 4191 (11.5m planetary nebula)


RA: $6.03^{\circ} \mid 0$ h $24.1^{\prime}-$ DEC: $-72.08^{\circ} \mid-72^{\circ} 4^{\prime}$
C106 (47 Tucanae, NGC 104) is a magnitude 4.0 globular cluster. Angular size is $31^{\prime}$.
$5.1^{\circ} \mathrm{N}$ from mag. 2.9 bet Hyi.

The second-brightest globular cluster in our sky (after Omega Centauri), Caldwell 106 (better known as 47 Tucanae) has a very dense, bright core. The Hubble Space Telescope scanned this cluster for exoplanets. Based on the density of exoplanets near Earth around a dozen exoplanets should have been discovered, but none were - indicating planets might be rare in globulars. So there might be no-one living in these clusters, looking up at the brightest skies imaginable.

Also visible:
(1) C104 ( $6.6_{m}$ globular cluster)
(2) SMC (2.7m Magellanic barred spiral galaxy)


RA: $80.9^{\circ} \mid 5 \mathrm{~h} 23.6^{\prime}-$ DEC: $-69.75^{\circ} \mid-69^{\circ} 44^{\prime}$

LMC is a magnitude 0.91 Magellanic barred spiral galaxy. Angular size is $645^{\prime}$.
$7.3^{\circ} \mathrm{S}$ from mag. 3.81 bet Dor.

With a mass of 10 billion Suns, the Large Magellanic Cloud is $1 \%$ of the mass of the Milky Way. It is the fourth-largest galaxy in the Local Group, after M31, the Milky Way and M33. It is classified as Magellanic Spiral (a galaxy with a core and one spiral arm). It lies 163,000 light-years from Earth, and spans 10 degrees in our sky.

Also visible:
(1) C103 (4.0 $\mathrm{O}_{\mathrm{m}}$ bright nebula)
(2) S Dor (8.6 $\boldsymbol{\sigma}_{m}$ variable star)


R Oct (HD 040857) is a magnitude 6.4 variable star. Magnitude ranges from 13.2 to 6.4 ( $\Delta$ mag. 6.8 ) with a period of 405 d .
$12.3^{\circ}$ SSE from mag. 2.9 bet Hyi.


This moderately red pulsating giant star is approximately 1,650 light-years from Earth.


RA: $152.38^{\circ} \mid 10 \mathrm{~h} 9.5^{\prime}-$ DEC: $-80.87^{\circ} \mid-80^{\circ} 51^{\prime}$
C109 (NGC 3195) is a magnitude 11.6 planetary nebula. Angular size is $1.3^{\prime}$.
$10.8^{\circ} \mathrm{S}$ from mag. 3.56 ome Car.


A small and dim planetary nebula, discovered by John Herschel in 1835.


C105 (The Southern Butterfly, NGC 4833) is a magnitude 7.3 globular cluster. Angular size is $14^{\prime}$.
0.7 ${ }^{\circ}$ NNW from mag.3.63 del Mus.

Somewhat fainter and sparser than Caldwell 104, this globular cluster to 22,000 light-years from Earth.

Also visible:
(1) C 108 ( $7.8_{m}$ globular cluster)
(2) NGC 5189 (10.0 $0_{m}$ planetary nebula)
(3) IC 4191 ( $11.5_{m}$ planetary nebula)
(4) NGC 4372 ( $7.24_{m}$ globular cluster)

## Acknowledgements

This book would not be possible without the prior work of many others, so I would like to offer my thanks here to a number of my main resources. Sadly there are other resources that I have used over the years to which I also owe thanks but which have become part of the furniture of my mind and have thus been overlooked on the following list.

The Messier Catalog was one of the first and certainly the most famous of catalogs of deep sky objects. Intended as list of extended objects that were not comets as an aid to comet hunters, I must say that even in a small telescope very few Messiers look at all like a comet.
https://en.wikipedia.org/wiki/Messier_object
Being the first, Messier managed to collect many of the very best objects but by no means all. Patrick Moore's Caldwell catalog addresses this by collecting many more notable objects. A drawback of the Caldwell catalog is that a number of the objects are very difficult.
https://en.wikipedia.org/wiki/Caldwell_catalogue
In addition to these catalogs there are still a further number of galaxies, nebulae and clusters that were worth including in this book. These were mostly cataloged by the Herschels in the titanic NGC catalog.
https://en.wikipedia.org/wiki/
New_General_Catalogue
The planetarium software Stellarium has been helpful, particularly in selecting double stars that are bright and easily locatable:
http://stellarium.org/
Wikipedia is a phenomenal resource and more trustworthy than many think - but double checking is always advisable, both for Wikipedia and any other online resource!
https://en.wikipedia.org/wiki/Main_Page
The Henry Draper catalog compiled by Annie Jump Cannon is the main resource for star positions in the
charts. Some variables and high proper motion star have been manually corrected by me.
https://ui.adsabs.harvard.edu/
abs/1993yCat.3135....0C/abstract
The charts would be labelled with incomprehensible Henry Draper numbers without the help of this catalog that cross references Flamsteed and Bayer designations with HD numbers.
https://cdsarc.unistra.fr/viz-bin/cat/IV/27A
In addition to the Bayer and Flamsteed designations, the IAU provides official and usually widely accepted names for a few stars:
https://www.iau.org/public/themes/naming_stars/
The Apache Software Foundation and the Apache FOP contributors provided the PDF and SVG rendering software required for this book. They said FOP 0.96 was dead, but 20 years later it is still going strong. Version 2.6 and counting!
https://xmlgraphics.apache.org/fop/
The World Wide Web Consortium developed the XSL FO and SVG standards used by Apache FOP. The work that went into creating these standards opened up technical publishing technology to the world.
https://www.w3.org/TR/xsl/
https://www.w3.org/TR/SVG2/
The cover image was taken by me and shows Messiers 42 and 43, as well as the Running Man nebula, which is three NGC objects for the price of one: NGC 1973, NGC 1975 and NGC 1977.

This book received constructive input from the members of the Stargazer's Lounge and Cloudy Nights forums, which brought about many improvements, from diagram colors to indexing of information.

Linda Clarke contributed many hours of patient editing, diagram adjustment, and data normalization that made this book possible.

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| 29 | C 17 | $00 \mathrm{~h} 33.2^{\prime}$ | $+48^{\circ} 30.0^{\prime}$ | 9.3 | elliptical <br> galaxy | 65,61 |


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| 48 | C34 | 20h 45.7' | $+30^{\circ} 43.0^{\prime}$ | 8.0 | supernova <br> remnant | 281, 277, 287 |
| 49 | C35 | 13h 00.1' | +27 ${ }^{\circ} 59.0^{\prime}$ | 11.4 | elliptical galaxy | 177, 170 |
| 50 | C36 | 12h 36.0' | $+27^{\circ} 58.0^{\prime}$ | 9.8 | spiral galaxy | 177, 184, 170 |
| 51 | C37 | 20h 12.0' | $+26^{\circ} 29.0^{\prime}$ | 5.7 | open cluster | 288, 285 |
| 52 | C38 | 12h 36.3' | $+25^{\circ} 59.0^{\prime}$ | 9.6 | barred spiral galaxy | 184, 184, 177 |
| 53 | C39 | 07h 29.2' | $+20^{\circ} 55.0^{\prime}$ | 9.9 | planetary <br> nebula | 117 |
| 54 | C4 | 21h 01.8' | $+68^{\circ} 12.0^{\prime}$ | 6.8 | bright nebula | 55, 51 |
| 55 | C40 | 11h 20.1' | $+18^{\circ} 02.0^{\prime}$ | 10.9 | barred spiral galaxy | 178 |
| 56 | C41 | 04h 27.0' | $+16^{\circ} 00.0^{\prime}$ | 1.0 | open cluster | 91, 114 |
| 57 | C42 | 21h 01.5' | $+16^{\circ} 11.0^{\prime}$ | 10.6 | globular cluster | 289 |
| 58 | C43 | 00h 03.3' | $+16^{\circ} 09.0^{\prime}$ | 10.5 | barred spiral galaxy | 68 |
| 59 | C44 | 23h 04.9' | $+12^{\circ} 19.0^{\prime}$ | 11.0 | barred spiral galaxy | 69, 69 |
| 60 | C45 | 13h 37.5' | $+8^{\circ} 53.0^{\prime}$ | 10.2 | spiral galaxy | 217 |
| 61 | C46 | 06h 39.2' | $+8^{\circ} 44.0^{\prime}$ | 10.0 | bright nebula | 119, 116, 119 |
| 62 | C47 | 20h 34.2' | $+7^{\circ} 24.0^{\prime}$ | 8.9 | globular cluster | 289 |
| 63 | C48 | 09h 10.3' | $+7^{\circ} 02.0^{\prime}$ | 10.3 | spiral galaxy | 143, 143 |
| 64 | C49 | 06h 32.3' | $+5^{\circ} 03.0^{\prime}$ | 9.0 | bright nebula | $119,116,119$ |
| 65 | C5 | 03h 46.8' | $+68^{\circ} 06.0^{\prime}$ | 9.2 | spiral galaxy | 54 |
| 66 | C50 | 06h 32.4' | $+4^{\circ} 52.0^{\prime}$ | 4.8 | open cluster | 116, 119, 119 |
| 67 | C51 | 01h 04.8' | $+2^{\circ} 07.0^{\prime}$ | 9.0 | irregular galaxy | 70 |
| 68 | C52 | 12h 48.6' | $-5^{\circ} 0-47.9^{\prime}$ | 9.3 | elliptical galaxy | 192 |
| 69 | C53 | 10h 05.2' | $-7^{\circ} 0-42.9{ }^{\prime}$ | 9.1 | elliptical galaxy | 150 |


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| 70 | C54 | 08h 00.2' | $-10^{\circ} 0-46.9^{\prime}$ | 7.6 | open cluster | 150, 148, 151 |
| 71 | C55 | 21h 04.2' | $-11^{\circ} 0-21.9$ | 8.3 | planetary <br> nebula | 295, 293, 296 |
| 72 | C56 | 00h 47.0' | $-11^{\circ} 0-52.9^{\prime}$ | 8.0 | planetary nebula | 72 |
| 73 | C57 | 19h 44.9' | $-14^{\circ} 0-47.9^{\prime}$ | 9.3 | irregular galaxy | 296, 294 |
| 74 | C58 | 07h 17.8' | $-15^{\circ} 0-36.9{ }^{\prime}$ | 7.2 | open cluster | 129, 125, 148, 151 |
| 75 | C59 | 10h $24.8{ }^{\prime}$ | -180 0-37.9' | 8.6 | planetary nebula | 149 |
| 76 | C6 | 17h 58.6' | $+66^{\circ} 38.0^{\prime}$ | 8.8 | planetary nebula | 234 |
| 77 | C60 | 12h 01.9' | $-18^{\circ} 0-51.9^{\prime}$ | 11.3 | spiral galaxy | 194, 193 |
| 78 | C61 | 12h 01.9' | $-18^{\circ} 0-52.9^{\prime}$ | 13.0 | spiral galaxy | 194, 194 |
| 79 | C62 | 00h 47.1' | $-20^{\circ} 0-45.9^{\prime}$ | 8.9 | spiral galaxy | 74, 74 |
| 80 | C63 | 22h 29.6' | -20 ${ }^{\circ} 0-47.9^{\prime}$ | 6.5 | planetary <br> nebula | 297 |
| 81 | C64 | 07h 18.8' | $-24^{\circ} 0-56.9^{\prime}$ | 4.1 | open cluster | 130, 126, 130 |
| 82 | C65 | 00h 47.6' | $-25^{\circ} 0-16.9^{\prime}$ | 7.1 | spiral galaxy | 74, 74 |
| 83 | C66 | 14h 39.6' | $-26^{\circ} 0-31.9^{\prime}$ | 10.2 | globular cluster | 221 |
| 84 | C67 | 02h 46.3' | $-30^{\circ} 0-16.9^{\prime}$ | 9.2 | barred spiral galaxy | 100, 98, 100 |
| 85 | C68 | 19h 01.9' | $-36^{\circ} 0-56.9^{\prime}$ | 9.7 | bright nebula | 263, 263 |
| 86 | C69 | 17h 13.7' | -37 0 0-5.9 ${ }^{\prime}$ | 12.8 | planetary <br> nebula | 268 |
| 87 | C7 | 07h 36.9' | $+65^{\circ} 36.0^{\prime}$ | 8.9 | spiral galaxy | 136 |
| 88 | C70 | 00h 54.9' | -37 ${ }^{\circ} 0-40.9^{\prime}$ | 8.1 | spiral galaxy | 77 |
| 89 | C71 | 07h 52.3' | -38 ${ }^{\circ} 0-32.9{ }^{\prime}$ | 5.8 | open cluster | 155, 155, 159 |
| 90 | C72 | 00h 14.9' | -39 ${ }^{\circ} 0-10.9^{\prime}$ | 8.2 | barred spiral galaxy | 77 |
| 91 | C73 | 05h 14.1' | $-40^{\circ} 0-2.9^{\prime}$ | 7.3 | globular <br> cluster | 133 |


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| 92 | C74 | 10h 07.7 | -40 ${ }^{\circ} 0-25.9{ }^{\prime}$ | 8.2 | planetary nebula | 160 |
| 93 | C75 | 16h 25.6 | -40 ${ }^{\circ} 0-39.9{ }^{\prime}$ | 5.8 | open cluster | 225 |
| 94 | C76 | 16h 54.0' | $-41^{\circ} 0-47.9^{\prime}$ | 2.6 | open cluster | 264, 264, 268 |
| 95 | C77 | 13h 25.5 ' | -43 ${ }^{\circ} 0-0.9{ }^{\prime}$ | 7.0 | peculiar galaxy | 199, 202 |
| 96 | C78 | 18h 08.0' | $-43^{\circ} 0-41.9^{\prime}$ | 6.6 | globular cluster | 269 |
| 97 | C79 | 10h 17.6 | $-46^{\circ} 0-24.9{ }^{\prime}$ | 6.7 | globular cluster | 156 |
| 98 | C8 | 01h 29.5' | $+63^{\circ} 18.0^{\prime}$ | 9.5 | open cluster | 85, 85, 59 |
| 99 | C80 | 13h $26.8^{\prime}$ | -47 ${ }^{\circ} 0-28.9^{\prime}$ | 3.6 | globular cluster | 202, 203, 230, 199 |
| 100 | C81 | 17h $25.5{ }^{\prime}$ | $-48^{\circ} 0-24.9{ }^{\prime}$ | 8.1 | globular cluster | 269 |
| 101 | C82 | 16h 41.3' | $-48^{\circ} 0-45.9{ }^{\prime}$ | 5.2 | open cluster | 264, 265 |
| 102 | C83 | 13h 05.4' | -49 ${ }^{\circ} 0-27.9^{\prime}$ | 9.5 | spiral galaxy | 199, 203 |
| 103 | C84 | 13h 46.4' | $-51^{\circ} 0-21.9^{\prime}$ | 7.6 | globular cluster | 199, 229, 230 |
| 104 | C86 | 17h 40.7' | $-53^{\circ} 0-39.9{ }^{\prime}$ | 5.6 | globular cluster | 270 |
| 105 | C87 | 03h 12.3' | $-55^{\circ} 0-12.9{ }^{\prime}$ | 8.4 | globular cluster | 103 |
| 106 | C88 | 15h 05.7 | $-55^{\circ} 0-35.9$ | 7.9 | open cluster | 227, 231 |
| 107 | C89 | 16h 18.9' | $-57^{\circ} 0-53.9^{\prime}$ | 5.4 | open cluster | 226, 230, 232, 227 |
| 108 | C9 | 22h 56.8' | $+62^{\circ} 37.0^{\prime}$ | 7.7 | bright nebula | 59, 63, 63 |
| 109 | C90 | 09h $21.4{ }^{\prime}$ | $-58^{\circ} 0-18.9{ }^{\prime}$ | 9.7 | planetary nebula | 161, 158, 163 |
| 110 | C91 | 11h 06.4' | $-58^{\circ} 0-39.9{ }^{\prime}$ | 3.0 | open cluster | 200, 204, 204, 200 |
| 111 | C92 | 10h 43.8' | $-59^{\circ} 0-51.9^{\prime}$ | 6.2 | bright nebula | 200, 202, 200, 157 |
| 112 | C93 | 19h 10.9' | -59 ${ }^{\circ} 0-58.9^{\prime}$ | 5.4 | globular cluster | 270, 265 |
| 113 | C94 | 12h 53.6' | $-60^{\circ} 0-19.9{ }^{\prime}$ | 4.2 | open cluster | 201, 201, 205, 203 |


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| 114 | C95 | 16h 03.7' | -60 ${ }^{\circ} 0-29.9{ }^{\prime}$ | 5.1 | open cluster | 227, 232 |
| 115 | C96 | 07h 58.3' | $-60^{\circ} 0-51.9^{\prime}$ | 3.8 | open cluster | 158 |
| 116 | C97 | 11h 36.1' | $-61^{\circ} 0-36.9^{\prime}$ | 5.3 | open cluster | 200, 204, 204 |
| 117 | C98 | 12h 42.3' | $-62^{\circ} 0-57.9^{\prime}$ | 6.9 | open cluster | 201, 203, 201 |
| 118 | C99 | 12h 53.0' | $-63^{\circ} 00.0^{\prime}$ | 20.0 | dark nebula | 201, 201, 203, 205 |
| 119 | CH Cyg | 19h $24.6{ }^{\prime}$ | $+50^{\circ} 14.5{ }^{\prime}$ | 5.6 | variable star | 234, 274, 274, 278 |
| 120 | DR Ser | 18h 47.4' | $+5^{\circ} 30.0^{\prime}$ | 8.4 | carbon star | 241, 242, 243 |
| 121 | ESO224-8 | 15h 39.1' | -50 $0-3.1{ }^{\prime}$ | 14.0 | globular cluster | 226 |
| 122 | ESO452-SC11 | 16h 39.4' | $-28^{\circ} 0-23.8^{\prime}$ | 12.0 | globular cluster | 222, 222 |
| 123 | EU Del | 20h 37.9' | $+18^{\circ} 16.1^{\prime}$ | 5.79 | variable star | 288 |
| 124 | Fornax Dwarf | 02h 40.0' | $-34^{\circ} 0-26.9^{\prime}$ | 9.3 | irregular galaxy | 100, 100 |
| 125 | g Her | 16h 28.6' | $+41^{\circ} 52.9^{\prime}$ | 4.3 | variable star | 210,209 |
| 126 | GCl 38 | 16h 11.0' | $+14^{\circ} 57.5^{\prime}$ | 14.74 | globular cluster | 216 |
| 127 | GCl 50 | 16h 59.9' | $+0^{\circ} 0-32.2^{\prime}$ | 14.0 | globular cluster | 247, 247 |
| 128 | GK Ori | 06h 17.7' | $+8^{\circ} 36.0^{\prime}$ | 9.5 | carbon star | 116, 119 |
| 129 | HK Lyr | 18h 42.8' | $+37^{\circ} 00.0^{\prime}$ | 8.5 | carbon star | 236 |
| 130 | IC 2574 | 10h 28.4' | $+68^{\circ} 25.0^{\prime}$ | 10.8 | Magellanic barred spiral galaxy | 49, 49, 50, 53 |
| 131 | IC 1297 | 19h 17.4' | $-39^{\circ} 0-36.9^{\prime}$ | 9.8 | planetary <br> nebula | 263 |
| 132 | IC 1369 | 21h 12.1' | $+47^{\circ} 44.0^{\prime}$ | 6.8 | open cluster | 275, 279 |
| 133 | IC 2003 | 03h 56.4' | $+33^{\circ} 52.0^{\prime}$ | 12.0 | planetary <br> nebula | 84 |
| 134 | IC 2395 | 08h 41.1' | $-48^{\circ} 0-11.9^{\prime}$ | 4.6 | open cluster | 160 |
| 135 | IC 2501 | 09h 38.8' | $-60^{\circ} 0-4.9{ }^{\prime}$ | 11.0 | planetary <br> nebula | 159,161 |
| 136 | IC 2553 | 10h 09.3' | $-62^{\circ} 0-36.9^{\prime}$ | 11.5 | planetary <br> nebula | 202 |


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| 137 | IC 2581 | 10h 27.4' | -57º 0-37.9' | 4.3 | open cluster | 200, 157, 162, 162 |
| 138 | IC 2621 | 11h 00.3' | -65 ${ }^{\circ} 0-14.9{ }^{\prime}$ | 11.5 | planetary nebula | 202, 204, 204 |
| 139 | IC 351 | 03h 47.5' | $+35^{\circ} 03.0^{\prime}$ | 12.0 | planetary nebula | 84 |
| 140 | IC 4191 | $13 \mathrm{~h} \mathrm{08.8}{ }^{\prime}$ | $-67^{\circ} 0-38.9^{\prime}$ | 11.5 | planetary nebula | 205, 309, 311 |
| 141 | IC 4406 | 14h $22.4{ }^{\prime}$ | $-44^{\circ} 0-8.9{ }^{\prime}$ | 10.5 | planetary nebula | 229 |
| 142 | IC 4634 | 17h 01.6' | $-21^{\circ} 0-49.9{ }^{\prime}$ | 11.0 | planetary nebula | 255 |
| 143 | IC 4651 | 17h $24.7{ }^{\prime}$ | $-49^{\circ} 0-56.9^{\prime}$ | 6.9 | open cluster | 269, 270 |
| 144 | IC 4663 | 17h 45.5' | -44 ${ }^{\circ} 0-53.9{ }^{\prime}$ | 12.5 | planetary nebula | 269, 269 |
| 145 | IC 4665 | 17h 46.3' | $+5^{\circ} 43.0^{\prime}$ | 4.2 | open cluster | 243 |
| 146 | IC 4699 | 18h 18.5' | -45 ${ }^{\circ} 0-58.9{ }^{\prime}$ | 12.5 | planetary nebula | 269 |
| 147 | IC 4756 | 18h 39.0' | $+5^{\circ} 27.0^{\prime}$ | 5.0 | open cluster | 241, 242, 243 |
| 148 | IC 4997 | 20h 20.2' | $+16^{\circ} 45.0^{\prime}$ | 11.5 | planetary nebula | 286, 288 |
| 149 | KS Mon | 06h 19.8' | $-5^{\circ} 0-17.9^{\prime}$ | 8.5 | carbon star | 127, 127 |
| 150 | 1 Car | 09h 45.2' | -62 ${ }^{\circ} 0-30.4{ }^{\prime}$ | 3.28 | variable star | $\begin{aligned} & 158,159,161,162 \text {, } \\ & 162,163 \end{aligned}$ |
| 151 | LMC | 05h 23.6' | $-69^{\circ} 0-44.9{ }^{\prime}$ | 0.91 | Magellanic barred spiral galaxy | 310, 308 |
| 152 | M1 | 05h 34.5' | $+22^{\circ} 01.0^{\prime}$ | 8.4 | supernova remnant | 114 |
| 153 | M10 | 16h 57.1' | $-4^{\circ} 0-5.9^{\prime}$ | 6.6 | globular cluster | 247, 247 |
| 154 | M100 | $12 \mathrm{~h} 23.0{ }^{\prime}$ | $+15^{\circ} 50.0^{\prime}$ | 9.3 | spiral galaxy | $\begin{aligned} & 180,181,185,188 \text {, } \\ & 188,185 \end{aligned}$ |
| 155 | M101 | 14h 03.2' | $+54^{\circ} 21.0^{\prime}$ | 7.9 | spiral galaxy | 208 |
| 156 | M102 | 15h 06.5' | $+55^{\circ} 46.0^{\prime}$ | 9.9 | lenticular galaxy | 208 |
| 157 | M103 | 01h 33.2' | $+60^{\circ} 42.0^{\prime}$ | 7.4 | open cluster | 59, 64, 81, 85, 85 |


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| 158 | M104 | 12h 40.0' | - $11^{\circ} 0-36.9{ }^{\prime}$ | 8.0 | spiral galaxy | 192 |
| 159 | M105 | 10h 47.8' | $+12^{\circ} 35.0^{\prime}$ | 9.3 | elliptical galaxy | 182, 189, 180 |
| 160 | M106 | 12h 18.9' | $+47^{\circ} 19.0^{\prime}$ | 8.4 | spiral galaxy | 173, 172, 169 |
| 161 | M107 | 16h 32.5' | -13 ${ }^{\circ} 0-2.9{ }^{\prime}$ | 7.9 | globular <br> cluster | 220, 249 |
| 162 | M108 | 11h 11.5' | $+55^{\circ} 40.0^{\prime}$ | 10.0 | spiral galaxy | 171, 168 |
| 163 | M109 | 11h 57.6' | $+53^{\circ} 23.0^{\prime}$ | 9.8 | barred spiral galaxy | 171, 172, 167 |
| 164 | M11 | 18h 51.1' | $-6^{\circ} 0-15.9{ }^{\prime}$ | 6.3 | open cluster | 248, 253, 248 |
| 165 | M110 | 00h 40.4' | $+41^{\circ} 41.0^{\prime}$ | 8.5 | elliptical galaxy | 65, 62, 62, 66 |
| 166 | M12 | 16h 47.2' | $-1^{\circ} 0-56.9{ }^{\prime}$ | 6.7 | globular cluster | 247, 247 |
| 167 | M13 | 16h 41.7' | $+36^{\circ} 28.0^{\prime}$ | 5.8 | globular cluster | 236 |
| 168 | M14 | 17h 37.6' | $-3^{\circ} 0-14.9{ }^{\prime}$ | 7.6 | globular cluster | 253, 252 |
| 169 | M15 | 21h 30.0' | $+12^{\circ} 10.0^{\prime}$ | 6.2 | globular cluster | 287 |
| 170 | M16 | 18h $18.8^{\prime}$ | $-13^{\circ} 0-46.9^{\prime}$ | 6.4 | open cluster | $\begin{aligned} & 249,250,254,255, \\ & 254 \end{aligned}$ |
| 171 | M17 | 18h $20.8^{\prime}$ | $-16^{\circ} 0-10.9{ }^{\prime}$ | 7.0 | diffuse nebula | 254, 254, 256, 249 |
| 172 | M18 | 18h 19.9' | $-17^{\circ} 0-7.9^{\prime}$ | 7.5 | open cluster | 254, 255 |
| 173 | M19 | 17h 02.6' | $-26^{\circ} 0-15.9^{\prime}$ | 6.8 | globular cluster | 261, 261, 257 |
| 174 | M2 | 21h 33.5' | $+0^{\circ} 0-48.9{ }^{\prime}$ | 6.5 | globular cluster | 293 |
| 175 | M20 | 18h 02.6' | $-23^{\circ} 0-1.9^{\prime}$ | 9.0 | diffuse nebula | 250, 256 |
| 176 | M21 | 18h 04.6' | $-22^{\circ} 0-29.9{ }^{\prime}$ | 6.5 | open cluster | $\begin{aligned} & 250,250,252,256, \\ & 257,251 \end{aligned}$ |
| 177 | M22 | 18h 36.4' | $-23^{\circ} 0-53.9{ }^{\prime}$ | 5.1 | globular cluster | 252, 256, 251 |
| 178 | M23 | 17h $56.8^{\prime}$ | -19 ${ }^{\circ} 0-0.9{ }^{\prime}$ | 6.9 | open cluster | 251, 256, 250 |
| 179 | M24 | 18h 16.9' | -18 ${ }^{\circ} 0-29.9{ }^{\prime}$ | 4.6 | star cloud | $\begin{aligned} & 249,251,254,254 \\ & 255,256,256,250 \end{aligned}$ |


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| 180 | M25 | 18h 31.6' | -19 ${ }^{\circ} 0-14.9{ }^{\prime}$ | 6.5 | open cluster | $\begin{aligned} & 249,250,251,254, \\ & 255,256 \end{aligned}$ |
| 181 | M26 | 18h 45.2' | -9 ${ }^{\circ} 0-23.9{ }^{\prime}$ | 8.0 | open cluster | 248, 253, 248 |
| 182 | M27 | 19h 59.6' | $+22^{\circ} 43.0^{\prime}$ | 7.4 | planetary nebula | 285, 286, 288 |
| 183 | M28 | 18h 24.5 | $-24^{\circ} 0-51.9^{\prime}$ | 6.8 | globular cluster | 251, 251, 257, 252 |
| 184 | M29 | 20h $23.9{ }^{\prime}$ | $+38^{\circ} 32.0^{\prime}$ | 7.1 | open cluster | 276, 276, 280, 280 |
| 185 | M3 | 13h 42.2' | $+28^{\circ} 23.0^{\prime}$ | 6.2 | globular cluster | 214 |
| 186 | M30 | 21h 40.4' | $-23^{\circ} 0-10.9{ }^{\prime}$ | 7.2 | globular cluster | 297 |
| 187 | M31 | 00h 41.8' | $+41^{\circ} 16.0^{\prime}$ | 3.4 | spiral galaxy | 62, 65, 66, 62 |
| 188 | M32 | 00h 42.8' | $+40^{\circ} 52.0^{\prime}$ | 8.1 | elliptical galaxy | 66,62, 62, 65 |
| 189 | M33 | 01h 33.9' | $+30^{\circ} 39.0^{\prime}$ | 5.7 | spiral galaxy | 88 |
| 190 | M34 | 02h 42.0' | $+42^{\circ} 47.0^{\prime}$ | 5.5 | open cluster | 83, 87 |
| 191 | M35 | 06h 08.9' | $+24^{\circ} 20.0^{\prime}$ | 5.3 | open cluster | 113, 118 |
| 192 | M36 | 05h 36.1' | $+34^{\circ} 08.0^{\prime}$ | 6.3 | open cluster | $\begin{aligned} & 109,108,108,109, \\ & 110 \end{aligned}$ |
| 193 | M37 | 05h 52.4' | $+32^{\circ} 33.0^{\prime}$ | 6.2 | open cluster | 110, 109 |
| 194 | M38 | 05h 28.7 | $+35^{\circ} 50.0^{\prime}$ | 7.4 | open cluster | 108, 108, 109, 109 |
| 195 | M39 | 21h 32.2' | $+48^{\circ} 26.0^{\prime}$ | 4.6 | open cluster | 279, 279 |
| 196 | M4 | 16h 23.6' | -26 ${ }^{\circ} 0-31.9^{\prime}$ | 5.6 | globular <br> cluster | 221, 222, 222 |
| 197 | M40 | 12h $22.4{ }^{\prime}$ | $+58^{\circ} 05.0^{\prime}$ | 8.4 | double star | 171, 167 |
| 198 | M41 | 06h 47.0' | $-20^{\circ} 0-43.9^{\prime}$ | 4.6 | open cluster | 129 |
| 199 | M42 | 05h 35.4' | -5 ${ }^{\circ} 0-26.9{ }^{\prime}$ | 4.0 | diffuse nebula | 128, 123, 123, 124 |
| 200 | M43 | 05h 35.6' | $-5^{\circ} 0-15.9{ }^{\prime}$ | 9.0 | diffuse nebula | 124, 123, 123, 128 |
| 201 | M44 | 08h 40.1' | $+19^{\circ} 59.0^{\prime}$ | 3.7 | open cluster | 140, 140 |
| 202 | M45 | 03h 47.0' | $+24^{\circ} 07.0^{\prime}$ | 1.6 | open cluster | 90 |
| 203 | M46 | 07h 41.8' | $-14^{\circ} 0-48.9^{\prime}$ | 6.0 | open cluster | 151, 148, 150,151 |
| 204 | M47 | 07h 36.6' | -14 ${ }^{\circ} 0-29.9{ }^{\prime}$ | 5.2 | open cluster | $\begin{aligned} & 148,125,129,147, \\ & 151,151 \end{aligned}$ |
| 205 | M48 | 08h 13.8' | -5 ${ }^{\circ} 0-47.9^{\prime}$ | 5.5 | open cluster | 147 |
| 206 | M49 | 12h $29.8{ }^{\prime}$ | $+8^{\circ} 00.0^{\prime}$ | 8.4 | elliptical galaxy | $\begin{aligned} & 181,182,183,183, \\ & 188,189,190 \end{aligned}$ |


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| 207 | M5 | 15h 18.6' | $+2^{\circ} 05.0^{\prime}$ | 5.6 | globular cluster | 218 |
| 208 | M50 | 07h 03.2' | -8 $8^{\circ} 0-19.9{ }^{\prime}$ | 6.3 | open cluster | 124 |
| 209 | M51 | 13h 30.0' | $+47^{\circ} 11.0^{\prime}$ | 8.4 | spiral galaxy | 168 |
| 210 | M52 | 23h $24.2^{\prime}$ | $+61^{\circ} 35.0^{\prime}$ | 7.3 | open cluster | 63, 63, 59 |
| 211 | M53 | 13h 12.9' | $+18^{\circ} 10.0^{\prime}$ | 7.6 | globular cluster | 184, 178 |
| 212 | M54 | 18h 55.1' | $-30^{\circ} 0-28.9{ }^{\prime}$ | 7.6 | globular cluster | 262, 263, 266 |
| 213 | M55 | 19h 40.0' | $-30^{\circ} 0-57.9^{\prime}$ | 6.3 | globular cluster | 300, 300 |
| 214 | M56 | 19h 16.6' | $+30^{\circ} 11.0^{\prime}$ | 8.3 | globular cluster | 237 |
| 215 | M57 | 18h 53.6' | $+33^{\circ} 02.0^{\prime}$ | 8.8 | planetary nebula | 236 |
| 216 | M58 | 12h 37.7' | $+11^{\circ} 49.0^{\prime}$ | 9.7 | barred spiral galaxy | 182 |
| 217 | M59 | 12h 42.0' | $+11^{\circ} 39.0^{\prime}$ | 9.6 | elliptical galaxy | 190, 189 |
| 218 | M6 | 17h 40.1' | $-32^{\circ} 0-12.9^{\prime}$ | 4.2 | open cluster | 262, 267, 266 |
| 219 | M60 | 12h 43.7' | $+11^{\circ} 33.0^{\prime}$ | 8.8 | elliptical galaxy | $\begin{aligned} & 179,180,181,181, \\ & 182,186,188,188 \\ & 189,190,183 \end{aligned}$ |
| 220 | M61 | 12h $21.9^{\prime}$ | $+4^{\circ} 28.0^{\prime}$ | 9.7 | spiral galaxy | 190, 190, 183 |
| 221 | M62 | $17 \mathrm{~h} 01.2^{\prime}$ | -30 ${ }^{\circ} 0-6.9{ }^{\prime}$ | 6.5 | globular cluster | 257, 261, 261 |
| 222 | M63 | 13h 15.8' | $+42^{\circ} 02.0^{\prime}$ | 8.6 | spiral galaxy | 173, 169 |
| 223 | M64 | 12h $56.7^{\prime}$ | $+21^{\circ} 41.0^{\prime}$ | 8.5 | spiral galaxy | 177, 178, 184 |
| 224 | M65 | 11h 18.9' | $+13^{\circ} 05.0^{\prime}$ | 9.3 | spiral galaxy | 187, 187 |
| 225 | M66 | 11h $20.2^{\prime}$ | $+12^{\circ} 59.0^{\prime}$ | 8.9 | spiral galaxy | 187, 187 |
| 226 | M67 | 08h 50.4' | $+11^{\circ} 49.0^{\prime}$ | 6.1 | open cluster | 141 |
| 227 | M68 | 12h 39.5' | $-26^{\circ} 0-44.9^{\prime}$ | 7.8 | globular cluster | 196 |
| 228 | M69 | 18h 31.4' | $-32^{\circ} 0-20.9{ }^{\prime}$ | 7.6 | globular cluster | 262, 267 |


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| 245 | M84 | 12h $25.1^{\prime}$ | $+12^{\circ} 53.0^{\prime}$ | 9.1 | lenticular galaxy | $\begin{aligned} & 179,179,180,181, \\ & 181,182,183,185, \\ & 186,186,188,188 \end{aligned}$ |
| 246 | M85 | 12h $25.5{ }^{\prime}$ | $+18^{\circ} 12.0^{\prime}$ | 9.1 | lenticular galaxy | $\begin{aligned} & 179,179,185,186, \\ & 186,185 \end{aligned}$ |
| 247 | M86 | 12h $26.2^{\prime}$ | $+12^{\circ} 57.0^{\prime}$ | 8.9 | lenticular galaxy | $\begin{aligned} & 179,179,180,181, \\ & 182,183,185,186, \\ & 186,188,188,189, \\ & 181 \end{aligned}$ |
| 248 | M87 | $12 \mathrm{~h} 30.8^{\prime}$ | $+12^{\circ} 24.0^{\prime}$ | 8.6 | elliptical galaxy | $\begin{aligned} & 179,179,180,181, \\ & 181,182,183,185, \\ & 186,186,188,189, \\ & 190,188 \end{aligned}$ |
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| 404 | NGC 6356 | 17h 23.6' | $-17^{\circ} 0-48.7^{\prime}$ | 8.25 | globular cluster | 255, 255 |
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| 406 | NGC 6380 | 17h 34.5' | $-39^{\circ} 0-4.1^{\prime}$ | 11.31 | globular cluster | 268 |
| 407 | NGC 6384 | 17h 32.4' | $+7^{\circ} 04.0^{\prime}$ | 11.14 | barred spiral galaxy | 243 |
| 408 | NGC 6388 | 17h 36.3' | $-44^{\circ} 0-44.0^{\prime}$ | 6.72 | globular cluster | 264, 269 |
| 409 | NGC 6416 | 17h 44.4' | -32 ${ }^{\circ} 0-20.9{ }^{\prime}$ | 5.7 | open cluster | 262, 266, 267, 262 |
| 410 | NGC 6426 | 17h 44.9' | $+3^{\circ} 10.2^{\prime}$ | 11.01 | globular cluster | 243 |
| 411 | NGC 6440 | 17h 48.9' | $-20^{\circ} 0-21.5^{\prime}$ | 9.2 | globular cluster | 250 |


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| 412 | NGC 6441 | 17h 50.2' | -37 $0-3.0^{\prime}$ | 7.15 | globular cluster | 262, 267 |
| 413 | NGC 6453 | $17 \mathrm{~h} 50.8^{\prime}$ | $-34^{\circ} 0-35.9^{\prime}$ | 10.08 | globular cluster | 266 |
| 414 | NGC 6496 | 17h 59.0' | $-44^{\circ} 0-15.9{ }^{\prime}$ | 8.54 | globular cluster | 269 |
| 415 | NGC 6503 | 17h 49.5' | $+70^{\circ} 09.0^{\prime}$ | 10.91 | spiral galaxy | 234 |
| 416 | NGC 6517 | 18h 01.8' | - $8^{\circ} 0-57.4{ }^{\prime}$ | 10.23 | globular cluster | 253 |
| 417 | NGC 6522 | 18h 03.6' | $-30^{\circ} 0-1.9^{\prime}$ | 8.27 | globular cluster | 262 |
| 418 | NGC 6539 | 18h $04.8{ }^{\prime}$ | $-7^{\circ} 0-35.1^{\prime}$ | 9.33 | globular cluster | 253 |
| 419 | NGC 654 | 01h 44.1' | $+61^{\circ} 53.0{ }^{\prime}$ | 6.5 | open cluster | 59, 64, 81, 85, 85 |
| 420 | NGC 6544 | 18h 07.3' | $-24^{\circ} 0-59.7{ }^{\prime}$ | 7.77 | globular cluster | 252, 257 |
| 421 | NGC 6553 | 18h 09.3' | $-25^{\circ} 0-54.4$ | 8.06 | globular cluster | 257 |
| 422 | NGC 6558 | 18h 10.3' | $-31^{\circ} 0-45.7{ }^{\prime}$ | 9.26 | globular cluster | 267 |
| 423 | NGC 6567 | 18h 13.7 | $-19^{\circ} 0-4.9^{\prime}$ | 11.5 | planetary nebula | 250 |
| 424 | NGC 6569 | 18h 13.6 | $-31^{\circ} 0-49.5{ }^{\prime}$ | 8.55 | globular cluster | 267, 267 |
| 425 | NGC 6572 | 18h 12.1' | $+6^{\circ} 51.0^{\prime}$ | 8.5 | planetary nebula | 242 |
| 426 | NGC 6605 | 18h 17.1' | $-14^{\circ} 0-57.9^{\prime}$ | 6.0 | open cluster | $\begin{aligned} & 249,250,254,255, \\ & 256,254 \end{aligned}$ |
| 427 | NGC 6624 | 18h 23.7 | $-30^{\circ} 0-21.6$ | 7.87 | globular cluster | 262, 267 |
| 428 | NGC 6633 | 18h $27.7^{\prime}$ | $+6^{\circ} 34.0^{\prime}$ | 4.6 | open cluster | 241, 243, 242 |
| 429 | NGC 6638 | 18h 30.9' | -25 ${ }^{\circ} 0-29.7{ }^{\prime}$ | 9.02 | globular cluster | 251 |
| 430 | NGC 6642 | 18h 31.9' | $-23^{\circ} 0-28.4{ }^{\prime}$ | 9.13 | globular cluster | 251 |


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| 431 | NGC 6652 | $18 \mathrm{~h} 35.8^{\prime}$ | $-32^{\circ} 0-59.3^{\prime}$ | 8.62 | globular <br> cluster | $262,266,267$ |
| 432 | NGC 6709 | $18 \mathrm{~h} 51.5^{\prime}$ | $+10^{\circ} 21.0^{\prime}$ | 6.7 | open cluster | $241,242,241$ |
| 433 | NGC 6712 | $18 \mathrm{~h} 53.1^{\prime}$ | $-8^{\circ} 0-42.3^{\prime}$ | 8.1 | globular <br> cluster | 248,253 |


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| 451 | NGC 7217 | 22h 07.9' | $+31^{\circ} 22.0^{\prime}$ | 11.02 | barred spiral galaxy | 285 |
| 452 | NGC 7354 | 22h 40.4' | $+61^{\circ} 17.0{ }^{\prime}$ | 12.5 | planetary nebula | 63 |
| 453 | NGC 7686 | 23h $30.2^{\prime}$ | $+49^{\circ} 08.0^{\prime}$ | 5.6 | open cluster | 64, 60 |
| 454 | NGC 7789 | 23h $57.0^{\prime}$ | $+56^{\circ} 44.0^{\prime}$ | 6.7 | open cluster | 59, 60 |
| 455 | NGC 7793 | 23h $57.8^{\prime}$ | $-32^{\circ} 0-34.9{ }^{\prime}$ | 9.63 | spiral galaxy | 76 |
| 456 | Palomar 10 | 19h 18.0' | $+18^{\circ} 34.3{ }^{\prime}$ | 13.22 | globular cluster | 240 |
| 457 | Palomar 13 | 23h 06.7' | $+12^{\circ} 46.3^{\prime}$ | 13.47 | globular cluster | 69, 69 |
| 458 | Palomar 5 | 15h 16.1' | $+0^{\circ} 06.7{ }^{\prime}$ | 11.75 | globular cluster | 218 |
| 459 | R And | 00h 24.0' | $+38^{\circ} 34.6^{\prime}$ | 5.8 | variable star | 62, 65, 66, 62 |
| 460 | R Aql | 19h 06.4' | $+8^{\circ} 13.8^{\prime}$ | 5.5 | variable star | 241, 242 |
| 461 | R Aqr | 23h 43.8' | $-15^{\circ} 0-17.0^{\prime}$ | 5.8 | variable star | 73, 72 |
| 462 | R Ari | 02h 16.1' | $+25^{\circ} 03.4$ | 7.4 | variable star | 90 |
| 463 | R Aur | 05h 17.3' | $+53^{\circ} 35.2^{\prime}$ | 6.7 | variable star | 106 |
| 464 | R Boo | 14h 37.2' | $+26^{\circ} 44.2^{\prime}$ | 6.2 | variable star | 215 |
| 465 | R Car | 09h 32.2' | -62 ${ }^{\circ} 0-47.2^{\prime}$ | 3.9 | variable star | $\begin{aligned} & 163,158,159,161, \\ & 162,162 \end{aligned}$ |
| 466 | R Cas | 23h $58.4{ }^{\prime}$ | $+51^{\circ} 23.3{ }^{\prime}$ | 4.7 | variable star | 64, 60 |
| 467 | R Cen | 14h 16.6' | -59 0 0-54.7' | 5.3 | variable star | $\begin{aligned} & 228,228,231,232, \\ & 231 \end{aligned}$ |
| 468 | R CMa | 07h 19.5' | -16 ${ }^{\circ} 0-23.6$ | 5.7 | variable star | 125, 129, 148, 151 |
| 469 | R Cnc | 08h 16.6' | $+11^{\circ} 43.6^{\prime}$ | 6.07 | variable star | 141 |
| 470 | R Com | 12h 04.3' | $+18^{\circ} 46.9^{\prime}$ | 7.1 | variable star | 179, 185, 186 |
| 471 | R CrB | 15h 48.6' | $+28^{\circ} 09.4{ }^{\prime}$ | 5.71 | variable star | 214, 215 |
| 472 | R Cru | 12h $23.6{ }^{\prime}$ | $-61^{\circ} 0-37.6^{\prime}$ | 6.4 | variable star | 201, 203, 205, 201 |
| 473 | R Crv | 12h 19.6' | $-19^{\circ} 0-15.3^{\prime}$ | 6.7 | variable star | 193, 194, 194 |
| 474 | R Cyg | 19h 36.8' | $+50^{\circ} 12.0^{\prime}$ | 6.1 | variable star | 278, 274, 274 |
| 475 | R For | 02h $29.3{ }^{\prime}$ | -26 ${ }^{\circ} 0-5.8^{\prime}$ | 7.5 | variable star | 98, 100 |
| 476 | R Gem | 07h 07.4' | $+22^{\circ} 42.2^{\prime}$ | 6.0 | variable star | 113 |
| 477 | R Hor | 02h 53.9' | -49 ${ }^{\circ} 0-53.3^{\prime}$ | 4.7 | variable star | 102 |
| 478 | R Hya | 13h $29.7{ }^{\prime}$ | $-23^{\circ} 0-16.8^{\prime}$ | 3.5 | variable star | 195 |


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| 479 | R Leo | 09h 47.6' | $+11^{\circ} 25.7$ | 4.4 | variable star | 142 |
| 480 | R Lep | 04h 59.6' | $-14^{\circ} 0-47.9^{\prime}$ | 5.9 | carbon star | 125,128, 128 |
| 481 | R LMi | 09h 45.6' | $+34^{\circ} 30.7{ }^{\prime}$ | 6.3 | variable star | 138 |
| 482 | R Lyr | 18h 55.3' | $+43^{\circ} 56.8^{\prime}$ | 3.88 | variable star | 235 |
| 483 | R Nor | 15h 36.0' | -49 ${ }^{\circ} 0-30.4{ }^{\prime}$ | 6.5 | variable star | 226 |
| 484 | R Oct | 05h 26.1' | -86 ${ }^{\circ} 0-23.2^{\prime}$ | 6.4 | variable star | 310 |
| 485 | R Ori | 04h 59.0' | $+8^{\circ} 07.8^{\prime}$ | 9.05 | variable star | 115 |
| 486 | R Peg | 23h 06.7' | $+10^{\circ} 32.6^{\prime}$ | 6.9 | variable star | 69,69 |
| 487 | R Pic | 04h 46.2' | -49 ${ }^{\circ} 0-14.7{ }^{\prime}$ | 6.35 | variable star | 133 |
| 488 | R Psc | 01h 30.6' | $+2^{\circ} 52.9^{\prime}$ | 7.0 | variable star | 93 |
| 489 | R Scl | 01h 27.0' | -32 ${ }^{\circ} 0-29.9{ }^{\prime}$ | 6.1 | carbon star | 76 |
| 490 | R Sct | 18h 47.5' | $-5^{\circ} 0-42.2{ }^{\prime}$ | 4.2 | variable star | 248, 248, 253 |
| 491 | R Ser | 15h 50.7' | $+15^{\circ} 08.0^{\prime}$ | 5.16 | variable star | 216 |
| 492 | R Sgr | 19h 16.7' | $-19^{\circ} 0-18.4^{\prime}$ | 6.7 | variable star | 296 |
| 493 | R Tau | 04h 28.3' | $+10^{\circ} 09.7{ }^{\prime}$ | 7.6 | variable star | 93 |
| 494 | R Tri | 02h 37.0' | $+34^{\circ} 15.9^{\prime}$ | 5.4 | variable star | 84 |
| 495 | R UMa | 10h 44.6' | $+68^{\circ} 46.5^{\prime}$ | 6.5 | variable star | 49, 49, 50, 53 |
| 496 | R Vir | 12h $38.5{ }^{\prime}$ | $+6^{\circ} 59.3$ ' | 6.1 | variable star | 183, 183, 189, 190 |
| 497 | RR Sco | 16h 56.6' | $-30^{\circ} 0-34.7{ }^{\prime}$ | 5.0 | variable star | 257, 261, 261 |
| 498 | RR Sgr | 19h 55.9' | $-29^{\circ} 0-11.3^{\prime}$ | 5.4 | variable star | 300, 300 |
| 499 | RS Cyg | 20h 13.4' | $+38^{\circ} 42.0^{\prime}$ | 6.6 | carbon star | 276,280, 280 |
| 500 | RS Oph | $17 \mathrm{~h} 50.2^{\prime}$ | -6 ${ }^{\circ} 0-42.4{ }^{\prime}$ | 4.3 | variable star | 252, 253 |
| 501 | RS Pup | 08h 13.1' | $-34^{\circ} 0-34.6^{\prime}$ | 6.52 | variable star | 155, 155 |
| 502 | RS Sco | 16h 55.6' | $-45^{\circ} 0-6.1^{\prime}$ | 6.2 | variable star | 264, 264, 265, 268 |
| 503 | RT Cap | 20h 17.0' | $-21^{\circ} 0-17.9^{\prime}$ | 6.5 | carbon star | 295, 294 |
| 504 | RT Cyg | 19h 43.6' | $+48^{\circ} 46.7^{\prime}$ | 6.0 | variable star | 234, 274, 274, 278 |
| 505 | RT Pup | 08h 05.4' | $-38^{\circ} 0-41.9^{\prime}$ | 8.5 | carbon star | 155, 155, 159 |
| 506 | RT Sco | 17h 03.5' | $-36^{\circ} 0-55.1^{\prime}$ | 7.0 | variable star | 268 |
| 507 | RT Sgr | 20h 17.7' | -39 ${ }^{\circ} 0-6.7^{\prime}$ | 6.0 | variable star | 301, 302, 301 |
| 508 | RT UMa | 09h 18.4' | $+51^{\circ} 24.0^{\prime}$ | 8.6 | carbon star | 137, 136 |
| 509 | RU Sgr | 19h $58.7{ }^{\prime}$ | $-41^{\circ} 0-50.9{ }^{\prime}$ | 6.0 | variable star | 301, 302 |
| 510 | Ruprecht 106 | 12h 38.7' | $-51^{\circ} 0-8.9{ }^{\prime}$ | 10.9 | globular cluster | 203 |


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| 511 | RX Lep | 05h 11.4' | $-11^{\circ} 0-50.8^{\prime}$ | 5.0 | variable star | 128, 125 |
| 512 | RY Dra | 12h 56.4' | $+66^{\circ} 00.0^{\prime}$ | 6.0 | carbon star | 167, 52 |
| 513 | RY Sgr | 19h 16.5' | $-33^{\circ} 0-31.2^{\prime}$ | 5.8 | variable star | 263, 266, 263 |
| 514 | S Car | 10h 09.4' | $-61^{\circ} 0-32.8^{\prime}$ | 4.5 | variable star | $\begin{aligned} & 200,202,162,157, \\ & 158,162,163 \end{aligned}$ |
| 515 | S CMi | 07h 32.7' | $+8^{\circ} 19.1{ }^{\prime}$ | 6.6 | variable star | 142 |
| 516 | SCnc | 08h 43.9' | $+19^{\circ} 02.1^{\prime}$ | 8.29 | variable star | 140, 140 |
| 517 | S CrB | 15h 21.4 | $+31^{\circ} 22.0^{\prime}$ | 5.8 | variable star | 211,211 |
| 518 | S Dor | 05h 18.2' | -69 ${ }^{\circ} 0-14.9^{\prime}$ | 8.6 | variable star | 308, 310 |
| 519 | S Gem | 07h 43.0' | $+23^{\circ} 27.0^{\prime}$ | 8.0 | variable star | 117 |
| 520 | S Gru | 22h $26.1^{\prime}$ | -48 ${ }^{\circ} 0-26.2^{\prime}$ | 6.0 | variable star | 302 |
| 521 | S Her | 16h 51.9' | $+14^{\circ} 56.5^{\prime}$ | 6.4 | variable star | 240 |
| 522 | S Hya | 08h 53.6' | $+3^{\circ} 04.1{ }^{\prime}$ | 7.2 | variable star | 143, 143 |
| 523 | S Pav | 19h 55.2' | $-59^{\circ} 0-11.6^{\prime}$ | 6.6 | variable star | 303 |
| 524 | S Scl | 00h 15.4' | $-32^{\circ} 0-2.6^{\prime}$ | 5.5 | variable star | 76 |
| 525 | S Sct | 18h 50.3' | $-7^{\circ} 0-53.9{ }^{\prime}$ | 7.3 | carbon star | 248, 248, 253 |
| 526 | S Ser | 15h $21.7^{\prime}$ | $+14^{\circ} 18.9^{\prime}$ | 7.0 | variable star | 217 |
| 527 | S Vir | 13h 33.0' | -7 ${ }^{\circ} 0-11.6^{\prime}$ | 6.3 | variable star | 220 |
| 528 | Sculptor | 01h 00.2' | $-33^{\circ} 0-42.9^{\prime}$ | 10.5 | irregular galaxy | 76,77 |
| 529 | Segue 3 | 21h $21.5^{\prime}$ | $+19^{\circ} 07.0^{\prime}$ | 14.9 | globular cluster | 289 |
| 530 | SMC | 00h 52.6' | $-72^{\circ} 0-47.9^{\prime}$ | 2.7 | Magellanic barred spiral galaxy | 307, 309, 307 |
| 531 | SS Cyg | 21h 42.7' | $+43^{\circ} 35.2^{\prime}$ | 7.7 | variable star | 279 |
| 532 | SS Sgr | 18h 30.4' | $-16^{\circ} 0-53.9^{\prime}$ | 9.0 | carbon star | 254 |
| 533 | SS Vir | 12h 25.3 ' | $+0^{\circ} 42.0^{\prime}$ | 6.0 | carbon star | 183, 190 |
| 534 | ST Cas | 00h 17.6' | $+50^{\circ} 18.0^{\prime}$ | 9.0 | carbon star | 60, 61, 65 |
| 535 | SU Cyg | 19h 44.8' | $+29^{\circ} 15.9^{\prime}$ | 6.44 | variable star | 281 |
| 536 | SU Sco | 16h 40.6' | -32 ${ }^{\circ} 0-23.9^{\prime}$ | 8.0 | carbon star | 261, 261 |
| 537 | SU Tau | 05h 49.1' | $+19^{\circ} 04.4{ }^{\prime}$ | 9.1 | variable star | 114, 118 |
| 538 | SX Sco | 17h 47.5' | $-35^{\circ} 0-41.9^{\prime}$ | 8.5 | carbon star | 262, 266, 267 |
| 539 | T Cen | 13h 41.8' | $-33^{\circ} 0-35.7{ }^{\prime}$ | 5.5 | variable star | 225, 229 |
| 540 | T Cep | 21h 09.5' | $+68^{\circ} 29.5{ }^{\prime}$ | 5.2 | variable star | 55,51 |


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| 541 | T Cet | 00h 21.8' | -20 ${ }^{\circ} 0-3.4{ }^{\prime}$ | 5.0 | variable star | 73 |
| 542 | T Cnc | 08h 56.7' | $+19^{\circ} 48.0^{\prime}$ | 7.8 | carbon star | 140, 140, 140, 140 |
| 543 | T Col | 05h 19.3' | -33 ${ }^{\circ} 0-42.4{ }^{\prime}$ | 6.6 | variable star | 132 |
| 544 | T CrB | 15h 59.5' | $+25^{\circ} 55.2^{\prime}$ | 2.0 | variable star | 214, 214, 215 |
| 545 | T Eri | 03h 55.2' | -24 ${ }^{\circ} 0-1.8^{\prime}$ | 7.2 | variable star | 97 |
| 546 | T Gem | 07h 49.3' | $+23^{\circ} 44.1^{\prime}$ | 8.0 | variable star | 117 |
| 547 | T Lyr | 18h 32.3' | $+37^{\circ} 00.0^{\prime}$ | 7.5 | carbon star | 236 |
| 548 | T Nor | 15h 44.1' | $-54^{\circ} 0-59.1^{\prime}$ | 6.2 | variable star | 230, 226 |
| 549 | T Sct | 18h 55.4' | $-8^{\circ} 0-5.9{ }^{\prime}$ | 8.9 | carbon star | 248 |
| 550 | T Tau | 04h 22.0' | $+19^{\circ} 32.1^{\prime}$ | 9.3 | variable star | 91 |
| 551 | T UMa | 12h 36.4' | $+59^{\circ} 29.2^{\prime}$ | 6.6 | variable star | 171 |
| 552 | Terzan 6 | $17 \mathrm{~h} 50.8^{\prime}$ | -31 ${ }^{\circ} 0-16.4^{\prime}$ | 13.85 | globular cluster | 266 |
| 553 | TU Gem | 06h 10.9' | $+26^{\circ} 00.0^{\prime}$ | 7.5 | carbon star | 113 |
| 554 | TW Peg | 22h 04.0' | $+28^{\circ} 18.0^{\prime}$ | 7.0 | carbon star | 285 |
| 555 | TX Psc | 23h 46.4' | $+3^{\circ} 29.2^{\prime}$ | 4.79 | variable star | 70 |
| 556 | U Ant | $10 \mathrm{~h} \mathrm{35.2}$ | -39 0 0-35.9' | 5.7 | carbon star | 199 |
| 557 | U Ara | 17h 53.6' | $-51^{\circ} 0-41.1^{\prime}$ | 7.7 | variable star | 269, 270 |
| 558 | U Ari | 03h 11.1' | $+14^{\circ} 48.0^{\prime}$ | 7.2 | variable star | 92 |
| 559 | U Cep | 01h 02.3' | $+81^{\circ} 52.5$ | 6.75 | variable star | 51, 53, 54 |
| 560 | U Cet | 02h 33.7' | $-13^{\circ} 0-8.8^{\prime}$ | 6.8 | variable star | 96 |
| 561 | UCrB | 15h 18.2' | $+31^{\circ} 38.8^{\prime}$ | 7.66 | variable star | 211, 211 |
| 562 | U Cyg | 20h 19.6' | $+47^{\circ} 53.7^{\prime}$ | 5.9 | variable star | 275 |
| 563 | U Del | 20h 45.5' | $+18^{\circ} 05.4^{\prime}$ | 7.6 | variable star | 288, 289 |
| 564 | u Her | 17h 17.3' | $+33^{\circ} 06.0^{\prime}$ | 4.69 | variable star | 237 |
| 565 | U Hor | 03h 52.8' | -45 ${ }^{\circ} 0-49.7^{\prime}$ | 7.8 | variable star | 102 |
| 566 | U Hya | 10h 37.6' | $-13^{\circ} 0-23.0^{\prime}$ | 7.0 | variable star | 193, 193 |
| 567 | U Mic | 20h 29.3' | -40 0-24.9' | 7.0 | variable star | 301, 301 |
| 568 | U Mon | 07h 30.8' | -9 $9^{\circ} 0-46.5^{\prime}$ | 6.1 | variable star | 147, 148, 151 |
| 569 | U Ori | 05h 55.8' | $+20^{\circ} 10.5{ }^{\prime}$ | 4.8 | variable star | 118, 113 |
| 570 | U Sge | 19h 18.8' | $+19^{\circ} 36.6^{\prime}$ | 6.45 | variable star | 240 |
| 571 | U Sgr | 18h 31.9' | -19 ${ }^{\circ} 0-7.4^{\prime}$ | 6.28 | variable star | $\begin{aligned} & 249,250,254,255, \\ & 256 \end{aligned}$ |
| 572 | UU Aur | 06h 36.5' | $+38^{\circ} 30.0^{\prime}$ | 5.1 | carbon star | 107, 107 |


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| 573 | UV Aql | 18h 58.6' | $+14^{\circ} 24.0^{\prime}$ | 8.6 | carbon star | 241 |
| 574 | UX Dra | 19h 21.6 | $+76^{\circ} 36.0^{\prime}$ | 6.2 | carbon star | 50 |
| 575 | V Aql | 19h 04.4' | $-5^{\circ} 0-41.9^{\prime}$ | 6.6 | carbon star | 248, 248, 253 |
| 576 | V Ari | 02h 15.0' | $+12^{\circ} 12.0^{\prime}$ | 8.0 | carbon star | 92 |
| 577 | V CrB | 15h 49.5' | $+39^{\circ} 34.3{ }^{\prime}$ | 6.9 | variable star | 210 |
| 578 | V Cyg | 20h 41.3' | $+48^{\circ} 12.0^{\prime}$ | 7.8 | carbon star | 275, 275 |
| 579 | V Hya | 10h 51.6' | $-21^{\circ} 0-17.9^{\prime}$ | 6.5 | carbon star | 195 |
| 580 | V Mon | 06h $22.7{ }^{\prime}$ | $-2^{\circ} 0-11.6$ | 6.0 | variable star | 127, 127 |
| 581 | V Oph | 16h $26.7^{\prime}$ | $-12^{\circ} 0-25.5{ }^{\prime}$ | 7.3 | variable star | 249, 220 |
| 582 | V1942 Sgr | 19h 19.2' | $-15^{\circ} 0-53.9^{\prime}$ | 6.7 | carbon star | 296 |
| 583 | VV Ori | 05h 33.5' | $-1^{\circ} 0-9.3^{\prime}$ | 5.31 | variable star | $\begin{aligned} & 123,120,123,124, \\ & 128 \end{aligned}$ |
| 584 | VX And | 00h 19.9' | $+44^{\circ} 42.0^{\prime}$ | 8.0 | carbon star | 61, 62, 65, 65, 66 |
| 585 | VX Sgr | 18h 08.1' | -22 ${ }^{\circ} 0-13.3{ }^{\prime}$ | 6.52 | variable star | $\begin{aligned} & 250,251,252,256, \\ & 257 \end{aligned}$ |
| 586 | VY CMa | 07h 23.0' | $-25^{\circ} 0-46.0^{\prime}$ | 6.5 | variable star | $\begin{aligned} & 130,126,126,130, \\ & 130 \end{aligned}$ |
| 587 | VY UMa | 10h 45.0' | $+67^{\circ} 24.0^{\prime}$ | 6.0 | carbon star | 49, 49, 50, 53 |
| 588 | W Cas | 00h 54.9' | $+58^{\circ} 33.8^{\prime}$ | 7.8 | variable star | 59, 60, 64 |
| 589 | W Cet | 00h 02.1' | $-14^{\circ} 0-40.5{ }^{\prime}$ | 7.1 | variable star | 72, 73 |
| 590 | W CMa | 07h 08.0' | $-11^{\circ} 0-53.9$ | 7.0 | carbon star | 124, 125, 129 |
| 591 | W CrB | 16h 15.4' | $+37^{\circ} 47.7^{\prime}$ | 7.8 | variable star | 209, 210 |
| 592 | W Cyg | 21h 36.0' | $+45^{\circ} 22.5^{\prime}$ | 6.8 | variable star | 279, 279 |
| 593 | W Ori | 05h 05.4' | $+1^{\circ} 12.0^{\prime}$ | 6.5 | carbon star | 120 |
| 594 | W UMa | 09h 43.8' | $+55^{\circ} 57.2^{\prime}$ | 7.75 | variable star | 136, 137 |
| 595 | WLM | 00h 01.9' | - $15^{\circ} 0-26.9^{\prime}$ | 11.03 | irregular galaxy | 72, 73 |
| 596 | WW Cas | 01h 33.5' | $+57^{\circ} 48.0^{\prime}$ | 9.1 | carbon star | 64, 81, 85, 85 |
| 597 | WZ Cas | 00h 01.3' | $+60^{\circ} 21.3^{\prime}$ | 6.3 | variable star | 59, 60 |
| 598 | WZ Sge | 20h 07.6' | $+17^{\circ} 42.2^{\prime}$ | 7.0 | variable star | 286, 286 |
| 599 | X Cnc | 08h 55.4' | $+17^{\circ} 13.9^{\prime}$ | 5.6 | variable star | 140, 140 |
| 600 | X Cyg | 20h 43.4' | $+35^{\circ} 35.3^{\prime}$ | 5.85 | variable star | 276, 276, 277 |
| 601 | X Her | 16h 02.7' | $+47^{\circ} 14.4{ }^{\prime}$ | 6.0 | variable star | 209 |
| 602 | X Oph | 18h 38.4' | $+8^{\circ} 50.0^{\prime}$ | 5.9 | variable star | 241, 242, 243, 241 |
| 603 | X Peg | 21h $21.0^{\prime}$ | $+14^{\circ} 27.0^{\prime}$ | 8.8 | variable star | 287, 289 |
| 604 | X Per | 03h 55.4' | $+31^{\circ} 02.8^{\prime}$ | 6.03 | variable star | 84 |


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| 605 | X Sge | 20h 05.1' | $+20^{\circ} 36.0^{\prime}$ | 8.7 | carbon star | 286, 286, 288 |
| 606 | X Vel | 09h 55.5' | -41 ${ }^{\circ} 0-29.9{ }^{\prime}$ | 8.4 | carbon star | 160 |
| 607 | Y Cvn | 12h 45.1' | $+45^{\circ} 24.0^{\prime}$ | 5.0 | carbon star | 172, 173, 173, 169 |
| 608 | Y Lyn | 07h 28.2' | $+46^{\circ} 00.0^{\prime}$ | 6.9 | carbon star | 106 |
| 609 | Y Per | 03h $27.7^{\prime}$ | $+44^{\circ} 12.0^{\prime}$ | 8.1 | carbon star | 83 |
| 610 | Y Tau | 05h 45.7' | $+20^{\circ} 42.0^{\prime}$ | 7.1 | carbon star | 114, 118 |
| 611 | Z Psc | 01h 16.1' | $+25^{\circ} 48.0^{\prime}$ | 7.0 | carbon star | 68 |
| 612 | Z UMa | 11h 56.5' | +57 ${ }^{\circ} 52.3^{\prime}$ | 6.2 | variable star | 171, 172, 167 |
| 613 | $\beta$ Dor | 05h 33.6' | -62 ${ }^{\circ} 0-29.3{ }^{\prime}$ | 3.46 | variable star | 134 |
| 614 | o Cet | 02h 19.3' | $-2^{\circ} 0-58.6{ }^{\prime}$ | 2.0 | variable star | 96 |
| 615 | $\chi \mathrm{Cyg}$ | 19h 50.6' | $+32^{\circ} 54.8{ }^{\prime}$ | 3.3 | variable star | 281, 280 |

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[^0]:    Observed:

[^1]:    Observed:

[^2]:    Observed:

